

Energy Levels of Calcium, Ca I through Ca XX

Jack Sugar and Charles Corliss

National Measurement Laboratory, National Bureau of Standards, Washington, D.C. 20234

The energy levels of the calcium atom in all of its stages of ionization, as derived from the analyses of atomic spectra, have been critically compiled. In cases where only line classifications are reported in the literature, level values have been derived. Electron configurations, term designations, J-values, experimental g-values, and ionization energies are included. Calculated percentages of the two leading components of the eigenvectors of the levels are given.

Key words: Atomic energy levels; atomic spectra; calcium energy levels.

Contents

	Page		Page
Introduction	865	Ca X	903
Energy Level Tables	868	Ca XI	905
Ca I	868	Ca XII	906
Ca II	888	Ca XIII	907
Ca III	890	Ca XIV	908
Ca IV	895	Ca XV	910
Ca V	897	Ca XVI	912
Ca VI	898	Ca XVII	913
Ca VII	899	Ca XVIII	914
Ca VIII	900	Ca XIX	915
Ca IX	901	Ca XX	916

Introduction

At the time of the first compilation of atomic energy levels by Bacher and Goudsmit in 1932, only the first 5 of the 20 spectra of calcium had been studied. By 1949, Moore was able to compile the first 12 spectra of calcium. At that time, oxygen was the heaviest atom for which some levels of all stages of ionization were known.

A great amount of new experimental work has been carried out since then, particularly in the higher stages of ionization. Today, experimental results are available for every stage of ionization of calcium. This is the result of the development of more energetic light sources, which was stimulated by the need to interpret new spectroscopic observations of the sun at short wavelengths from rocket- and satellite-borne spectrographs. A new impetus for the interpretation of spectra of highly ionized atoms has arisen from the investigation of hot laboratory plasmas generated to achieve nuclear fusion.

These activities have produced a substantial increase in spectroscopic information, particularly for elements of the iron period, making the earlier compilations of energy

levels inadequate. The NBS Atomic Energy Levels Data Center has undertaken to provide new compilations of energy levels, including the elements of the iron period. The material on each atom and its ions is being published as a separate paper. A collection of these compilations, with revisions, is planned as one volume for the iron period. Already completed are the compilation for iron by Reader and Sugar (1975), chromium and vanadium by Sugar and Corliss (1977, 1978), and manganese and titanium by Corliss and Sugar (1977, 1979). The present work on calcium will be followed by a compilation of the energy levels of potassium.

The present compilation comprises the energy levels of the calcium atom and all of its ions, as derived from analyses of atomic spectra. For many of the ions the original papers do not give energy level values, but only classifications of observed lines. In these cases we have derived the level values. Although generally we used only published papers as sources of data, unpublished data have been included when they constituted a substantial improvement over material in the literature.

Ionization energies found in the literature are usually given in their equivalence in cm^{-1} . The conversion factor $8065.479 \pm 0.021 \text{ cm}^{-1}/\text{eV}$, as given by Cohen and Taylor

© 1979 by the U.S. Secretary of Commerce on behalf of the United States. This copyright is assigned to the American Institute of Physics and the American Chemical Society.

(1973), was used to obtain values in eV. In a few cases where adequate data were available but the ionization energy was not derived, we carried out the calculation. For a large number of ions, no suitable series are known. In these cases we have quoted values obtained by extrapolation along isoelectronic sequences. Although uncertainties are not usually provided with these extrapolated values, they are probably accurate to a few units of the last significant figure given.

Nearly all of the data are the result of observations of various types of laboratory light sources. However, they are sometimes supplemented by data obtained from solar observations. This is particularly true where spin-forbidden lines are required to establish the absolute energy of a system of excited levels and also where parity-forbidden transitions between levels of a ground configuration are used to obtain accurate relative energies for the low levels. Whenever both solar data and equivalent laboratory data are available preference is generally given to the laboratory measurements in order to avoid the problem of blended lines of various elements in the solar spectrum.

For a convenient source of wavelengths of calcium lines below 2000 Å we refer the reader to the compilation by Kelly and Palumbo (1973).

We sometimes assign a calculated value to a level of a term in a system not connected to the ground state. The error in the calculated value is indicated by the letter *x* following the level values of that system. For Ca xix and xx, which are isoelectronic with He I and H I we give only theoretical level values. They are much more accurate than experimental x-ray wavelengths from which level values may be obtained.

For a given configuration a certain number of terms of various types are theoretically expected, and spectroscopists have given names on the basis of *J*-values, *g*-values, intensities, arrangement of the levels, and, more recently, theoretical calculations. We have included the results of calculations, under the heading "Leading percentages", that express the percentage composition of levels in terms of the basis states of a single configuration, or more than one configuration where configuration interaction has been included. Where these results contradict an author's designation, we have accepted the theoretical term and configuration labeling of a level to conform with its calculated leading percentages. In some cases these are low and the labeling has less physical meaning.

The percentage compositions have the following meaning. Suppose that for a given configuration there is a set of *n* basis states, written symbolically as $\psi_1, \psi_2, \dots, \psi_n$. Usually these basis states are taken to be the *LS*-states for a configuration, but other coupling schemes are often used. (See Martin, Zalubas, and Hagan (1978) for coupling notation.) Then the eigenvector ψ_A of an actual energy level *A* can be expressed as

$$\psi_A = \alpha_1 \psi_1 + \alpha_2 \psi_2 + \dots + \alpha_n \psi_n,$$

where $\alpha_1^2 + \alpha_2^2 + \dots + \alpha_n^2 = 1$. The squared quantities α_1^2, α_2^2 , etc., multiplied by 100, represent the percentage composi-

tion of a given level. Levels are given names corresponding to the basis state having the largest percentage.

In the columns of the present tables headed "Leading percentages" we give first the percentage of the basis state corresponding to the level's name; next the second largest percentage together with the related basis state. We have not listed any second component representing less than 4 percent.

Of course, the percentage compositions cannot be considered to be as reliable as experimental quantities inasmuch as a new calculation using a different approximation, such as the introduction of configuration interaction where none had been used before, might yield a different set of percentages. For some levels the percentages may change drastically in a new calculation. In the present tables, the percentages are taken mostly from published least-squares level fitting calculations. When only *ab initio* calculations are found in the literature, we have used them if there appears to be a reasonable correspondence with the experimental data. For higher ionization stages there have been fewer publications relating quantitatively the theoretical results to the observations by means of least-squares calculations.

In assembling the data for each spectrum, we referred to the following bibliographies:

- i. papers cited by Moore (1949)
- ii. C. E. Moore (1968)
- iii. L. Hagan and W. C. Martin (1972).
- iv. L. Hagan (1977)
- v. card file of publications since June 1975 maintained by the NBS Atomic Energy Levels Data Center.

A selection of data was made that, in our judgment, represents the most accurate and reliable available. The text for each ion is not always a complete review of the literature but is intended to credit the major contributions. A final check for new data was made on September 1, 1978, at which time the compilations were considered completed.

Acknowledgements

Throughout this work we have made extensive use of the bibliographical files and reprint collection maintained in the Atomic Energy Level Data Center by Dr. Romuald Zalubas. Our thanks are extended to him for generous cooperation. The compilation has also benefited greatly from the preprints that were provided by many of our colleagues.

We thank Dr. W. C. Martin for a critical reading of the manuscript.

This work was supported by the U.S. Department of Energy, Division of Magnetic Fusion Energy, the Office of Standard Reference Data of the National Bureau of Standards, and by the National Aeronautics and Space Administration, Astrophysics Division.

References for Introduction

- Bacher, R. F., and Goudsmit, S. (1932). *Atomic Energy States*, (McGraw-Hill Book Co., New York)

- Cohen, E. R., and Taylor, B. N. (1973), J. Phys. Chem. Ref. Data **2**, 663.
- Corliss, C., and Sugar, J. (1977), Energy Levels of Manganese, Mn I through Mn xxv, J. Phys. Chem. Ref. Data, **6**, 1253.
- Corliss, C., and Sugar, J. (1979), Energy Levels of Titanium, Ti I through Ti xxii, J. Phys. Chem. Ref. Data **8**, 1.
- Hagan, L. (1977), Bibliography on Atomic Energy Levels and Spectra, July 1971 through June 1975, Nat. Bur. Stand. (U.S.) Spec. Publ. 363, Suppl. 1 (U.S. Gov't Printing Office, Washington, D.C.).
- Hagan, L., and Martin, W. C. (1972), Bibliography on Atomic Energy Levels and Spectra, July 1968 through June 1971, Nat. Bur. Stand. (U.S.) Spec. Publ. 363 (U.S. Gov't Printing Office, Washington, D.C.).
- Kelly, R. L., and Palumbo, L. J. (1973), Atomic and Ionic Emission Lines Below 2000 Angstroms—Hydrogen Through Krypton, NRL Report 7599 (U.S. Gov't Printing Office, Washington, D.C.).
- Martin, W. C., Zalubas, R., and Hagan, L. (1978), Atomic Energy Levels—The Rare Earth Elements, Nat. Stand. Ref. Data Ser., Nat. Bur. Stand. (U.S.), 60.
- Moore, C. E. (1949), Atomic Energy Levels, Nat. Bur. Stand. (U.S.) Circ. 467, Vol. I (reissued as Nat. Stand. Ref. Data Ser., Nat. Bur. Stand. (U.S.), 35.)
- Moore, C. E. (1968), Bibliography on the Analyses of Optical Atomic Spectra, Section 1, Nat. Bur. Stand. (U.S.), Spec. Publ. 306 (U.S. Gov't Printing Office, Washington, D.C.).
- Nielson, C. W., and Koster, G. F. (1963), Spectroscopic Coefficients for the p^n , d^n , and f^n Configurations (The M.I.T. Press, Cambridge).
- Reader, J., and Sugar, J. (1975), Energy Levels of Iron, Fe I through Fe xxvi, J. Phys. Chem. Ref. Data **4**, 353.
- Sugar, J., and Corliss, C. (1977), Energy Levels of Chromium, Cr I through Cr xxiv, J. Phys. Chem. Ref. Data **6**, 317.
- Sugar, J., and Corliss, C. (1978), Energy Levels of Vanadium, V I through V xxiii, J. Phys. Chem. Ref. Data **7**, 1191.

Energy Level Tables

Ca I

$Z=20$

Ground state: $1s^2 2s^2 2p^6 3s^2 3p^6 4s^2 \ ^1S_0$

Ionization energy = $49\ 305.96 \pm 0.08\ \text{cm}^{-1}$ ($6.11321 \pm 0.00002\ \text{eV}$)

Early History

The observation and analysis of the first spectrum of calcium have played an important role in the development of experimental methods and theoretical ideas in the field of atomic spectra. Saunders (1920) summarized the considerable early work on this spectrum which included observations of the $4snp$, $4pns$, $4pnd$, and $3dnf$ series of triplets and singlets. The longest was $4snp\ ^1P^{\circ}$, reaching $n=11$. Shortly thereafter Russell and Saunders (1925), and independently Bohr (1923), recognized the excitation of two electrons to form terms above the ionization energy. In the same paper Russell and Saunders proposed the now universally accepted designations for levels in LS -coupling. They also introduced the concept of ordering level values from zero for the level of highest binding energy. The paper reports an extension of the analysis of Ca I which classified the remaining strong lines of this spectrum. The 3P terms of $4p^2$, $3d^2$, and $3dn$ ($n=4-6$) are given, the latter occurring entirely above the principal ionization limit. For the configuration $3d4p$ the terms $^3P^{\circ}$, $^3D^{\circ}$, and $^3F^{\circ}$ were found, and for $3d5p$ the $^3D^{\circ}$ was observed above the limit. Several undesigned levels are included whose interpretation was provided by later workers. In a later paper Russell (1927) identified the $4s5s\ ^3S_1$ and 1S_0 , the $4p\ ^1D$, and the $3d4d\ ^3S_1$ and 3D_1 .

Measurements of the Zeeman effect are due to Back (1925), who observed the $3d4s-3d4p$ transition array in a field of 39 000 gauss (3.9 T) and derived the following g-values for Ca I:

Config.	Term	J	g
3d4s	3D	1	0.502
		2	1.162
		3	1.328
	1D	2	1.007
		2	0.754
		3	1.076
		4	1.245
3d4p	$^3F^{\circ}$	2	0.893
	$^1D^{\circ}$	2	0.893

An additional term of the $3d^2$ configuration, the 3F , was discovered by Humphreys (1951) from observations in the infrared of the $3d4p-3d^2\ ^3F^{\circ}-^3F$ and $^3D^{\circ}-^3F$ multiplets.

Absorption measurements in the range of 1590–2400 Å by Garton and Codling (1965) produced 82 newly observed lines of calcium. Their classifications extended the known series as follows: $4snp\ ^1P^{\circ}$, to $n=33$, $3dn$ $^1P^{\circ}$, to $n=29$, $3dn$ $^3P^{\circ}$, to $n=25$, $3dn$ $^3D^{\circ}$, to $n=19$, $3dnf\ ^1P^{\circ}$, to $n=15$, $3dnf\ ^3P^{\circ}$, to $n=10$, and $3dnf\ ^3D^{\circ}$, to $n=15$. A single line at 1740.32 Å was classified as $4s^2\ ^1S_0-4s5p\ ^1P^{\circ}$. Their classification of the level at $43\ 933\ \text{cm}^{-1}$ as mainly $3d4p\ ^1P^{\circ}$, was accepted in several subsequent papers by other authors but was abandoned in the light of recent theoretical work cited below. It is now apparent that no level is preponderantly of this nature.

Sources of Level Values

The emission spectrum was completely reobserved by Risberg (1968) from 195 Å to 30 000 Å by means of a hollow cathode discharge. In this range 275 lines were measured, about 100 of which were observed for the first time. Autoionization and the low pressure of this light source prevented the detection of all but four levels above the ionization limit, 3P_1 and $^3D_{1,2,3}$ of the $3d4d$ configuration. The other members of the $3d4d$ and $3d5d$ configurations are derived from the classified lines of Russell and Saunders (1925) combined with the lower level values given by Risberg.

Several revisions and extensions are proposed in Risberg's paper. The earlier assignments of $4s6s\ ^1S_0$ and $4p^2\ ^1S_0$ to $41\ 786\ \text{cm}^{-1}$ and $40\ 690\ \text{cm}^{-1}$ respectively are interchanged. The $J=0$ level of $4s7p\ ^3P^{\circ}$ and the $J=1,2$ levels of $4s8p\ ^3P^{\circ}$ were found. The 1D_2 terms of $3d5s$ and $3d^2$ were also discovered, and in the $4snf$ series the 1F_3 of $4s12f$ was added.

Risberg has incorporated the interferometric measurements of the strong lines by Wagman (1937) and the accurate measurements of Grafenberger (1937) with her own extensive measurements to redetermine all the energy level values. Risberg's results are quoted here.

The data available for Ca I are particularly rich in long series. Those series for which the maximum observed value of n is 10 or more are listed below.

Brown, Tilford, and Ginter (1973) observed absorption spectra of calcium series in the range of 1500–1770 Å and 2020–2090 Å at high dispersion. They recorded the $4snp\ ^1P^{\circ}$ series for $n=11-79$ and deduced from these data an ionization energy of $49\ 305.99 \pm 0.12\ \text{cm}^{-1}$.

Series in Ca I

Config.	Term	<i>n</i> _{max}	References
1. 4snp	³ P°	60	Armstrong et al. (1979)
2. 4snp	¹ P°	79	Brown et al. (1973)
3. 4snd	³ D	17	Risberg (1968), Camus (1974)
4. 4snd	¹ D	62	Borgström and Rubbmark (1977), Armstrong et al. (1977)
5. 4sns	³ S	18	Risberg (1968), Camus (1974)
6. 4sns	¹ S	32	Borgström and Rubbmark (1977), Armstrong et al. (1977)
7. 3dnp	³ D°	58	Brown et al. (1973)
8. 3dnp	³ P°	33	Brown et al. (1973)
9. 3dnp	¹ P°	61	Brown et al. (1973)
10. 4snf	³ F°	13	Risberg (1968), Camus (1974)
11. 4snf	¹ F°	28	Borgström and Rubbmark (1977)
12. 3dnf	³ D°	30	Brown et al. (1973)
13. 3dnf	³ P°	38	Brown et al. (1973)
14. 3dnf	¹ P°	39	Brown et al. (1973)
15. 3p ⁴ (³ P° _{3,2}) 4s ² nd	² [3/2]°	16	Mansfield and Newsom (1977)
16. 3p ⁵ (² P° _{1,2}) 4s ² nd	² [3/2]°	17	Mansfield and Newsom (1977)
17. 3p ⁵ (² P° _{3,2}) 4s ² ns	² [3/2]°	10	Mansfield and Newsom (1977)
18. 3p ⁵ (² P° _{1,2}) 4s ² ns	² [1/2]°	14	Mansfield and Newsom (1977)

Series of doubly excited configurations converging to the 3d²D limit were extended. The 3dnp 3D°₁ series was measured through *n*=58, the 3dnp 1P°, through *n*=61, the 3dnp 3P°₁, through *n*=33, the 3dnf 1P°, through *n*=39, the 3dnf 3P°₁, through *n*=38 and the 3dnf 3D°, through *n*=30. They note that "the 3dnp 1P°, series lines are the strongest and most diffuse of the six observed series." Two levels that perturb this series at *n*=6 and *n*=7 were measured. They have been identified as 5s4p 3P°₁ and 1P°₁ by Newsom (1966). The results of Brown, Tilford, and Ginter are quoted here for series members beyond the observations of Risberg.

In a paper by Armstrong, Esherick, and Wynne (1979) they report observations by means of multiphoton absorption experiments of the 4snp 3P°₀ series to *n*=60. This was previously given to *n*=7 by Risberg (1968).

Armstrong, Esherick, and Wynne (1977) greatly extended the observations of the 4sns 1S₀ series from *n*=12 to 32 and the 4snd 1D₂ series from *n*=8 to 62. In the latter series they substituted the level 46 199.23 cm⁻¹ for 4s7d, rejecting the level 46 308.257 given by Risberg. The level 47 449.083 cm⁻¹ assigned by Risberg to 3d² 1D₂ is renamed 3d5s 1D₂. This designation was used by Risberg for the level 48 083.383 cm⁻¹, which is now named 4s10d 1D₂ by Armstrong et al. These changes are confirmed by additional observations by Palenius and Risberg (1977). No level was found to replace the 3d² 1D₂.

Borgström and Rubbmark (1977) observed absorption series from the 4s4p 1P°₁ level. They report values for 4sns 1S₀ from *n*=12 to 28 and 4snd 1D₂ from *n*=7 to 59. These two series have been extended to *n*=32 and *n*=62, respectively, by Armstrong, Esherick, and

Wynne (1977). Borgström and Rubbmark observed the 4snf 1F°₃ series from *n*=13 to 28.

Mansfield and Newsom (1977) have observed absorption from the 3p shell between 320 and 500 Å. We have given the 3p⁵4s²ns and 3p⁵4s²nd series from their paper.

Ionization Potential

Borgström and Rubbmark have determined a value of the ionization energy from the 4snf 1F°₃ series equal to 49 305.92 ± 0.10 cm⁻¹ which is within the error limits of the value of Brown et al., 49 305.99 ± 0.12 cm⁻¹. We have adopted the mean of the two values, 49 305.96 ± 0.08 cm⁻¹ as the best value. The uncertainty in the conversion factor, 8065.479 ± 0.021 cm⁻¹/eV, determines the uncertainty in the value given in eV.

Perturbed Series

The identification of 3d4p 1P°₁, which perturbs the 4snp 1P°₁ series, has been the subject of much discussion in the literature. Russell and Shenstone (1932) decided from intensity and energy considerations that it must be the level at 36 731 cm⁻¹. Roth (1969) attributes to Racah his identification of this term with the level at 41 679 cm⁻¹. Friedrich and Trefftz (1969), on the basis of multi-configuration calculations, retained it at 36 731 cm⁻¹. Garston and Codling (1965), from intensity considerations, placed it at 43 933 cm⁻¹. This assignment has been adopted by Risberg in her analysis of Ca I and by Moores (1966) in his multichannel quantum defect theory (MQDT) calculation of the 4snp 1P°₁ series. The investigation of the 4snp 1P°₁ series by Armstrong et al. (unpublished) using MQDT has shown that all of the above levels contain less than 30% 3dnp 1P°₁ composition (summed over all *n*). The 3d4p 1P°₁ state is diluted beyond recognition by mixing into the 4snp series. Among these levels the largest percentage of 3dnp 1P°₁ (27%) is present in the level at 43 933 cm⁻¹. Since this level is identified in the modern literature with 3d4p 1P°₁ and the rest of the quantum numbers are assigned accordingly, we retain this nominal labeling. In the light of this mixing, Roth's calculation of the (3d+4s)4p interaction is inappropriate and his results are not included here.

Another well-known perturbed series in Ca I is the 4snd 3D series, strongly perturbed by 3d5s 3D between *n*=8 and 9. This has been studied by Seaton (1966) using his MQDT method.

Some parity-forbidden series have been observed in Ca I. In Saunders' (1920) paper the first four members of 4s² 1S₀-4snd 1D₂ series are reported. Transitions to the ground state from 4p² 1S₀ and from 4s7s and 4s8s 1S₀ are also given. McIlrath (1974) has observed lines in absorption between 2048 and 2091 Å which Palenius and Risberg have identified as transitions to 4s9d-14d 1D₂ and 4s13s-17s 1S₀ from the 4s² 1S₀ ground state.

Arrangement of Tables

The first table of energy levels presented here for neutral calcium is arranged in the usual way; the terms are listed in order of increasing energy without regard to configuration assignments. Because many long Rydberg series have been observed in Ca I, we present a second table for this spectrum in which the series are listed separately, followed by their series limits. This table reveals more clearly the character of the observed spectrum. The series are listed in order of increasing first series member. The series member with the largest value of n for each term type is followed by the limit or limits of that series in Ca II. The assignment of the limits is discussed by Brown et al. and graphically displayed in their figure 10.

References

- Armstrong, J. A., Esherick, P., and Wynne, J. J. (1977), Phys. Rev. A **15**, 180.
 Armstrong, J. A., Esherick, P., and Wynne, J. J. (1979), J. Opt. Soc. Am. **69**, 211.
 Back, E. (1925), Zeits. f. Phys. **33**, 579.

- Bohr, N. This discovery was noted by Wentzel, G. (1923), Phys. Zeits. **24**, 106.
 Borgström, S. A., and Rubbmark, J. R. (1977), J. Phys. B **10**, 3607.
 Brown, C. M., Tilford, S. G., and Ginter, M. L. (1973), J. Opt. Soc. Am. **63**, 1454.
 Camus, P. (1974), J. Phys. B **7**, 1154, and unpublished data.
 Friedrich, H., and Trefftz, E. (1969), J. Quant. Spectrosc. Radiat. Transfer **9**, 333.
 Garton, W. R. S., and Codling, K. (1965), Proc. Phys. Soc. **86**, 1067.
 Grafenberger, H. (1937), Ann. d. Phys. [5] **30**, 267.
 Humphreys, C. J. (1951), J. Res. Nat. Bur. Stand. **47**, 262.
 Mansfield, M. W. D., and Newsom, G. H. (1977), Proc. R. Soc. London A **357**, 77.
 McIlrath, T. J. (1974), J. Phys. B **7**, 393.
 Moores, D. L. (1966), Proc. Phys. Soc. **88**, 843.
 Newsom, G. H. (1966), Proc. Phys. Soc. **87**, 975.
 Palenius, H. P., and Risberg, G. (1977), J. Phys. B **10**, L435.
 Risberg, G. (1968), Ark. Fys. **37**, 231.
 Roth, C. (1969), J. Res. Nat. Bur. Stand. **73A**, 497.
 Russell, H. N. (1927), Astrophys. J. **66**, 184.
 Russell, H. N., and Saunders, F. A. (1925), Astrophys. J. **61**, 38.
 Russell, H. N., and Shenstone, A. G. (1932), Phys. Rev. **39**, 415.
 Saunders, F. A. (1920), Astrophys. J. **52**, 256.
 Seaton, M. J. (1966), Proc. Phys. Soc. **88**, 815.
 Wagman, N. E. (1937), Univ. Pittsburgh Bull. **34**, 325.

Ca I: Ordered by term values

Configuration	Term	J	Level (cm $^{-1}$)	Configuration	Term	J	Level (cm $^{-1}$)
4s ²	¹ S	0	0.000	3d 4p	³ D°	1	38 192.392
4s4p	³ P°	0	15 157.901			2	38 219.118
		1	15 210.063			3	38 259.124
		2	15 315.943	4p ²	³ P	0	38 417.543
3d 4s	³ D	1	20 335.360			1	38 464.808
		2	20 349.260			2	38 551.558
		3	20 371.000	3d 4p	³ P°	0	39 333.382
3d 4s	¹ D	2	21 849.634			1	39 335.322
						2	39 340.080
4s4p	¹ P°	1	23 652.304	4s6s	³ S	1	40 474.241
4s5s	³ S	1	31 539.495	3d 4p	¹ F°	3	40 537.893
4s5s	¹ S	0	33 317.264	4s6s	¹ S	0	40 690.435
3d 4p	³ F°	2	35 730.454	4p ²	¹ D	2	40 719.847
		3	35 818.713				
		4	35 896.889	4s6p	¹ P°	1	41 679.008
3d 4p	¹ D°	2	35 835.413				
4s5p	³ P°	0	36 547.688	4p ²	¹ S	0	41 786.276
		1	36 554.749	4s4f	³ F°	2	42 170.214
		2	36 575.119			3	42 170.558
4s5p	¹ P°	1	36 731.615			4	42 171.026
4s4d	¹ D	2	37 298.287	4s4f	¹ F°	3	42 343.587
4s4d	³ D	1	37 748.197	4s6p	¹ P°	0	42 514.845
		2	37 751.867			1	42 518.708
		3	37 757.449			2	42 526.591

Ca I: Ordered by term values—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
4s5d	³ D	1	42 743.002	4s7f	³ F°	2	47 006.194
		2	42 744.716			3	47 006.280
		3	42 747.387			4	47 006.400
4s5d	¹ D	2	42 919.053	4s7f	¹ F°	3	47 015.141
3d ²	³ F	2	43 474.827	4s8d	³ D	1	47 036.225
		3	43 489.119			2	47 040.007
		4	43 508.088			3	47 045.241
3d 4p	¹ P°	1	43 933.477	4s9p	³ P°	1	47 085.38
4s7s	³ S	1	43 980.767	4s9p	¹ P°	1	47 184.370
4s7s	¹ S	0	44 276.538	4s10s	³ S	1	47 382.048
4s5f	³ F°	2	44 762.620	4s10s	¹ S	0	47 437.471
		3	44 762.839				
		4	44 763.118				
4s5f	¹ F°	3	44 804.878	3d 5s	¹ D	2	47 449.083
4s7p	³ P°	0	44 955.67	3d 5s	³ D	1	47 456.452
		1	44 957.655			2	47 466.014
		2	44 961.757			3	47 475.915
4s6d	¹ D	2	44 989.830	4s8f	³ F°	2	47 550.214
		3				3	47 550.271
		4				4	47 550.371
4s6d	³ D	1	45 049.073	4s8f	¹ F°	3	47 555.23
		2	45 050.419				
		3	45 052.374				
4s7p	¹ P°	1	45 425.358	4s10p	³ P°	1	47 604.75
4s8s	³ S	1	45 738.684	4s10p	¹ P°	1	47 662.10
4s8s	¹ S	0	45 887.200	4s9d	³ D	1	47 752.655
4s6f	³ F°	2	46 164.644	4s11s	³ S	1	47 757.286
		3	46 164.785			2	47 765.697
		4	46 164.971			3	
4s6f	¹ F°	3	46 182.399	4s9d	¹ D	2	47 806.17
4s7d	¹ D	2	46 200.13	4s11s	¹ S	0	47 812.39
4s8p	³ P°	0	46 284.12	4s9f	³ F°	2	47 843.76
		1	46 285.23			3	47 921.87
		2	46 287.63			4	47 921.981
4s7d	³ D	1	46 301.973	4s9f	¹ F°	3	47 922.033
		2	46 303.649				
		3	46 306.059				
4s8p	¹ P°	1	46 479.813	4s11p	¹ P°	1	47 997.49
4s9s	³ S	1	46 748.283	4s10d	³ D	1	48 031.58
4s9s	¹ S	0	46 835.055	4s10d	³ D	2	48 033.23
4s8d	¹ D	2	46 948.98	4s10d	¹ D	2	48 036.212
							48 083.41

Ca I: Ordered by term values—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
4s12s	³ S	1	48 103.93	4s15s	³ S	1	48 609.75
4s12s	¹ S	0	48 130.75	4s15s	¹ S	0	48 621.53
4s10f	³ F°	2	48 187.045	4s13f	³ F°	4	48 646.38
		3	48 187.075	4s13f	¹ F°	3	48 647.30
		4	48 187.118				
4s10f	¹ F°	3	48 188.990	4s15p	³ P°	2	48 660.23
4s12p	³ P°	1	48 215.81	4s15p	¹ P°	1	48 669.83
		2	48 216.36	4s14d	³ D	1	48 675.68
4s12p	¹ P°	1	48 240.53			2	48 675.87
						3	48 676.13
4s11d	³ D	1	48 258.30	4s14d	¹ D	2	48 678.97
		2	48 258.98				
		3	48 260.25	4s16s	³ S	1	48 708.67
4s11d	¹ D	2	48 290.85	4s16s	¹ S	0	48 718.02
4s13s	³ S	1	48 321.00	4s14f	¹ F°	3	48 738.54
4s13s	¹ S	0	48 340.75	4s16p	³ P°	2	48 749.04
4s11f	³ F°	2	48 382.70	4s16p	¹ P°	1	48 756.45
		3	48 382.781	4s15d	¹ D	2	48 760.14
		4	48 382.801	4s15d	³ D	1	48 761.11
4s11f	¹ F°	3	48 384.039	4s15d		2	48 761.21
						3	48 761.31
4s13p	³ P°	1	48 404.57				
		2	48 404.95	4s17s	³ S	1	48 787.89
4s13p	¹ P°	1	48 422.09	4s17s	¹ S	0	48 795.46
4s12d	³ D	1	48 433.23	4s15f	¹ F°	3	48 812.09
		2	48 433.65				
		3	48 434.36	4s17p	³ P°	2	48 820.60
4s12d	¹ D	2	48 451.73	4s17p	¹ P°	1	48 826.54
4s14s	³ S	1	48 484.12	4s11p			
4s14s	¹ S	0	48 499.14	4s16d	¹ D	2	48 827.05
3d ²	³ P	0	48 524.093	4s16d	³ D	2	48 830.41
		1	48 537.623			1	48 830.44
		2	48 563.522	4s18s	³ S	1	48 830.64
4s12f	³ F°	3,4	48 531.04	4s18s	¹ S	0	48 852.18
4s12f	¹ F°	3	48 532.139				
4s14p	³ P°	1	48 548.30	4s16f	¹ F°	3	48 858.59
		2	48 548.51	4s18p	³ P°	2	48 872.20
4s14p	¹ P°	1	48 561.10	4s17d	¹ D	2	48 879.31
4s13d	³ D	1	48 568.95	4s18p	¹ P°	1	48 882.37
		2	48 569.16				
		3	48 569.59	4s17d	³ D	3	48 884.06
4s13d	¹ D	2	48 578.32			2	48 887.58
							48 887.68

ENERGY LEVELS OF CALCIUM

Ca I: Ordered by term values—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
4s19s	¹ S	0	48 910.65	4s23f	¹ F°	3	49 096.74
4s17f	¹ F°	3	48 921.95	4s25p	³ P°	2	49 099.25
4s19p	³ P°	2	48 927.93	4s25p	¹ P°	1	49 100.72
4s18d	¹ D	2	48 928.79	4s26s	¹ S	0	49 109.85
4s19p	¹ P°	1	48 931.82	4s25d	¹ D	2	49 113.41
4s20s	¹ S	0	48 954.13	4s24f	¹ F°	3	49 113.85
4s18f	¹ F°	3	48 963.67	4s26p	³ P°	2	49 116.07
4s19d	¹ D	2	48 968.10	4s26p	¹ P°	1	49 117.38
4s20p	³ P°	2	48 968.67	4s27s	¹ S	0	49 125.44
4s20p	¹ P°	1	48 971.93	4s26d	¹ D	2	49 128.34
4s21s	¹ S	0	48 990.83	4s25f	¹ F°	3	49 129.00
4s19f	¹ F°	3	48 998.89	4s27p	³ P°	2	49 130.94
4s20d	¹ D	2	49 001.66	4s27p	¹ P°	1	49 132.09
4s21p	³ P°	2	49 003.21	4s28s	¹ S	0	49 139.31
4s21p	¹ P°	1	49 005.92	4s27d	¹ D	2	49 141.63
4s22s	¹ S	0	49 022.02	4s26f	¹ F°	3	49 142.38
4s20f	¹ F°	3	49 028.93	4s28p	³ P°	2	49 144.19
4s21d	¹ D	2	49 030.55	4s28p	¹ P°	1	49 145.15
4s22p	³ P°	2	49 032.70	4s29s	¹ S	0	49 151.59
4s22p	¹ P°	1	49 034.98	4s28d	¹ D	2	49 153.49
4s23s	¹ S	0	49 048.85	4s27f	¹ F°	3	49 154.20
4s21f	¹ F°	3	49 054.80	4s29p	³ P°	2	49 155.90
4s22d	¹ D	2	49 055.62	4s29p	¹ P°	1	49 156.78
4s23p	³ P°	2	49 058.02	4s30s	¹ S	0	49 162.47
4s23p	¹ P°	1	49 060.02	4s29d	¹ D	2	49 164.12
4s24s	¹ S	0	49 071.99	4s28f	¹ F°	3	49 164.99
4s22f	¹ F°	3	49 077.20	4s30p	³ P°	2	49 166.43
4s23d	¹ D	2	49 077.48	4s30p	¹ P°	1	49 167.20
4s24p	³ P°	2	49 080.04	4s31s	¹ S	0	49 172.43
4s24p	¹ P°	1	49 081.75	4s30d	¹ D	2	49 173.70
4s25s	¹ S	0	49 092.19	4s31p	³ P°	2	49 175.87
4s24d	¹ D	2	49 096.52	4s31p	¹ P°	1	49 176.58

Ca I: Ordered by term values—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
4s32s	¹ S	0	49 181.49	4s41d	¹ D	2	49 236.45
4s31d	¹ D	2	49 182.36	4s42p	³ P°	2	49 237.55
4s32p	³ P°	2	49 184.37	4s42p	¹ P°	1	49 237.82
4s32p	¹ P°	1	49 185.02	4s42d	¹ D	2	49 239.80
4s32d	¹ D	2	49 190.17	4s43p	³ P°	2	49 240.81
4s33p	³ P°	2	49 192.08	4s43p	¹ P°	1	49 241.09
4s33p	¹ P°	1	49 192.68	4s43d	¹ D	2	49 242.93
4s33d	¹ D	2	49 197.30	4s44p	³ P°	2	49 243.87
4s34p	³ P°	2	49 199.09	4s44p	¹ P°	1	49 244.13
4s34p	¹ P°	1	49 199.62	4s44d	¹ D	2	49 245.80
4s34d	¹ D	2	49 203.78	4s45p	³ P°	2	49 246.74
4s35p	³ P°	2	49 205.47	4s45p	¹ P°	1	49 246.98
4s35p	¹ P°	1	49 205.95	4s45d	¹ D	2	49 248.52
4s35d	¹ D	2	49 209.68	4s46p	³ P°	2	49 249.37
4s36p	³ P°	2	49 211.34	4s46p	¹ P°	1	49 249.61
4s36p	¹ P°	1	49 211.72	4s46d	¹ D	2	49 251.05
4s36d	¹ D	2	49 215.11	4s47p	³ P°	2	49 251.87
4s37p	³ P°	2	49 216.61	4s47p	¹ P°	1	49 252.08
4s37p	¹ P°	1	49 217.02	4s47d	¹ D	2	49 253.40
4s37d	¹ D	2	49 220.10	4s48p	³ P°	2	49 254.22
4s38p	³ P°	2	49 221.51	4s48p	¹ P°	1	49 254.41
4s38p	¹ P°	1	49 221.87	4s48d	¹ D	2	49 255.64
4s38d	¹ D	2	49 224.70	4s49p	³ P°	2	49 256.42
4s39p	³ P°	2	49 226.02	4s49d	¹ D	2	49 257.74
4s39p	¹ P°	1	49 226.35	4s50p	³ P°	2	49 258.45
4s39d	¹ D	2	49 228.92	4s50p	¹ P°	1	49 258.60
4s40p	³ P°	2	49 230.13	4s50d	¹ D	2	49 259.68
4s40p	¹ P°	1	49 230.45	4s51p	³ P°	2	49 260.36
4s40d	¹ D	2	49 232.81	4s51p	¹ P°	1	49 260.51
4s41p	³ P°	2	49 234.00	4s51d	¹ D	2	49 261.51
4s41p	¹ P°	1	49 234.28	4s52p	³ P°	2	49 262.15

ENERGY LEVELS OF CALCIUM

Ca I: Ordered by term values—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
4s52p	¹ P°	1	49 262.30	4s66p	¹ P°	1	49 279.30
4s52d	¹ D	2	49 263.31	4s67p	¹ P°	1	49 280.10
4s53p	³ P°	2	49 263.85	4s68p	¹ P°	1	49 280.88
4s53p	¹ P°	1	49 263.99	4s69p	¹ P°	1	49 281.62
4s53d	¹ D	2	49 264.77	4s70p	¹ P°	1	49 282.34
4s54p	³ P°	2	49 265.44	4s71p	¹ P°	1	49 283.01
4s54p	¹ P°	1	49 265.59	4s72p	¹ P°	1	49 283.67
4s54d	¹ D	2	49 266.36	4s73p	¹ P°	1	49 284.30
4s55p	³ P°	2	49 266.99	4s74p	¹ P°	1	49 284.88
4s55p	¹ P°	1	49 267.10	4s75p	¹ P°	1	49 285.46
4s55d	¹ D	2	49 267.84	4s76p	¹ P°	1	49 286.00
4s56p	³ P°	2	49 268.40	4s77p	¹ P°	1	49 286.54
4s56p	¹ P°	1	49 268.52	4s78p	¹ P°	1	49 287.05
4s56d	¹ D	2	49 269.33	4s79p	¹ P°	1	49 287.57
4s57p	³ P°	2	49 269.77	Ca II (² S _{1/2})		<i>Limit</i>	49 205.96
4s57p	¹ P°	1	49 269.88	3d 4d	³ D	1	51 351.74
4s57d	¹ D	2	49 270.40			2	51 369.38
4s58p	³ P°	2	49 271.01			3	51 396.32
4s58p	¹ P°	1	49 271.14	3d 4f	³ S	1	51 571.7
4s58d	¹ D	2	49 271.79	3d 4f	³ D°	1	51 709.5
4s59p	³ P°	2	49 272.24	3d 4f		2	51 734.6
4s59p	¹ P°	1	49 272.35	3d 6p	³ D°	1	51 767.0
4s59d	¹ D	2	49 273.09	3d 5d	³ P	0	54 282.3
4s60p	³ P°	2	49 273.37	3d 6p		1	54 288.74
4s60p	¹ P°	1	49 273.50	3d 5d	³ S	1	54 304.6
4s60d	¹ D	2	49 274.29	3d 6p	³ D°	1	55 902.8
4s61p	¹ P°	1	49 274.60	4p 5s	³ D°	1	55 946.6
4s61d	¹ D	2	49 275.58	3d 5d	³ P°	1	55 982.3
4s62p	¹ P°	1	49 275.62		¹ P°	1	56 254.
4s62d	¹ D	2	49 275.87	4p 5s	³ D°	1	56 469.0
4s63p	¹ P°	1	49 276.61	3d 5f	³ P°	1	56 532.63
4s64p	¹ P°	1	49 277.55		³ S	1	56 558.9
4s65p	¹ P°	1	49 278.44	3d 5f	³ D°	1	57 611.2
					³ P°	0	57 617.9
						1	57 638.4
						2	57 960.
					¹ P°	1	58 431.31
					³ D°	1	58 491.91
					³ P°	1	

Ca I: Ordered by term values—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
3d5f	¹ P°	1	58 505.89	3d11f	³ D°	1	62 009.17
3d7p	³ D°	1	58 798.92	3d11f	³ P°	1	62 052.49
3d7p	³ P°	1	59 010.87	3d13p	³ D°	1	62 084.92
3d7p	¹ P°	1	59 197.	3d11f	¹ P°	1	62 084.92
3d6d	³ P	1	59 368.	3d13p	³ P°	1	62 097.08
		2	59 391.	3d12f	¹ P°	1	62 154.62
3d6f	³ D°	1	59 802.21	3d13p	³ D°	1	62 161.61
3d6f	³ P°	1	59 862.26	3d12f	³ P°	1	62 198.92
3d6f	¹ P°	1	59 878.52	3d12f	³ P°	1	62 223.68
3d8p	³ D°	1	60 046.13	3d14p	³ D°	1	62 234.47
3d8p	³ P°	1	60 150.26	3d12f	¹ P°	1	62 237.12
3d8p	¹ P°	1	60 300.	3d14p	³ P°	1	62 279.33
3d7f	³ D°	1	60 632.18	3d13f	³ D°	1	62 312.63
3d7f	³ P°	1	60 690.27	3d13f	³ P°	1	62 331.04
3d7f	¹ P°	1	60 709.54	3d14p	¹ P°	1	62 346.47
3d9p	³ D°	1	60 807.28	3d13f	¹ P°	1	62 351.93
3d9p	³ P°	1	60 869.28	3d15p	³ D°	1	62 375.29
3d9p	¹ P°	1	60 973.6	3d13f	³ D°	1	62 392.01
3d8f	³ D°	1	61 172.54	3d15p	¹ P°	1	62 403.22
3d8f	³ P°	1	61 228.65	3d14f	³ P°	1	62 416.47
3d8f	¹ P°	1	61 251.39	3d16p	³ D°	1	62 433.38
3d10p	³ D°	1	61 306.65	3d14f	¹ P°	1	62 443.64
3d10p	³ P°	1	61 345.74	3d16p	³ P°	1	62 453.85
3d10p	¹ P°	1	61 428.1	3d15f	³ D°	1	62 472.14
3d9f	³ D°	1	61 543.95	3d15f	³ P°	1	62 478.26
3d9f	³ P°	1	61 596.69	3d16p	¹ P°	1	62 486.45
3d9f	¹ P°	1	61 622.60	3d17p	³ D°	1	62 506.62
3d11p	³ D°	1	61 652.25	3d15f	¹ P°	1	62 511.97
3d11p	³ P°	1	61 676.62	3d16p	³ P°	1	62 520.06
3d11p	¹ P°	1	61 747.2	3d17p	³ D°	1	62 533.90
3d10f	³ D°	1	61 810.68	3d15f	¹ P°	1	
3d10f	³ P°	1	61 859.22	3d16f	³ D°	1	
3d10f	¹ P°	1	61 888.22	3d17p	³ P°	1	
3d12p	³ D°	1	61 901.04	3d16f	¹ P°	1	
3d12p	³ P°	1	61 916.38	3d16f	³ D°	1	
3d12p	¹ P°	1	61 980.49	3d16f	³ P°	1	

ENERGY LEVELS OF CALCIUM

877

Ca I: Ordered by term values—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
3d 18p	³ D°	1	62 539.39	3d 22f	³ D°	1	62 720.41
3d 17p	¹ P°	1	62 548.16	3d 21p	¹ P°	1	62 721.80
4p4d	³ D°?	1	62 552.97	3d 22f	³ P°	1	62 729.19
4p4d	³ P°?	1	62 557.45	3d 24p	³ D°	1	62 734.25
4p4d	¹ P°?	1	62 561.72	3d 22p	³ P°	1	62 736.10
3d 17f	³ D°	1	62 565.68	3d 20f	¹ P°	1	62 736.10
3d 16f	¹ P°	1	62 572.66	3d 23f	³ D°	1	62 743.11
3d 17f	³ P°	1	62 578.91	3d 23f	³ P°	1	62 748.53
3d 18p	³ P°	1	62 584.50	3d 22p	¹ P°	1	62 750.58
3d 19p	³ D°	1	62 587.83	3d 25p	³ D°	1	62 753.60
3d 18p	¹ P°	1	62 602.72	3d 23p	³ P°	1	62 762.86
3d 18f	³ D°	1	62 610.75	3d 21f	¹ P°	1	62 762.86
3d 18f	³ P°	1	62 617.02	3d 24f	³ P°	1	62 764.10
3d 17f	¹ P°	1	62 625.37	3d 24f	³ D°	1	62 768.10
3d 20p	³ D°	1	62 626.03	3d 26p	³ D°	1	62 769.47
3d 19p	³ P°	1	62 628.76	3d 23p	¹ P°	1	62 774.75
3d 19f	³ D°	1	62 641.52	3d 25f	³ P°	1	62 780.30
3d 19p	¹ P°	1	62 648.37	3d 27p	³ D°	1	62 783.86
3d 19f	³ P°	1	62 653.99	3d 22f	¹ P°	1	62 785.02
3d 21p	³ D°	1	62 660.15	3d 24p	³ P°	1	62 787.31
3d 20p	³ P°	1	62 671.51	3d 26f	³ D°	1	62 790.08
3d 18f	¹ P°	1	62 671.51	3d 26f	³ P°	1	62 794.19
3d 20f	³ D°	1	62 675.56	3d 24p	¹ P°	1	62 795.55
3d 20f	³ P°	1	62 683.89	3d 28p	³ D°	1	62 797.46
3d 22p	³ D°	1	62 686.88	3d 23f	¹ P°	1	62 804.38
3d 20p	¹ P°	1	62 690.32	3d 27f	³ P°	1	62 806.02
3d 19f	¹ P°	1	62 704.43	3d 25p	³ P°	1	62 807.90
3d 21f	³ D°	1	62 704.43	3d 29p	³ D°	1	62 808.51
3d 21f	³ P°	1	62 707.25	3d 28f	³ D°	1	62 812.00
3d 21p	³ P°	1	62 711.37	3d 25p	¹ P°	1	62 814.13
3d 23p	³ D°	1	62 713.24				

Ca I: Ordered by term values—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
3d 28f	³ P°	1	62 816.51	3d 30p	³ P°	1	62 875.40
3d 30p	³ D°	1	62 818.76	3d 27f	³ P°	1	62 875.40
3d 24f	¹ P°	1	62 820.31	3d 39p	³ D°	1	62 876.98
3d 26p	³ P°	1	62 824.17			1	62 878.40
3d 29f	³ P°	1	62 824.17	3d 30p	¹ P°	1	62 879.48
3d 31p	³ D°	1	62 827.75	3d 38f	³ P°	1	62 880.04
3d 26p	¹ P°	1	62 830.67	3d 40p	³ D°	1	62 881.24
3d 30f	³ D°	1	62 832.31	3d 31p	³ P°	1	62 884.77
3d 30f	³ P°	1	62 834.11	3d 41p	³ D°	1	62 884.77
3d 32p	³ D°	1	62 836.22			1	62 886.45
3d 25f	¹ P°	1	62 836.85	3d 31p	¹ P°	1	62 889.00
3d 27p	³ P°	1	62 838.45	3d 42p	³ D°	1	62 889.00
3d 31f	³ P°	1	62 840.23	3d 43p	³ D°	1	62 891.65
3d 33p	³ D°	1	62 843.72	3d 32p	³ P°	1	62 893.66
3d 27p	¹ P°	1	62 845.27	3d 44p	³ D°	1	62 894.65
3d 32f	³ P°	1	62 848.77	3d 32p	¹ P°	1	62 897.57
3d 34p	³ D°	1	62 850.62	3d 45p	³ D°	1	62 897.57
3d 26f	¹ P°	1	62 852.33	3d 46p	³ D°	1	62 900.09
3d 28p	³ P°	1	62 853.79	3d 33p	³ P°	1	62 901.85
3d 33f	³ P°	1	62 853.79	3d 47p	³ D°	1	62 902.75
3d 35p	³ D°	1	62 856.56	3d 33p	¹ P°	1	62 904.95
3d 28p	¹ P°	1	62 858.11	3d 48p	³ D°	1	62 904.95
3d 34f	³ P°	1	62 861.06	3d 49p	³ D°	1	62 906.98
3d 36p	³ D°	1	62 862.60	3d 50p	³ D°	1	62 908.88
3d 29p	³ P°	1	62 863.99	3d 51p	³ D°	1	62 910.88
3d 35f	³ P°	1	62 865.39	3d 34p	¹ P°	1	62 911.51
3d 37p	³ D°	1	62 867.87	3d 52p	³ D°	1	62 912.67
3d 29p	¹ P°	1	62 869.38	3d 53p	³ D°	1	62 914.47
3d 36f	³ P°	1	62 871.46	3d 54p	³ D°	1	62 915.91
3d 38p	³ D°	1	62 872.69	3d 35p	¹ P°	1	62 917.83
				3d 55p	³ D°	1	62 917.83
				3d 56p	³ D°	1	62 918.80
				3d 57p	³ D°	1	62 920.32

ENERGY LEVELS OF CALCIUM
Ca II: Ordered by term values—Continued

879

Configuration	Term	J	Level (cm ⁻¹)	Configuration	Term	J	Level (cm ⁻¹)
3d 58p	³ D°	1	62 921.43	3p ⁵ 4s ² 4d	³ D°	1	267 956
3d 36p	¹ P°	1	62 923.62	3p ⁵ (² P _{1/2})4s ² 6s	² [_{1/2}]°	1	269 706
3d 35f	¹ P°	1	62 926.36	3p ⁵ 4s ² 4d	¹ P°	1	270 640
3d 37p	¹ P°	1	62 928.81	3p ⁵ (² P _{3/2})4s ² 5d	² [_{3/2}]°	1	270 889
3d 36f	¹ P°	1	62 931.29	3p ⁵ (² P _{3/2})4s ² 7s	² [_{3/2}]°	1	271 245
3d 38p	¹ P°	1	62 933.61	3p ⁵ (² P _{3/2})4s ² 6d	² [_{3/2}]°	1	273 134
3d 37f	¹ P°	1	62 935.88	3p ⁵ (² P _{1/2})4s ² 7s	² [_{1/2}]°	1	273 598
3d 39p	¹ P°	1	62 938.02	3p ⁵ (² P _{3/2})4s ² 7d	² [_{3/2}]°	1	273 959
3d 38f	¹ P°	1	62 939.79	3p ⁵ (² P _{1/2})4s ² 5d	² [_{3/2}]°	1	274 040
3d 40p	¹ P°	1	62 942.10	3p ⁵ (² P _{3/2})4s ² 9s	² [_{3/2}]°	1	274 276
3d 39f	¹ P°	1	62 943.85	3p ⁵ (² P _{3/2})4s ² 8d	² [_{3/2}]°	1	274 652
3d 41p	¹ P°	1	62 945.88	3p ⁵ (² P _{3/2})4s ² 10s	² [_{3/2}]°	1	274 777
3d 42p	¹ P°	1	62 949.34	3p ⁵ (² P _{3/2})4s ² 9d	² [_{3/2}]°	1	275 272
3d 43p	¹ P°	1	62 952.58				
3d 44p	¹ P°	1	62 955.57				
Ca II (² D _{3/2})	<i>Limit</i>		62 956.15	3p ⁵ (² P _{3/2})4s ² 10d	² [_{3/2}]°	1	275 503
3d 45p	¹ P°	1	62 958.38	3p ⁵ (² P _{3/2})4s ² 11d	² [_{3/2}]°	1	275 726
3d 46p	¹ P°	1	62 961.01	3p ⁵ (² P _{3/2})4s ² 12d	² [_{3/2}]°	1	275 898
3d 47p	¹ P°	1	62 963.43	3p ⁵ (² P _{3/2})4s ² 13d	² [_{3/2}]°	1	276 038
3d 48p	¹ P°	1	62 965.72	3p ⁵ (² P _{3/2})4s ² 14d	² [_{3/2}]°	1	276 138
3d 49p	¹ P°	1	62 967.84	3p ⁵ (² P _{3/2})4s ² 15d	² [_{3/2}]°	1	276 214
3d 50p	¹ P°	1	62 969.89	3p ⁵ (² P _{3/2})4s ² 16d	² [_{3/2}]°	1	276 283
3d 51p	¹ P°	1	62 971.78	3p ⁵ (² P _{1/2})4s ² 7d	² [_{3/2}]°	1	276 631
3d 52p	¹ P°	1	62 973.55	Ca II 3p ⁵ 4s ² ² P _{3/2}	<i>Limit</i>		276 750
3d 53p	¹ P°	1	62 975.23				
3d 54p	¹ P°	1	62 976.81				
3d 55p	¹ P°	1	62 978.26				
3d 56p	¹ P°	1	62 979.72				
3d 57p	¹ P°	1	62 981.03				
3d 58p	¹ P°	1	62 982.35				
3d 59p	¹ P°	1	62 983.48				
3d 60p	¹ P°	1	62 984.63				
3d 61p	¹ P°	1	62 985.70				
Ca II (² D _{5/2})	<i>Limit</i>		63 016.84	3p ⁵ (² P _{1/2})4s ² 9d	² [_{3/2}]°	1	277 917
3p ⁵ 4s ² 3d	³ P°	1	200 096	3p ⁵ (² P _{1/2})4s ² 11s	² [_{1/2}]°	1	277 991
3p ⁵ 4s ² 3d	³ D°	1	218 991	3p ⁵ (² P _{1/2})4s ² 10d	² [_{3/2}]°	1	278 228
3p ⁵ 4s ² 3d	¹ P°	1	253 310	3p ⁵ (² P _{1/2})4s ² 12s	² [_{1/2}]°	1	278 302
3p ⁵ 4s ² 4d	³ P°	1	255 022	3p ⁵ (² P _{1/2})4s ² 11d	² [_{3/2}]°	1	278 469
3p ⁵ (² P _{3/2})4s ² 5s	² [_{3/2}]°	1	257 737	3p ⁵ (² P _{1/2})4s ² 13s	² [_{1/2}]°	1	278 534
3p ⁵ (² P _{1/2})4s ² 5s	² [_{1/2}]°	1	260 193	3p ⁵ (² P _{1/2})4s ² 12d	² [_{3/2}]°	1	278 647
3p ⁵ (² P _{3/2})4s ² 6s	² [_{3/2}]°	1	267 417				

Ca I: Ordered by term values—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
$3p^5(^2P_{1/2}^o)4s^2 14s$	$^2[\frac{1}{2}]^o$	1	278 702	$3p^5(^2P_{1/2}^o)4s^2 16d$	$^2[\frac{3}{2}]^o$	1	279 050
$3p^5(^2P_{1/2}^o)4s^2 13d$	$^2[\frac{3}{2}]^o$	1	278 796	$3p^5(^2P_{1/2}^o)4s^2 17d$	$^2[\frac{3}{2}]^o$	1	279 117
$3p^5(^2P_{1/2}^o)4s^2 14d$	$^2[\frac{3}{2}]^o$	1	278 900	Ca II $3p^54s^2 ^2P_{1/2}^o$	<i>Limit</i>		279 530
$3p^5(^2P_{1/2}^o)4s^2 15d$	$^2[\frac{3}{2}]^o$	1	278 983				

ENERGY LEVELS OF CALCIUM

881

Ca I: Ordered by series

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
4s ²	¹ S	0	0.000	4s33p	³ P°	2	49 192.08
*****	***	*	*****	4s34p	³ P°	2	49 199.09
4s4p	³ P°	0	15 157.901	4s35p	³ P°	2	49 205.47
		1	15 210.063	4s36p	³ P°	2	49 211.34
		2	15 315.943	4s37p	³ P°	2	49 216.61
4s5p	³ P°	0	36 547.688	4s38p	³ P°	2	49 221.51
		1	36 554.749	4s39p	³ P°	2	49 226.02
		2	36 575.119	4s40p	³ P°	2	49 230.13
4s6p	³ P°	0	42 514.845	4s41p	³ P°	2	49 234.00
		1	42 518.708	4s42p	³ P°	2	49 237.55
		2	42 526.591	4s43p	³ P°	2	49 240.81
4s7p	³ P°	0	44 955.67	4s44p	³ P°	2	49 243.87
		1	44 957.655	4s45p	³ P°	2	49 246.74
		2	44 961.757	4s46p	³ P°	2	49 249.37
4s8p	³ P°	0	46 284.12	4s47p	³ P°	2	49 251.87
		1	46 285.23	4s48p	³ P°	2	49 254.22
		2	46 287.63	4s49p	³ P°	2	49 256.42
4s9p	³ P°	1	47 085.38	4s50p	³ P°	2	49 258.45
		2	47 086.99	4s51p	³ P°	2	49 260.36
4s10p	³ P°	1	47 604.75	4s52p	³ P°	2	49 262.15
		2	47 605.77	4s53p	³ P°	2	49 263.85
4s11p	³ P°	1	47 960.87	4s54p	³ P°	2	49 265.44
		2	47 961.53	4s55p	³ P°	2	49 266.99
4s12p	³ P°	1	48 215.81	4s56p	³ P°	2	49 268.40
		2	48 216.36	4s57p	³ P°	2	49 269.77
4s13p	³ P°	1	48 404.57	4s58p	³ P°	2	49 271.01
		2	48 404.95	4s59p	³ P°	2	49 272.24
4s14p	³ P°	1	48 548.30	4s60p	³ P°	2	49 273.37
		2	48 548.51	Ca II (² S _{1/2})		<i>Limit</i>	49 305.96
4s15p	³ P°	2	48 660.23	4s4p	¹ P°	1	23 652.304
4s16p	³ P°	2	48 749.04	4s5p	¹ P°	1	36 731.615
4s17p	³ P°	2	48 820.60	4s6p	¹ P°	1	41 679.008
4s18p	³ P°	2	48 879.31	4s7p	¹ P°	1	45 425.358
4s19p	³ P°	2	48 927.93	4s8p	¹ P°	1	46 179.813
4s20p	³ P°	2	48 968.67	4s9p	¹ P°	1	47 184.370
4s21p	³ P°	2	49 003.21	4s10p	¹ P°	1	47 662.10
4s22p	³ P°	2	49 032.70	4s11p	¹ P°	1	47 997.49
4s23p	³ P°	2	49 058.02	4s12p	¹ P°	1	48 240.53
4s24p	³ P°	2	49 080.04	4s13p	¹ P°	1	48 422.09
4s25p	³ P°	2	49 099.25	4s14p	¹ P°	1	48 561.10
4s26p	³ P°	2	49 116.07	4s15p	¹ P°	1	48 669.83
4s27p	³ P°	2	49 130.94	4s16p	¹ P°	1	48 756.45
4s28p	³ P°	2	49 144.19	4s17p	¹ P°	1	48 826.54
4s29p	³ P°	2	49 155.90	4s18p	¹ P°	1	48 884.06
4s30p	³ P°	2	49 166.43	4s19p	¹ P°	1	48 931.82
4s31p	³ P°	2	49 175.87	4s20p	¹ P°	1	48 971.93
4s32p	³ P°	2	49 184.37	4s21p	¹ P°	1	49 005.92
				4s22p	¹ P°	1	49 034.98
				4s23p	¹ P°	1	49 060.02
				4s24p	¹ P°	1	49 081.75
				4s25p	¹ P°	1	49 100.72
				4s26p	¹ P°	1	49 117.38
				4s27p	¹ P°	1	49 132.09
				4s28p	¹ P°	1	49 145.15
				4s29p	¹ P°	1	49 156.78
				4s30p	¹ P°	1	49 167.20
				4s31p	¹ P°	1	49 176.58
				4s32p	¹ P°	1	49 185.02

Ca I: Ordered by series—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
4s33p	¹ P°	1	49 192.68	4s5d	³ D	1	42 743.002
4s34p	¹ P°	1	49 199.62			2	42 744.716
4s35p	¹ P°	1	49 205.95			3	42 747.387
4s36p	¹ P°	1	49 211.72	4s6d	³ D	1	45 049.078
4s37p	¹ P°	1	49 217.02			2	45 050.419
4s38p	¹ P°	1	49 221.87			3	45 052.374
4s39p	¹ P°	1	49 226.35	4s7d	³ D	1	46 301.978
4s40p	¹ P°	1	49 230.45			2	46 303.649
4s41p	¹ P°	1	49 234.28			3	46 306.059
4s42p	¹ P°	1	49 237.82	4s8d	³ D	1	47 036.225
4s43p	¹ P°	1	49 241.09			2	47 040.007
4s44p	¹ P°	1	49 244.13			3	47 045.241
4s45p	¹ P°	1	49 246.98	4s9d	³ D	1	47 752.655
4s46p	¹ P°	1	49 249.61			2	47 757.286
4s47p	¹ P°	1	49 252.08			3	47 765.697
4s48p	¹ P°	1	49 254.41	4s10d	³ D	1	48 031.58
4s49p	¹ P°	1	49 256.56			2	48 033.23
4s50p	¹ P°	1	49 258.60			3	48 036.212
4s51p	¹ P°	1	49 260.51	4s11d	³ D	1	48 258.30
4s52p	¹ P°	1	49 262.30			2	48 258.98
4s53p	¹ P°	1	49 263.99			3	48 260.25
4s54p	¹ P°	1	49 265.59	4s12d	³ D	1	48 433.23
4s55p	¹ P°	1	49 267.10			2	48 433.65
4s56p	¹ P°	1	49 268.52			3	48 434.36
4s57p	¹ P°	1	49 269.88	4s13d	³ D	1	48 568.95
4s58p	¹ P°	1	49 271.14			2	48 569.16
4s59p	¹ P°	1	49 272.35			3	48 569.59
4s60p	¹ P°	1	49 273.50	4s14d	³ D	1	48 675.68
4s61p	¹ P°	1	49 274.60			2	48 675.87
4s62p	¹ P°	1	49 275.62			3	48 676.13
4s63p	¹ P°	1	49 276.61	4s15d	³ D	1	48 761.11
4s64p	¹ P°	1	49 277.55			2	48 761.21
4s65p	¹ P°	1	49 278.44			3	48 761.31
4s66p	¹ P°	1	49 279.30	4s16d	³ D	2	48 830.41
4s67p	¹ P°	1	49 280.10			1	48 830.44
4s68p	¹ P°	1	49 280.88			3	48 830.64
4s69p	¹ P°	1	49 281.62	4s17d	³ D	3	48 887.58
4s70p	¹ P°	1	49 282.34			2	48 887.68
4s71p	¹ P°	1	49 283.01	Ca II (² S _{1/2})	<i>Limit</i>		49 305.96
4s72p	¹ P°	1	49 283.67				
4s73p	¹ P°	1	49 284.30	Ca II (² S _{1/2})	<i>Limit</i>		49 305.96
4s74p	¹ P°	1	49 284.88				
4s75p	¹ P°	1	49 285.46	4s3d	¹ D	2	21 849.634
4s76p	¹ P°	1	49 286.00	4s4d	¹ D	2	37 298.287
4s77p	¹ P°	1	49 286.54	4s5d	¹ D	2	42 919.053
4s78p	¹ P°	1	49 287.05	4s6d	¹ D	2	44 989.830
4s79p	¹ P°	1	49 287.57	4s7d	¹ D	2	46 200.13
				4s8d	¹ D	2	46 948.98
Ca II (² S _{1/2})	<i>Limit</i>		49 305.96				
4s3d	³ D	1	20 335.360				
		2	20 349.260	4s3d	¹ D	2	
		3	20 371.000	4s4d	¹ D	2	
4s4d	³ D	1	37 748.197	4s5d	¹ D	2	
		2	37 751.867	4s6d	¹ D	2	
		3	37 757.449	4s7d	¹ D	2	
				4s8d	¹ D	2	

ENERGY LEVELS OF CALCIUM

Ca I: Ordered by series—Continued

Configuration	Term	J	Level (cm ⁻¹)	Configuration	Term	J	Level (cm ⁻¹)
4s9d	¹ D	2	47 812.39	4s8s	³ S	1	45 738.684
4s10d	¹ D	2	48 083.41	4s9s	³ S	1	46 748.283
4s11d	¹ D	2	48 290.85	4s10s	³ S	1	47 382.048
4s12d	¹ D	2	48 451.73	4s11s	³ S	1	47 806.17
4s13d	¹ D	2	48 578.32	4s12s	³ S	1	48 103.93
4s14d	¹ D	2	48 678.97	4s13s	³ S	1	48 321.00
4s15d	¹ D	2	48 760.14	4s14s	³ S	1	48 484.12
4s16d	¹ D	2	48 827.05	4s15s	³ S	1	48 609.75
4s17d	¹ D	2	48 882.87	4s16s	³ S	1	48 708.67
4s18d	¹ D	2	48 928.79	4s17s	³ S	1	48 787.89
4s19d	¹ D	2	48 968.10	4s18s	³ S	1	48 852.18
4s20d	¹ D	2	49 001.66				
4s21d	¹ D	2	49 030.55	Ca II (² S _{1/2})		<i>Limit</i>	49 305.96
4s22d	¹ D	2	49 055.62				
4s23d	¹ D	2	49 077.48	4s5s	¹ S	0	33 317.264
4s24d	¹ D	2	49 096.52	4s6s	¹ S	0	40 690.435
4s25d	¹ D	2	49 113.41	4s7s	¹ S	0	44 276.538
4s26d	¹ D	2	49 128.34	4s8s	¹ S	0	45 887.200
4s27d	¹ D	2	49 141.63	4s9s	¹ S	0	46 835.055
4s28d	¹ D	2	49 153.49	4s10s	¹ S	0	47 437.471
4s29d	¹ D	2	49 164.12	4s11s	¹ S	0	47 843.76
4s30d	¹ D	2	49 173.70	4s12s	¹ S	0	48 130.75
4s31d	¹ D	2	49 182.36	4s13s	¹ S	0	48 340.75
4s32d	¹ D	2	49 190.17	4s14s	¹ S	0	48 499.14
4s33d	¹ D	2	49 197.30	4s15s	¹ S	0	48 621.53
4s34d	¹ D	2	49 203.78	4s16s	¹ S	0	48 718.02
4s35d	¹ D	2	49 209.68	4s17s	¹ S	0	48 795.46
4s36d	¹ D	2	49 215.11	4s18s	¹ S	0	48 858.59
4s37d	¹ D	2	49 220.10	4s19s	¹ S	0	48 910.65
4s38d	¹ D	2	49 224.70	4s20s	¹ S	0	48 954.13
4s39d	¹ D	2	49 228.92	4s21s	¹ S	0	48 990.83
4s40d	¹ D	2	49 232.81	4s22s	¹ S	0	49 022.02
4s41d	¹ D	2	49 236.45	4s23s	¹ S	0	49 048.85
4s42d	¹ D	2	49 239.80	4s24s	¹ S	0	49 071.99
4s43d	¹ D	2	49 242.93	4s25s	¹ S	0	49 092.19
4s44d	¹ D	2	49 245.80	4s26s	¹ S	0	49 109.85
4s45d	¹ D	2	49 248.52	4s27s	¹ S	0	49 125.44
4s46d	¹ D	2	49 251.05	4s28s	¹ S	0	49 139.81
4s47d	¹ D	2	49 253.40	4s29s	¹ S	0	49 151.59
4s48d	¹ D	2	49 255.64	4s30s	¹ S	0	49 162.47
4s49d	¹ D	2	49 257.74	4s31s	¹ S	0	49 172.43
4s50d	¹ D	2	49 259.68	4s32s	¹ S	0	49 181.49
4s51d	¹ D	2	49 261.51				
4s52d	¹ D	2	49 263.31	Ca II (² S _{1/2})		<i>Limit</i>	49 305.96
4s53d	¹ D	2	49 264.77				
4s54d	¹ D	2	49 266.36	3d 4p	³ F°	2	35 730.454
4s55d	¹ D	2	49 267.84			3	35 818.713
4s56d	¹ D	2	49 269.33			4	35 896.889
4s57d	¹ D	2	49 270.40				
4s58d	¹ D	2	49 271.79	3d 4p	¹ D°	2	35 835.443
4s59d	¹ D	2	49 273.09				
4s60d	¹ D	2	49 274.29	3d 4p	³ D°	1	38 192.392
4s61d	¹ D	2	49 275.58			2	38 219.118
4s62d	¹ D	2	49 275.87			3	38 259.124
Ca II (² S _{1/2})	<i>Limit</i>		49 305.96	3d 4p	³ P°	0	39 339.382
4s5s	³ S	1	31 539.495			1	39 335.322
4s6s	³ S	1	40 474.241			2	39 340.080
4s7s	³ S	1	43 980.767	3d 4p	¹ F°	3	40 537.893

Ca I: Ordered by series—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
3d 4p	¹ P°	1	43 933.477	3d 55p	³ D°	1	62 917.83
3d 5p	³ D°	1	51 709.5	3d 56p	³ D°	1	62 918.80
		2	51 734.6	3d 57p	³ D°	1	62 920.32
		3	51 767.0	3d 58p	³ D°	1	62 921.43
				Ca II (² D _{3/2})		<i>Limit</i>	62 956.15
3d 5p	³ P°	1	51 908.	3d 6p	³ P°	1	56 532.63
3d 5p	¹ P°	1	53 100.	3d 7p	³ P°	1	59 010.87
3d 6p	³ D°	1	56 254.	3d 8p	³ P°	1	60 150.26
3d 7p	³ D°	1	58 798.92	3d 9p	³ P°	1	60 869.28
3d 8p	³ D°	1	60 046.13	3d 10p	³ P°	1	61 345.74
3d 9p	³ D°	1	60 807.28	3d 11p	³ P°	1	61 676.62
3d 10p	³ D°	1	61 306.65	3d 12p	³ P°	1	61 916.38
3d 11p	³ D°	1	61 652.25	3d 13p	³ P°	1	62 097.08
3d 12p	³ D°	1	61 901.04	3d 14p	³ P°	1	62 237.12
3d 13p	³ D°	1	62 084.92	3d 15p	³ P°	1	62 351.93
3d 14p	³ D°	1	62 223.68	3d 16p	³ P°	1	62 443.64
3d 15p	³ D°	1	62 331.04	3d 17p	³ P°	1	62 511.97
3d 16p	³ D°	1	62 416.47	3d 18p	³ P°	1	62 584.50
3d 17p	³ D°	1	62 486.45	3d 19p	³ P°	1	62 628.76
3d 18p	³ D°	1	62 539.39	3d 20p	³ P°	1	62 671.51
3d 19p	³ D°	1	62 587.83	3d 21p	³ P°	1	62 711.37
3d 20p	³ D°	1	62 626.03	3d 22p	³ P°	1	62 736.10
3d 21p	³ D°	1	62 660.15	3d 23p	³ P°	1	62 762.86
3d 22p	³ D°	1	62 686.88	3d 24p	³ P°	1	62 787.31
3d 23p	³ D°	1	62 713.24	3d 25p	³ P°	1	62 807.90
3d 24p	³ D°	1	62 734.25	3d 26p	³ P°	1	62 824.17
3d 25p	³ D°	1	62 753.60	3d 27p	³ P°	1	62 838.45
3d 26p	³ D°	1	62 769.47	3d 28p	³ P°	1	62 853.79
3d 27p	³ D°	1	62 783.86	3d 29p	³ P°	1	62 863.99
3d 28p	³ D°	1	62 797.46	3d 30p	³ P°	1	62 875.40
3d 29p	³ D°	1	62 808.51	3d 31p	³ P°	1	62 884.77
3d 30p	³ D°	1	62 818.76	3d 32p	³ P°	1	62 893.66
3d 31p	³ D°	1	62 827.75	3d 33p	³ P°	1	62 901.85
3d 32p	³ D°	1	62 836.22	Ca II (² D _{5/2})		<i>Limit</i>	63 016.84
3d 33p	³ D°	1	62 843.72				
3d 34p	³ D°	1	62 850.62	3d 6p	¹ P°	1	56 651.
3d 35p	³ D°	1	62 856.56	3d 7p	¹ P°	1	59 197.
3d 36p	³ D°	1	62 862.60	3d 8p	¹ P°	1	60 300.
3d 37p	³ D°	1	62 867.87	3d 9p	¹ P°	1	60 973.6
3d 38p	³ D°	1	62 872.69	3d 10p	¹ P°	1	61 428.1
3d 39p	³ D°	1	62 876.98	3d 11p	¹ P°	1	61 747.2
3d 40p	³ D°	1	62 881.24	3d 12p	¹ P°	1	61 980.49
3d 41p	³ D°	1	62 884.77	3d 13p	¹ P°	1	62 154.62
3d 42p	³ D°	1	62 889.00	3d 14p	¹ P°	1	62 288.02
3d 43p	³ D°	1	62 891.65	3d 15p	¹ P°	1	62 392.01
3d 44p	³ D°	1	62 894.65	3d 16p	¹ P°	1	62 478.26
3d 45p	³ D°	1	62 897.57	3d 17p	¹ P°	1	62 548.16
3d 46p	³ D°	1	62 900.09	3d 18p	¹ P°	1	62 602.72
3d 47p	³ D°	1	62 902.75	3d 19p	¹ P°	1	62 648.37
3d 48p	³ D°	1	62 904.95	3d 20p	¹ P°	1	62 690.32
3d 49p	³ D°	1	62 906.98	3d 21p	¹ P°	1	62 721.80
3d 50p	³ D°	1	62 908.88	3d 22p	¹ P°	1	62 750.58
3d 51p	³ D°	1	62 910.88	3d 23p	¹ P°	1	62 774.75
3d 52p	³ D°	1	62 912.67	3d 24p	¹ P°	1	62 795.55
3d 53p	³ D°	1	62 914.47	3d 25p	¹ P°	1	62 814.13
3d 54p	³ D°	1	62 915.91	3d 26p	¹ P°	1	62 830.67

Ca I: Ordered by series—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
3d 27p	¹ P°	1	62 845.27	4s7f	³ F°	2	47 006.194
3d 28p	¹ P°	1	62 858.11			3	47 006.280
3d 29p	¹ P°	1	62 869.38			4	47 006.400
3d 30p	¹ P°	1	62 879.48				
3d 31p	¹ P°	1	62 889.00	4s8f	³ F°	2	47 550.214
3d 32p	¹ P°	1	62 897.57			3	47 550.271
3d 33p	¹ P°	1	62 904.95			4	47 550.371
3d 34p	¹ P°	1	62 911.51				
3d 35p	¹ P°	1	62 917.83	4s9f	³ F°	2	47 921.87
3d 36p	¹ P°	1	62 923.62			3	47 921.981
3d 37p	¹ P°	1	62 928.81			4	47 922.033
3d 38p	¹ P°	1	62 933.61				
3d 39p	¹ P°	1	62 938.02	4s10f	³ F°	2	48 187.045
3d 40p	¹ P°	1	62 942.10			3	48 187.075
3d 41p	¹ P°	1	62 945.88			4	48 187.118
3d 42p	¹ P°	1	62 949.34				
3d 43p	¹ P°	1	62 952.58	4s11f	³ F°	2	48 382.70
3d 44p	¹ P°	1	62 955.57			3	48 382.781
3d 45p	¹ P°	1	62 958.38			4	48 382.801
3d 46p	¹ P°	1	62 961.01				
3d 47p	¹ P°	1	62 963.43	4s12f	³ F°	3,4	48 531.04
3d 48p	¹ P°	1	62 965.72				
3d 49p	¹ P°	1	62 967.84	4s13f	³ F°	4	48 646.38
3d 50p	¹ P°	1	62 969.89				
3d 51p	¹ P°	1	62 971.78	Ca II (² S _{1/2})	<i>Limit</i>		49 305.96
3d 52p	¹ P°	1	62 973.55				
3d 53p	¹ P°	1	62 975.23				
3d 54p	¹ P°	1	62 976.81	4s4f	¹ F°	3	42 343.587
3d 55p	¹ P°	1	62 978.26	4s5f	¹ F°	3	44 804.878
3d 56p	¹ P°	1	62 979.72	4s6f	¹ F°	3	46 182.399
3d 57p	¹ P°	1	62 981.03	4s7f	¹ F°	3	47 015.141
3d 58p	¹ P°	1	62 982.35	4s8f	¹ F°	3	47 555.23
3d 59p	¹ P°	1	62 983.48	4s9f	¹ F°	3	47 924.947
3d 60p	¹ P°	1	62 984.63	4s10f	¹ F°	3	48 188.990
3d 61p	¹ P°	1	62 985.70	4s11f	¹ F°	3	48 384.039
				4s12f	¹ F°	3	48 532.139
Ca II (² D _{5/2})	<i>Limit</i>		63 016.84	4s13f	¹ F°	3	48 647.30
4p ²	³ P	0	38 417.543	4s14f	¹ F°	3	48 738.54
		1	38 464.808	4s15f	¹ F°	3	48 812.09
		2	38 551.558	4s16f	¹ F°	3	48 872.20
				4s17f	¹ F°	3	48 921.95
				4s18f	¹ F°	3	48 963.67
4p ²	¹ D	2	40 719.847	4s19f	¹ F°	3	48 998.89
				4s20f	¹ F°	3	49 028.93
4p ²	¹ S	0	41 786.276	4s21f	¹ F°	3	49 054.80
				4s22f	¹ F°	3	49 077.20
*****	***	*	*****	4s23f	¹ F°	3	49 096.74
4s4f	³ F°	2	42 170.214	4s24f	¹ F°	3	49 113.85
		3	42 170.558	4s25f	¹ F°	3	49 129.00
		4	42 171.026	4s26f	¹ F°	3	49 142.38
				4s27f	¹ F°	3	49 154.29
				4s28f	¹ F°	3	49 164.99
4s5f	³ F°	2	44 762.620				
		3	44 762.839				
		4	44 763.118	Ca II (² S _{1/2})	<i>Limit</i>		49 305.96
4s6f	³ F°	2	46 164.644	3d ²	³ F	2	43 474.827
		3	46 164.785			3	43 489.119
		4	46 164.971			4	43 508.088

Ca I: Ordered by series—Continued

Configuration	Term	J	Level (cm ⁻¹)	Configuration	Term	J	Level (cm ⁻¹)
3d ²	³ P	0	48 524.093	3d 26f	³ D°	1	62 790.08
		1	48 537.623	3d 28f	³ D°	1	62 812.00
		2	48 563.522	3d 30f	³ D°	1	62 832.31
*****	***	*	*****	Ca II (² D _{3/2})	<i>Limit</i>		62 956.15
3d 5s	¹ D	2	47 449.083	3d4f	³ P°	1	55 946.6
3d 5s	³ D	1	47 456.452	3d5f	³ P°	1	58 491.91
		2	47 466.014	3d6f	³ P°	1	59 862.26
		3	47 475.915	3d7f	³ P°	1	60 690.27
*****	***	*	*****	3d8f	³ P°	1	61 228.65
				3d9f	³ P°	1	61 596.69
				3d 10f	³ P°	1	61 859.22
*****	***	*	*****	3d 11f	³ P°	1	62 052.49
3d 4d	³ D	1	51 351.74	3d 12f	³ P°	1	62 198.92
		2	51 369.88	3d 13f	³ P°	1	62 312.63
		3	51 396.32	3d 14f	³ P°	1	62 403.22
3d 4d	³ S	1	51 571.7	3d 15f	³ P°	1	62 472.14
3d 4d	³ P	0	54 282.3	3d 16f	³ P°	1	62 533.90
		1	54 288.74	3d 17f	³ P°	1	62 578.91
		2	54 304.6	3d 18f	³ P°	1	62 617.02
3d 5d	³ D	2	56 469.0	3d 19f	³ P°	1	62 653.99
3d 5d	³ S	1	56 558.9	3d 20f	³ P°	1	62 683.89
3d 5d	³ P	0	57 611.2	3d 21f	³ P°	1	62 707.25
		1	57 617.9	3d 22f	³ P°	1	62 729.19
		2	57 638.4	3d 23f	³ P°	1	62 748.53
3d 6d	³ P	1	59 368.	3d 24f	³ P°	1	62 764.10
		2	59 391.	3d 25f	³ P°	1	62 780.30
*****	***	*	*****	3d 26f	³ P°	1	62 794.19
3d4f	³ D°	1	55 902.8	3d 27f	³ P°	1	62 806.02
3d5f	³ D°	1	58 431.31	3d 28f	³ P°	1	62 816.51
3d6f	³ D°	1	59 802.21	3d 29f	³ P°	1	62 824.17
3d7f	³ D°	1	60 632.18	3d 30f	³ P°	1	62 834.11
3d8f	³ D°	1	61 172.54	3d 31f	³ P°	1	62 840.23
3d9f	³ D°	1	61 543.95	3d 32f	³ P°	1	62 848.77
3d 10f	³ D°	1	61 810.68	3d 33f	³ P°	1	62 853.79
3d 11f	³ D°	1	62 009.17	3d 34f	³ P°	1	62 861.06
3d 12f	³ D°	1	62 161.61	3d 35f	³ P°	1	62 865.39
3d 13f	³ D°	1	62 279.33	3d 36f	³ P°	1	62 871.46
3d 14f	³ D°	1	62 375.29	3d 37f	³ P°	1	62 875.40
3d 15f	³ D°	1	62 453.85	3d 38f	³ P°	1	62 880.04
3d 16f	³ D°	1	62 520.06	Ca II (² D _{3/2})	<i>Limit</i>		62 956.15
3d 17f	³ D°	1	62 565.68	3d4f	¹ P°	1	55 982.3
3d 18f	³ D°	1	62 610.75	3d5f	¹ P°	1	58 505.89
3d 19f	³ D°	1	62 641.52	3d6f	¹ P°	1	59 878.52
3d 20f	³ D°	1	62 675.56	3d7f	¹ P°	1	60 709.54
3d 21f	³ D°	1	62 704.43	3d8f	¹ P°	1	61 251.39
3d 22f	³ D°	1	62 720.41	3d9f	¹ P°	1	61 622.60
3d 23f	³ D°	1	62 743.11	3d 10f	¹ P°	1	61 888.22
3d 24f	³ D°	1	62 768.10	3d 11f	¹ P°	1	62 084.92
				3d 12f	¹ P°	1	62 234.47
				3d 13f	¹ P°	1	62 346.47
				3d 14f	¹ P°	1	62 433.38
				3d 15f	¹ P°	1	62 506.62
				3d 16f	¹ P°	1	62 572.66
				3d 17f	¹ P°	1	62 625.37
				3d 18f	¹ P°	1	62 671.51

Ca I: Ordered by series—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
3d 19f	¹ P°	1	62 704.43	3p ⁵ (² P _{3/2})4s ² 5s	² [³ / ₂] ^o	1	257 737
3d 20f	¹ P°	1	62 736.10	3p ⁵ (² P _{3/2})4s ² 6s	² [³ / ₂] ^o	1	267 417
3d 21f	¹ P°	1	62 762.86	3p ⁵ (² P _{3/2})4s ² 7s	² [³ / ₂] ^o	1	271 245
3d 22f	¹ P°	1	62 785.02	3p ⁵ (² P _{3/2})4s ² 9s	² [³ / ₂] ^o	1	274 276
3d 23f	¹ P°	1	62 804.38	3p ⁵ (² P _{3/2})4s ² 10s	² [³ / ₂] ^o	1	274 777
3d 24f	¹ P°	1	62 820.31				
3d 25f	¹ P°	1	62 836.85	Ca II 3p ⁵ 4s ² ² P _{3/2}	Limit		276 750
3d 26f	¹ P°	1	62 852.33				
3d 35f	¹ P°	1	62 926.36	3p ⁵ (² P _{1/2})4s ² 5s	² [¹ / ₂] ^o	1	260 193
3d 36f	¹ P°	1	62 931.29	3p ⁵ (² P _{1/2})4s ² 6s	² [¹ / ₂] ^o	1	269 706
3d 37f	¹ P°	1	62 935.88	3p ⁵ (² P _{1/2})4s ² 7s	² [¹ / ₂] ^o	1	273 598
3d 38f	¹ P°	1	62 939.79	3p ⁵ (² P _{1/2})4s ² 11s	² [¹ / ₂] ^o	1	277 991
3d 39f	¹ P°	1	62 943.85	3p ⁵ (² P _{1/2})4s ² 12s	² [¹ / ₂] ^o	1	278 302
Ca II (² D _{5/2})	Limit		63 016.84	3p ⁵ (² P _{1/2})4s ² 13s	² [¹ / ₂] ^o	1	278 534
				3p ⁵ (² P _{1/2})4s ² 14s	² [¹ / ₂] ^o	1	278 702
4p5s	³ P°	1	57 462.	Ca II 3p ⁵ 4s ² ² P _{1/2}	Limit		279 530
4p5s	¹ P°	1	57 960.	3p ⁵ (² P _{3/2})4s ² 5d	² [³ / ₂] ^o	1	270 889
*****	***	*	*****	3p ⁵ (² P _{3/2})4s ² 6d	² [³ / ₂] ^o	1	273 134
				3p ⁵ (² P _{3/2})4s ² 7d	² [³ / ₂] ^o	1	273 959
4p4d	³ D°?	1	62 552.97	3p ⁵ (² P _{3/2})4s ² 8d	² [³ / ₂] ^o	1	274 652
4p4d	³ P°?	1	62 557.45	3p ⁵ (² P _{3/2})4s ² 9d	² [³ / ₂] ^o	1	275 272
4p4d	³ D°?	1	62 557.45	3p ⁵ (² P _{3/2})4s ² 10d	² [³ / ₂] ^o	1	275 503
4p4d	¹ P°?	1	62 561.72	3p ⁵ (² P _{3/2})4s ² 11d	² [³ / ₂] ^o	1	275 726
4p4d	³ P°?	1	62 561.72	3p ⁵ (² P _{3/2})4s ² 12d	² [³ / ₂] ^o	1	275 898
*****	***	*	*****	3p ⁵ (² P _{3/2})4s ² 13d	² [³ / ₂] ^o	1	276 038
				3p ⁵ (² P _{3/2})4s ² 14d	² [³ / ₂] ^o	1	276 138
				3p ⁵ (² P _{3/2})4s ² 15d	² [³ / ₂] ^o	1	276 214
			1	62 878.40	3p ⁵ (² P _{3/2})4s ² 16d	² [³ / ₂] ^o	276 283
			1	62 886.45	Ca II 3p ⁵ 4s ² ² P _{3/2}	Limit	276 750
*****	***	*	*****	3p ⁵ (² P _{1/2})4s ² 5d	² [³ / ₂] ^o	1	274 040
3p ⁵ 4s ² 3d	³ P°	1	200 096	3p ⁵ (² P _{1/2})4s ² 7d	² [³ / ₂] ^o	1	276 631
3p ⁵ 4s ² 3d	³ D°	1	218 991	3p ⁵ (² P _{1/2})4s ² 8d	² [³ / ₂] ^o	1	277 430
3p ⁵ 4s ² 3d	¹ P°	1	253 310	3p ⁵ (² P _{1/2})4s ² 9d	² [³ / ₂] ^o	1	277 917
3p ⁵ 4s ² 3d	³ P°	1	255 022	3p ⁵ (² P _{1/2})4s ² 10d	² [³ / ₂] ^o	1	278 228
3p ⁵ 4s ² 4d	³ P°	1	267 956	3p ⁵ (² P _{1/2})4s ² 11d	² [³ / ₂] ^o	1	278 469
3p ⁵ 4s ² 4d	³ D°	1	270 640	3p ⁵ (² P _{1/2})4s ² 12d	² [³ / ₂] ^o	1	278 647
3p ⁵ 4s ² 4d	¹ P°	1	270 640	3p ⁵ (² P _{1/2})4s ² 13d	² [³ / ₂] ^o	1	278 796
3p ⁵ 4s ² 4d	³ P°	1	270 640	3p ⁵ (² P _{1/2})4s ² 14d	² [³ / ₂] ^o	1	278 900
3p ⁵ 4s ² 4d	³ D°	1	270 640	3p ⁵ (² P _{1/2})4s ² 15d	² [³ / ₂] ^o	1	278 983
3p ⁵ 4s ² 4d	¹ P°	1	270 640	3p ⁵ (² P _{1/2})4s ² 16d	² [³ / ₂] ^o	1	279 050
3p ⁵ 4s ² 4d	³ D°	1	270 640	3p ⁵ (² P _{1/2})4s ² 17d	² [³ / ₂] ^o	1	279 117

Ca II

 $Z=20$

K I isoelectronic sequence

Ground state: $1s^2 2s^2 2p^6 3s^2 3p^6 4s\ ^2S_{1/2}$ Ionization energy = $95\ 751.87 \pm 0.03$ ($11.87181 \pm .000004$ eV)

The strong resonance doublet of this spectrum, the Fraunhofer H and K lines, is used to measure the velocity of recession of galaxies in the farthest reaches of space and thereby provides us with the dimensions and age of the universe.

The analysis in its present state was nearly completed by Saunders and Russell (1925). The spectrum was re-observed between 3000 and 12 000 Å by Edlén and Risberg (1956), who recalculated the level values from their new grating measurements and the interferometric measurements of Wagman (1937) and found the 5g, 9s, and 10s terms. G. Risberg (1968) added 10d, 8f, 8h, and 10h and revised the $6p\ ^3P^{\circ}_{1/2}$ level.

The 9–10f and 11–16d terms are derived from unpub-

lished observations of Shenstone (1930) in the region 2890 to 3220 Å.

Edlén and Risberg derived the quoted ionization energy from the ng series.

The $3p^5 4s^2\ ^2P^{\circ}$ term was obtained from the limit of the $3p^5 4s^2 ns$ and nd series in Ca I by Mansfield and Newsom (1977).

References

- Edlén, B. and Risberg, P. (1956), Ark. Fys. **10**, 553.
 Mansfield, M. W. D., and Newsom, G. H. (1977), Proc. R. Soc London, Ser. A **357**, 77.
 Risberg, G. (1968), Ark. Fys. **37**, 231.
 Saunders, F. A., and Russell, H. N. (1925), Astrophys. J. **62**, 1.
 Wagman, N. E. (1937), Univ. Pitts. Bull. **34**, 325.

Ca II

Configuration	Term	J	Level (cm ⁻¹)	Configuration	Term	J	Level (cm ⁻¹)
$3p^6(^1S_0)4s$	2S	$\frac{1}{2}$	0.00	$3p^6(^1S_0)7s$	2S	$\frac{1}{2}$	79 448.28
$3p^6(^1S_0)3d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	13 650.19 13 710.88	$3p^6(^1S_0)6d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	80 521.53 80 526.16
$3p^6(^1S_0)4p$	$^2P^{\circ}$	$\frac{1}{2}$ $\frac{3}{2}$	25 191.51 25 414.40	$3p^6(^1S_0)6f$	$^2F^{\circ}$	$\frac{5}{2}, \frac{7}{2}$	83 458.08
$3p^6(^1S_0)5s$	2S	$\frac{1}{2}$	52 166.93	$3p^6(^1S_0)6g$	2G	$\frac{7}{2}, \frac{9}{2}$	83 540.00
$3p^6(^1S_0)4d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	56 839.25 56 858.46	$3p^6(^1S_0)8s$	2S	$\frac{1}{2}$	84 300.89
$3p^6(^1S_0)5p$	$^2P^{\circ}$	$\frac{1}{2}$ $\frac{3}{2}$	60 533.02 60 611.28	$3p^6(^1S_0)7d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	84 928.65 84 936.41
$3p^6(^1S_0)4f$	$^2F^{\circ}$	$\frac{5}{2}, \frac{7}{2}$	68 056.91	$3p^6(^1S_0)7f$	$^2F^{\circ}$	$\frac{5}{2}, \frac{7}{2}$	86 727.06
$3p^6(^1S_0)6s$	2S	$\frac{1}{2}$	70 677.62	$3p^6(^1S_0)7g$	2G	$\frac{7}{2}, \frac{9}{2}$	86 781.14
$3p^6(^1S_0)5d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	72 722.23 72 730.93	$3p^6(^1S_0)9s$	2S	$\frac{1}{2}$	87 267.86
$3p^6(^1S_0)6p$	$^2P^{\circ}$	$\frac{1}{2}$ $\frac{3}{2}$	74 484.92 74 521.75	$3p^6(^1S_0)8d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	87 671.93 87 673.72
$3p^6(^1S_0)5f$	$^2F^{\circ}$	$\frac{5}{2}, \frac{7}{2}$	78 034.39	$3p^6(^1S_0)8f$	$^2F^{\circ}$	$\frac{5}{2}, \frac{7}{2}$	88 847.31
$3p^6(^1S_0)5g$	2G	$\frac{7}{2}, \frac{9}{2}$	78 164.72	$3p^6(^1S_0)8g$	2G	$\frac{7}{2}, \frac{9}{2}$	88 884.54
				$3p^6(^1S_0)8h$	$^2H^{\circ}$	$\frac{9}{2}, \frac{11}{2}$	88 890.64

Ca II—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
$3p^6(^1S_0)10s$	2S	$\frac{1}{2}$	89 214.13	$3p^6(^1S_0)11d$	2D	$\frac{3}{2}, \frac{5}{2}$	91 672.0
$3p^6(^1S_0)9d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	89 487.93 89 489.13	$3p^6(^1S_0)12d$	2D	$\frac{3}{2}, \frac{5}{2}$	92 359.0
$3p^6(^1S_0)9f$	$^2F^\circ$	$\frac{5}{2}, \frac{7}{2}$	90 300.0	$3p^6(^1S_0)13d$	2D	$\frac{3}{2}, \frac{5}{2}$	92 883.0
$3p^6(^1S_0)9g$	2G	$\frac{7}{2}, \frac{9}{2}$	90 326.45	$3p^6(^1S_0)14d$	2D	$\frac{3}{2}, \frac{5}{2}$	93 297.6
$3p^6(^1S_0)10d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	90 753.92 90 754.80	$3p^6(^1S_0)15d$	2D	$\frac{3}{2}, \frac{5}{2}$	93 626.9
$3p^6(^1S_0)10f$	$^2F^\circ$	$\frac{5}{2}, \frac{7}{2}$	91 338.0	$3p^6(^1S_0)16d$	2D	$\frac{3}{2}, \frac{5}{2}$	93 894.5
$3p^6(^1S_0)10h$	$^2H^\circ$	$\frac{9}{2}, \frac{11}{2}$	91 361.00	Ca III (1S_0)	<i>Limit</i>		95 751.87
				$3p^5 4s^2$	$^2P^\circ$	$\frac{3}{2}$ $\frac{1}{2}$	227 444 230 224

Ca III $Z \approx 20$

Ar + isoelectronic sequence

Ground state: $1s^2 2s^2 2p^6 3s^2 3p^6 {}^1S_0$ Ionization energy = $410\ 642 \pm 2\text{ cm}^{-1}$ ($50.9135 \pm .0002\text{ eV}$)

The first observations and analysis of Ca III were carried out by Bowen (1928). He measured the spectrum in the range 400–4000 Å and determined levels in the configurations $3p^5 3d$, $4d$, $4s$, $4p$, and $5s$. A much more extensive analysis has since been completed by Borgström (1968, 1971), who has remeasured the spectrum from 440 to 9640 Å. His designations and level values are given here. He has also determined the ionization energy from the ng series.

Hansen, Persson, and Borgström (1975) found the $3s 3p^5 3d$ configuration, which strongly interacts with the $3p^6 nf$ series, and added two levels to $3p^6 6p$. They reported the percentage compositions given here for the $3s^2 3p^5 5p$, $6p$, $4f$, $5f$, $6f$ and $3s 3p^6 3d$ configurations. The percentages for $3p^4 4p$ in LS coupling are from Borgström (1971).

The two ${}^1P^o$ levels from $3s 3p^6 4p$ and $5p$ were determined by Kastner, Crooker, Behring, and Cohen (1977).

Schmitz, Breuckmann, and Mehlhorn (1976) reported the discovery of $3p^4 4s^2$ but their identification has been shown by Pejcev, Ottley, Rassi and Ross (1978) to be incorrect.

References

- Borgström, A. (1968), *Ark. Fys.* **38**, 243.
 Borgström, A. (1971), *Physica Scripta* **3**, 157.
 Bowen, I. S. (1928), *Phys. Rev.* **31**, 497.
 Hansen, J. E., Persson, W., and Borgström, A. (1975), *Physica Scripta* **11**, 31.
 Kastner, S. O., Crooker, A. M., Behring, W. E., and Cohen, L. (1977), *Phys. Rev.* **16A**, 577.
 Pejcev, Y., Ottley, T. W., Rassi, D., and Ross, K. J. (1978), *J. Phys. B* **11**, 531.
 Schmitz, W., Breuckmann, B., and Mehlhorn, W. (1976), *J. Phys. B* **9**, L493.

Ca III

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Leading percentages
$3s^2 3p^6$	1S	0	0.00	
$3s^2 3p^5 3d$	${}^3P^o$	0	203 373.22	
		1	203 851.95	
		2	204 842.64	
$3s^2 3p^5 3d$	${}^3F^o$	4	212 310.04	
		3	213 379.40	
		2	214 334.06	
$3s^2 3p^5 3d$	${}^1D^o$	2	225 826.22	
$3s^2 3p^5 3d$	${}^3D^o$	3	226 333.56	
		2	227 388.56	
		1	227 432.11	
$3s^2 3p^5 3d$	${}^1F^o$	3	228 413.95	
$3s^2 3p^5 4s$	${}^3P^o$	2	242 547.19	
		1	243 930.44	
		0	245 611.88	
$3s^2 3p^5 4s$	${}^1P^o$	1	247 696.39	
$3s^2 3p^5 4p$	3S	1	272 188.70	98

Ca III—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Leading percentages		
$3s^2 3p^5 4p$	³ D	3	277 022.40	100		
		2	277 380.86	77		
		1	278 621.01	83		
$3s^2 3p^5 3d$	¹ P°	1	279 353.64			
$3s^2 3p^5 4p$	¹ D	2	279 741.80	57		
$3s^2 3p^5 4p$	¹ P	1	281 139.55	57		
$3s^2 3p^5 4p$	³ P	2	281 882.24	63		
		0	282 075.15	98		
		1	282 571.43	64		
$3s^2 3p^5 4p$	¹ S	0	290 934.30	98		
$3s^2 3p^5 4d$	³ P°	0	322 663.28			
		1	323 003.56			
		2	323 655.06			
$3s^2 3p^5 4d$	³ F°	4	324 110.24			
		3	324 660.47			
		2	325 467.67			
$3s^2 3p^5 4d$	¹ F°	3	326 186.32			
$3s^2 3p^5 (^2P_{3/2}^o) 5s$	² [$\frac{3}{2}$]°	2	327 922.87			
		1	328 582.45			
$3s^2 3p^5 4d$	³ D°	1	327 962.11			
		3	328 588.76			
		2	328 606.78			
$3s^2 3p^5 4d$	¹ D°	2	328 090.99			
$3s^2 3p^5 (^2P_{1/2}^o) 5s$	² [$\frac{1}{2}$]°	0	331 048.86			
		1	331 403.20			
$3s^2 3p^5 4d$	¹ P°	1	336 749.11			
$3s^2 3p^5 (^2P_{3/2}^o) 5p$	² [$\frac{1}{2}$]	1	339 198.09	91	9	(² P _{1/2} ^o) ² [$\frac{1}{2}$]
		0	343 110.24	75	25	
$3s^2 3p^5 (^2P_{3/2}^o) 5p$	² [$\frac{5}{2}$]	3	340 580.15	100	8	(² P _{3/2} ^o) ² [$\frac{5}{2}$]
		2	340 748.72	90		
$3s^2 3p^5 (^2P_{3/2}^o) 5p$	² [$\frac{3}{2}$]	1	341 349.19	96	4	(² P _{1/2} ^o) ² [$\frac{3}{2}$]
		2	341 601.46	90	9	(² P _{3/2} ^o) ² [$\frac{3}{2}$]
$3s^2 3p^5 (^2P_{1/2}^o) 5p$	² [$\frac{3}{2}$]	1	343 784.96	95	4	(² P _{1/2} ^o) ² [$\frac{3}{2}$]
		2	344 149.81	97		
$3s^2 3p^5 (^2P_{1/2}^o) 5p$	² [$\frac{1}{2}$]	1	344 257.92	90	9	(² P _{1/2} ^o) ² [$\frac{1}{2}$]
		0	346 692.34	75	25	
$3s^2 3p^5 (^2P_{3/2}^o) 4f$	² [$\frac{5}{2}$]	1	346 732.19	99	8	(² P _{3/2} ^o) ² [$\frac{5}{2}$]
		2	346 896.26	90		

Ca III—Continued

Configuration	Term	J	Level (cm ⁻¹)	Leading percentages		
$3s^2 3p^5 (^2P_{3/2}) 4f$	$^2[9/2]$	5	347 344.37	100		
		4	347 383.51	100		
$3s^2 3p^5 (^2P_{3/2}) 4f$	$^2[5/2]$	3	347 417.05	93	6	$(^2P_{1/2}) ^2[5/2]$
		2	347 758.22	88	9	$(^2P_{3/2}) ^2[5/2]$
$3s^2 3p^5 (^2P_{3/2}) 4f$	$^2[7/2]$	3	348 028.13	97		
		4	348 051.86	97		
$3s^2 3p^5 (^2P_{1/2}) 4f$	$^2[5/2]$	3	350 741.19	93	7	$(^2P_{3/2}) ^2[5/2]$
		2	350 900.04	96	4	$(^2P_{3/2}) ^2[5/2]$
$3s^2 3p^5 (^2P_{1/2}) 4f$	$^2[7/2]$	3	350 779.47	97		
		4	350 805.32	98		
$3s^2 3p^5 (^2P_{3/2}) 5d$	$^2[1/2]^{\circ}$	0	358 940.91			
		1	359 156.83			
$3s^2 3p^5 (^2P_{3/2}) 5d$	$^2[3/2]^{\circ}$	2	359 520.60			
		1	361 794.08			
$3s^2 3p^5 (^2P_{3/2}) 5d$	$^2[7/2]^{\circ}$	4	359 543.08			
		3	359 800.57			
$3s^2 3p^5 (^2P_{3/2}) 5d$	$^2[5/2]^{\circ}$	2	360 207.16			
		3	360 388.68			
$3s^2 3p^5 (^2P_{3/2}) 6s$	$^2[3/2]^{\circ}$	2	361 154.07			
		1	361 353.31			
$3s^2 3p^5 (^2P_{1/2}) 5d$	$^2[5/2]^{\circ}$	2	362 893.47			
		3	363 118.75			
$3s^2 3p^5 (^2P_{1/2}) 5d$	$^2[3/2]^{\circ}$	2	363 021.27			
		1	365 363.88			
$3s^2 3p^5 (^2P_{1/2}) 6s$	$^2[1/2]^{\circ}$	0	364 229.89			
		1	364 343.85			
$3s 3p^6 3d$	3D	1	366 778.85	65	31	$3p^5 (^2P_{3/2}) 5f ^2[5/2]$
		2	366 926.33	55	17	$3p^5 (^2P_{3/2}) 6p ^2[5/2]$
		3	367 472.91	53	24	$3p^5 (^2P_{3/2}) 6p ^2[5/2]$
$3s^2 3p^5 (^2P_{3/2}) 6p$	$^2[5/2]$	3	367 026.65	75	18	$3s 3p^6 3d ^3D$
$3s^2 3p^5 (^2P_{3/2}) 5f$	$^2[9/2]$	5	370 141.77	100		
		4	370 172.43	100		
$3s^2 3p^5 (^2P_{3/2}) 5f$	$^2[5/2]$	2	370 304.54?	58	34	$3p^5 (^2P_{3/2}) 5f ^2[5/2]$
		3	371 059.95?	80	17	$3s 3p^6 3d ^3D$
$3s^2 3p^5 (^2P_{1/2}) 6p$	$^2[3/2]$	2	370 454.04	93	6	$3p^5 (^2P_{3/2}) 5f ^2[5/2]$
$3s^2 3p^5 (^2P_{3/2}) 5f$	$^2[7/2]$	3	370 515.86	99		
		4	370 533.07	99		
$3s^2 3p^5 (^2P_{3/2}) 5g$	$^2[5/2]^{\circ}$	2	370 901.88			
		3	370 903.12			

Ca III—Continued

Configuration	Term	J	Level (cm ⁻¹)	Leading percentages
$3s^2 3p^5(^2P_{3/2}) 5g$	$^2[1\frac{1}{2}]^{\circ}$	6	370 957.52	
		5	370 957.65	
$3s^2 3p^5(^2P_{3/2}) 5g$	$^2[7/2]^{\circ}$	3	371 061.30	
		4	371 061.37	
$3s^2 3p^5(^2P_{3/2}) 5g$	$^2[9/2]^{\circ}$	4	371 120.69	
		5	371 121.06	
$3s^2 3p^5(^2P_{3/2}) 5f$	$^2[3/2]$	2	371 274.48?	42
		1	371 447.58?	67
$3s^2 3p^5(^2P_{1/2}) 5f$	$^2[7/2]$	3	373 401.87	99
		4	373 425.51	99
$3s^2 3p^5(^2P_{1/2}) 5f$	$^2[5/2]$	2	373 626.85	92
		3	373 880.37	91
$3s^2 3p^5(^2P_{1/2}) 5g$	$^2[9/2]^{\circ}$	4	374 138.90	
		5	374 139.31	
$3s^2 3p^5(^2P_{1/2}) 5g$	$^2[7/2]^{\circ}$	4	374 143.44	
		3	374 143.84	
$3s^2 3p^5(^2P_{3/2}) 6d$	$^2[7/2]^{\circ}$	4	376 808.60	
		3	376 883.36	
$3s 3p^6 3d$	1D	2	377 168.1	93
$3s^2 3p^5(^2P_{3/2}) 6d$	$^2[5/2]^{\circ}$	2	380 152.21	
		3	380 230.50	
$3s^2 3p^5(^2P_{3/2}) 6g$	$^2[9/2]^{\circ}$	5	382 190.20	
		4	383 189.82	
$3s^2 3p^5(^2P_{3/2}) 6f$	$^2[9/2]$	5	382 565.1	100
		4	382 587.9	100
$3s^2 3p^5(^2P_{3/2}) 6f$	$^2[7/2]$	3	382 784.7	100
		4	382 798.5	99
$3s^2 3p^5(^2P_{3/2}) 6f$	$^2[5/2]$	3	382 791.5	99
		2	382 852.3	79
$3s^2 3p^5(^2P_{3/2}) 6g$	$^2[5/2]^{\circ}$	2	383 061.33	
		3	383 063.78	
$3s^2 3p^5(^2P_{3/2}) 6g$	$^2[11/2]^{\circ}$	6	383 094.79	
		5	383 095.10	
$3s^2 3p^5(^2P_{3/2}) 6g$	$^2[7/2]^{\circ}$	4	383 156.51	
		3	383 157.11	
$3s^2 3p^5(^2P_{1/2}) 6f$	$^2[7/2]$	3	385 757.6	100
		4	385 775.5	100
$3s^2 3p^5(^2P_{1/2}) 6f$	$^2[5/2]$	3	385 867.2	99
		2	385 906.9	99

Ca III—Continued

Configuration	Term	J	Level (cm ⁻¹)	Leading percentages
$3s^2 3p^5 (^2P_{1/2}^o) 6g$	$^2[\frac{9}{2}]^o$	4 5	386 248.39 386 248.68	
$3s^2 3p^5 (^2P_{1/2}^o) 6g$	$^2[\frac{7}{2}]^o$	4 3	386 249.78 386 250.31	
$3s^2 3p^5 (^2P_{3/2}^o) 7f$	$^2[\frac{9}{2}]$	5	390 054.0	
$3s^2 3p^5 (^2P_{3/2}^o) 7f$	$^2[\frac{7}{2}]$	4	390 207.6	
$3s^2 3p^5 (^2P_{3/2}^o) 7g$	$^2[\frac{5}{2}]^o$	3	390 392.62	
$3s^2 3p^5 (^2P_{3/2}^o) 7g$	$^2[\frac{11}{2}]^o$	6 5	390 411.58 390 411.99	
$3s^2 3p^5 (^2P_{3/2}^o) 7g$	$^2[\frac{7}{2}]^o$	3	390 451.97	
$3s^2 3p^5 (^2P_{3/2}^o) 7g$	$^2[\frac{9}{2}]^o$	4 5	390 471.79 390 472.32	
$3s^2 3p^5 (^2P_{1/2}^o) 7f$	$^2[\frac{7}{2}]$	4	393 224.9	
$3s^2 3p^5 (^2P_{1/2}^o) 7g$	$^2[\frac{9}{2}]^o$	5 4	393 551.86 393 561.59	
$3s^2 3p^5 (^2P_{1/2}^o) 7g$	$^2[\frac{7}{2}]^o$	4 3	393 552.31 393 552.81	
Ca IV ($^2P_{3/2}^o$)	Limit		410 642	
$3s3p^6 4p$	$^1P^o$	1	431 100	
$3s3p^6 5p$	$^1P^o$	1	492 850	

Ca IV

Z=20

Cl I isoelectronic sequence

Ground state: $1s^2 2s^2 2p^6 3s^2 3p^5 \ ^2P^o_{3/2}$ Ionization energy = 542 600 cm⁻¹ (67.27 eV)

The initial work on the analysis was by Bowen (1928), who identified the ground term, the $3s3p^6 \ ^2S$ level, and the $3p^4(^3P)4s \ ^2P_{3/2}$ level. Kruger and Phillips (1937) also identified the $3p^44s$ terms from their observations below 350 Å.

Levels of $3p^4 3d \ ^4D$, 2D and 2F terms and the $3p^4 4p$ and $5s$ levels were reported by Tsien (1939), who worked with the line list of Ekefors (1931). Svensson and Ekberg (1968), also using Ekefors' list, established the $3p^4(^3P)3d \ ^2P$ term.

A new analysis based on a new set of observations was provided by Smitt (1978), whose work is still in progress. He estimates an accuracy of ± 2 cm⁻¹ for the levels relative to the ground term and less than ± 1 cm⁻¹ for the ground term splitting. The new analysis provides three

times the number of previously known levels as well as some changes of designation and *J*-values. All levels compiled here are from this work except for the $3p^4 5s \ ^2D$ term retained from Tsien.

The ionization energy was obtained by extrapolation by Lotz (1967).

References

- Bowen, I. S. (1928), Phys. Rev. **31**, 497.
 Ekefors, E. (1931), Z. Phys. **71**, 53.
 Kruger, P. G., and Phillips, L. W. (1937), Phys. Rev. **51**, 1087.
 Lotz, W. (1967), J. Opt. Soc. Am. **57**, 873.
 Smitt, R. (1978), Private communication.
 Svensson, L. A., and Ekberg, J. O. (1968), Ark. Fys. **37**, 65.
 Tsien, W.-Z. (1939), Chinese J. Phys. **3**, 117.

Ca IV

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
$3s^2 3p^5$	$^2P^o$	$\frac{3}{2}$ $\frac{1}{2}$	0.0 3 118.2	$3s^2 3p^4(^1D)3d$	2F	$\frac{5}{2}$ $\frac{7}{2}$	252 286 252 903
$3s3p^6$	2S	$\frac{1}{2}$	152 439.6	$3s^2 3p^4(^1S)3d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	269 974 270 532
$3s^2 3p^4(^3P)3d$	4D	$\frac{7}{2}$ $\frac{5}{2}$ $\frac{3}{2}$ $\frac{1}{2}$	201 505 201 747 202 031 202 254	$3s^2 3p^4(^3P)4s$	4P	$\frac{5}{2}$ $\frac{3}{2}$ $\frac{1}{2}$	291 456 292 864 294 292
$3s^2 3p^4(^3P)3d$	4F	$\frac{9}{2}$ $\frac{7}{2}$ $\frac{5}{2}$ $\frac{3}{2}$	218 383 219 467 220 240 220 741	$3s^2 3p^4(^1D)3d$	2S	$\frac{1}{2}$	293 009
$3s^2 3p^4(^1D)3d$	2P	$\frac{1}{2}$ $\frac{3}{2}$	219 991 221 945	$3s^2 3p^4(^3P)3d$	2P	$\frac{3}{2}$ $\frac{1}{2}$	293 872 295 133
$3s^2 3p^4(^3P)3d$	4P	$\frac{1}{2}$ $\frac{3}{2}$ $\frac{5}{2}$	227 214 227 825 228 694	$3s^2 3p^4(^3P)4s$	2P	$\frac{5}{2}$ $\frac{1}{2}$	301 218 303 850
$3s^2 3p^4(^1D)3d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	228 436 230 119	$3s^2 3p^4(^1D)4s$	2D	$\frac{5}{2}$ $\frac{3}{2}$	312 517 312 650
$3s^2 3p^4(^3P)3d$	2F	$\frac{7}{2}$ $\frac{5}{2}$	231 288 233 851	$3s^2 3p^4(^3P)4p$	$^4P^o$	$\frac{5}{2}$ $\frac{3}{2}$ $\frac{1}{2}$	330 693 331 173 331 969
$3s^2 3p^4(^1D)3d$	2G	$\frac{9}{2}$ $\frac{7}{2}$	234 498 234 642				

Ca iv—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
$3s^2 3p^4(^3P)4p$	$^4D^\circ$	$\frac{7}{2}$ $\frac{5}{2}$ $\frac{3}{2}$ $\frac{1}{2}$	335 122 335 901 336 958 337 452	$3s^2 3p^4(^1D)4p$	$^2P^\circ$	$\frac{3}{2}$ $\frac{1}{2}$	359 575 361 170
$3s^2 3p^4(^1S)4s$	2S	$\frac{1}{2}$	337 214	$3s3p^5(^3P^\circ)3d$	$^4F^\circ$	$\frac{9}{2}$ $\frac{7}{2}$ $\frac{5}{2}$ $\frac{3}{2}$	367 174 367 971 368 660 369 197
$3s^2 3p^4(^3P)4p$	$^2D^\circ$	$\frac{5}{2}$ $\frac{3}{2}$	338 250 340 377	$3s^2 3p^4(^1S)4p$	$^2P^\circ$	$\frac{1}{2}$ $\frac{3}{2}$	379 774 380 043
$3s^2 3p^4(^3P)4p$	$^2P^\circ$	$\frac{1}{2}$ $\frac{3}{2}$	338 300 338 959	$3s3p^5(^3P^\circ)3d$	$^4D^\circ$	$\frac{7}{2}$ $\frac{5}{2}$ $\frac{3}{2}$ $\frac{1}{2}$	382 641 382 974 383 125 383 176
$3s^2 3p^4(^3P)4p$	$^2S^\circ$	$\frac{1}{2}$	342 567				
$3s^2 3p^4(^3P)4p$	$^4S^\circ$	$\frac{3}{2}$	342 915	$3s^2 3p^4(^1D)5s$	2D	$\frac{5}{2}$ $\frac{3}{2}$	399 755 400 956
$3s^2 3p^4(^1D)4p$	$^2F^\circ$	$\frac{5}{2}$ $\frac{3}{2}$	352 154 352 616	Ca v (3P_2)	<i>Limit</i>		542 600
$3s^2 3p^4(^1D)4p$	$^2D^\circ$	$\frac{3}{2}$ $\frac{5}{2}$	357 941 358 306				

Ca v

Z=20

S I isoelectronic sequence

Ground state: $1s^2 2s^2 2p^6 3s^2 3p^4$ 3P_2 Ionization energy = 681 600 cm⁻¹ (84.50 eV)

Spectra of calcium from 1035–135 Å were obtained by Ekefors (1931) and were supplemented by longer exposures below 600 Å by Bowen (1934). From the combined line-lists Bowen derived levels of the configurations $3s^2 3p^4$, $3s 3p^5$, $3s^2 3p^3 3d$, $3s^2 3p^4 4s$, and $3s^2 3p^3 5s$. By means of isoelectronic comparisons Svensson and Ekberg (1968) revised and extended the $3p^3 3d$ configuration.

New measurements of the transition array $3s^2 3p^4$ – $3s 3p^5$ by Smitt, Svensson, and Outred (1976) led to improved values for the levels of these configurations.

The ionization energy is an extrapolated value by Lotz (1967).

References

- Bowen, I. S. (1934), Phys. Rev. **46**, 791.
 Ekefors, E. (1931), Z. Phys. **71**, 53.
 Lotz, W. (1967), J. Opt. Soc. Am. **57**, 873.
 Smitt, R., Svensson, L. A., and Outred, M. (1976), Physica Scripta **13**, 293.
 Svensson, L. A., and Ekberg, J. O. (1968), Ark. Fys. **37**, 65.

Ca v

Configuration	Term	J	Level (cm ⁻¹)	Configuration	Term	J	Level (cm ⁻¹)	
$3s^2 3p^4$	3P	2	0.0	$3s^2 3p^3 ({}^2P^\circ) 3d$	${}^1F^\circ$	3	329 229	
		1	2 404.7		${}^3S^\circ$	1	350 914	
		0	3 275.6		${}^1P^\circ$	1	353 220	
$3s^2 3p^4$	1D	2	18 830.3	$3s^2 3p^3 ({}^2D^\circ) 4s$	${}^3D^\circ$	1	369 590	
		0	43 836.5			2	369 696	
						3	369 959	
$3s 3p^5$	${}^3P^\circ$	2	154 670.8	$3s^2 3p^3 ({}^2D^\circ) 4s$	${}^1D^\circ$	2	374 728	
		1	156 760.2		${}^3P^\circ$	0	387 039	
		0	157 900.5			1	387 226	
$3s 3p^5$	${}^1P^\circ$	1	197 844.5			2	387 652	
	$3s^2 3p^3 ({}^2D^\circ) 3d$	${}^1D^\circ$	2	254 124	$3s^2 3p^3 ({}^2P^\circ) 4s$	${}^1P^\circ$	1	393 283
				${}^3S^\circ$	1	501 127		
$3s^2 3p^3 ({}^2D^\circ) 3d$	${}^1F^\circ$	3	283 955	${}^3D^\circ$	1	524 651		
	$3s^2 3p^3 ({}^2D^\circ) 3d$	${}^3S^\circ$	1	293 785		2	524 770	
					3	525 053		
$3s^2 3p^3 ({}^2P^\circ) 3d$	${}^3P^\circ$	2	298 214	$3s^2 3p^3 ({}^2D^\circ) 5s$	${}^1D^\circ$	2	526 523	
		1	299 534		${}^3P^\circ$	1	542 249	
		0	300 594			2	542 650	
$3s^2 3p^3 ({}^2D^\circ) 3d$	${}^1P^\circ$	1	302 184	$3s^2 3p^3 ({}^2P^\circ) 5s$	${}^1P^\circ$	1	544 143	
	$3s^2 3p^3 ({}^2P^\circ) 3d$	${}^3D^\circ$	3	308 188				
		2	309 831					
$3s^2 3p^3 ({}^2P^\circ) 3d$		1	310 943					
	$3s^2 3p^3 ({}^2P^\circ) 3d$	${}^1D^\circ$	2	318 741	Ca VI (${}^4S_{3/2}$)	<i>Limit</i>	681 600	

Ca vi

 $Z=20$

P 1 isoelectronic sequence

Ground state: $1s^2 2s^2 2p^6 3s^2 3p^3 \ ^4S^{\circ}_{3/2}$ Ionization energy = 877 400 cm^{-1} (108.78 eV)

The present compilation is obtained from the work of Ekberg and Svensson (1970) and Smitt, Svensson, and Outred (1976). The level values for the $3s^2 3p^3$ and $3s 3p^4$ configurations are taken from the latter paper. They have an uncertainty of about $\pm 2 \text{ cm}^{-1}$. We have combined these values with the measurements and classifications given by Ekberg and Svensson to derive new level values for the $3p^2 3d$ and $4s$ configurations. The uncertainty of these upper levels is about $\pm 10 \text{ cm}^{-1}$. Since no inter-system transitions have been observed, all of the doublets have an added systematic error x , relative to the ground term $^4S^{\circ}$. The value of x depends on the accuracy of cal-

culations by Smitt, Svensson and Outred and is expected to be less than $\pm 20 \text{ cm}^{-1}$.

Most of the wavelengths used by Ekberg and Svensson are taken from Bowen (1934) or Ekefors (1931).

The ionization energy is from an extrapolation by Lotz (1967).

References

- Bowen, I. S. (1934), Phys. Rev. **46**, 791.
 Ekberg, J. O., and Svensson, L. A. (1970), Physica Scripta **2**, 283.
 Ekefors, E. (1931), Z. Phys. **71**, 53.
 Lotz, W. (1967), J. Opt. Soc. Am. **57**, 873.
 Smitt, R., Svensson, L. A., and Outred, M. (1976), Physica Scripta **13**, 293.

Ca vi

Configuration	Term	J	Level (cm^{-1})	Configuration	Term	J	Level (cm^{-1})
$3s^2 3p^3$	$^4S^{\circ}$	$\frac{3}{2}$	0.0	$3s^2 3p^2(^1D)3d$	2P	$\frac{1}{2}$ $\frac{3}{2}$	331 968 + x 333 324 + x
$3s^2 3p^3$	$^2D^{\circ}$	$\frac{3}{2}$ $\frac{5}{2}$	26 835.1 + x 27 246.6 + x	$3s^2 3p^2(^1D)3d$	2F	$\frac{5}{2}$ $\frac{7}{2}$	336 219 + x 336 631 + x
$3s^2 3p^3$	$^2P^{\circ}$	$\frac{1}{2}$ $\frac{3}{2}$	44 586.7 + x 45 142.7 + x	$3s^2 3p^2(^1S)3d$	2D	$\frac{5}{2}$ $\frac{3}{2}$	348 819 + x 349 645 + x
$3s 3p^4$	4P	$\frac{5}{2}$ $\frac{3}{2}$ $\frac{1}{2}$	155 786.5 157 767.5 158 830.5	$3s^2 3p^2(^3P)4s$	4P	$\frac{1}{2}$ $\frac{3}{2}$ $\frac{5}{2}$	433 849 435 286 437 392
$3s 3p^4$	2D	$\frac{3}{2}$ $\frac{5}{2}$	193 247.0 + x 193 444.8 + x	$3s^2 3p^2(^3P)4s$	2P	$\frac{1}{2}$ $\frac{3}{2}$	442 256 + x 444 724 + x
$3s 3p^4$	2P	$\frac{3}{2}$ $\frac{1}{2}$	222 749.7 + x 224 773.3 + x	$3s^2 3p^2(^1D)4s$	2D	$\frac{5}{2}$ $\frac{3}{2}$	457 294 + x 457 358 + x
$3s 3p^4$	2S	$\frac{1}{2}$	233 712.8 + x	$3s^2 3p^2(^1S)4s$	2S	$\frac{1}{2}$	483 882 + x
$3s^2 3p^2(^3P)3d$	2P	$\frac{3}{2}$ $\frac{1}{2}$	294 630 + x 297 088 + x	Ca VII (3P_0)	<i>Limit</i>		877 400
$3s^2 3p^2(^1D)3d$	2D	$\frac{5}{2}$ $\frac{3}{2}$	320 919 + x 321 411 + x				

Ca VII

 $Z=20$

Si I isoelectronic sequence

Ground state: $1s^2 2s^2 2p^6 3s^2 3p^2 {}^3P_0$ Ionization energy = $1\ 026\ 000\ \text{cm}^{-1}$ (127.2 eV)

The level values for $3s^2 3p^2$ and $3s3p^3$ are taken from Smitt, Svensson, and Outred (1976). Ekberg and Svensson (1970) revised and extended the interpretation of the $3p^2-3p3d$, $4s$ arrays published by several earlier workers. We have combined these new classifications with the $3p^2$ levels given by Smitt et al. Ekberg and Svensson have

obtained the value for the ionization energy by extrapolation.

References

- Ekberg, J. O., and Svensson, L. A. (1970), *Physica Scripta* **2**, 283.
 Smitt, R., Svensson, L. A., and Outred, M. (1976), *Physica Scripta* **13**, 293.

Ca VII

Configuration	Term	J	Level (cm^{-1})	Configuration	Term	J	Level (cm^{-1})
$3s^2 3p^2$	3P	0	0.0	$3s3p^3$	${}^1P^\circ$	1	252 489.9
		1	1 624.9			2	286 224
		2	4 071.4			1	288 160
$3s^2 3p^2$	1D	2	21 864.0	$3s^2 3p3d$	${}^3P^\circ$	0	289 004
$3s^2 3p^2$	1S	0	48 981.4	$3s^2 3p3d$	${}^3D^\circ$	1	295 138
$3s3p^3$	${}^3D^\circ$	1	160 157.5	2	295 772		
		2	160 220.3	3	296 132		
		3	160 529.2	$3s^2 3p3d$	${}^1F^\circ$	3	324 885
$3s3p^3$	${}^3P^\circ$	0	185 356.6	$3s^2 3p3d$	${}^1P^\circ$	1	333 501
		1	185 392.9	$3s^2 3p4s$	${}^3P^\circ$	0	490 059
		2	185 412.2	$3s^2 3p4s$	${}^3P^\circ$	1	490 919
$3s3p^3$	${}^1D^\circ$	2	203 616.1	$3s^2 3p4s$	${}^3P^\circ$	2	494 262
$3s3p^3$	${}^3S^\circ$	1	245 240.5	$3s^2 3p4s$	${}^1P^\circ$	1	498 683
				Ca VIII (${}^2P_{1/2}^\circ$)	<i>Limit</i>		1 026 000

Ca VIII

 $Z = 20$

Al I isoelectronic sequence

Ground state: $1s^2 2s^2 2p^6 3s^2 3p^2 P^{\circ}_{1/2}$ Ionization energy = $1\ 187\ 600 \pm 1000\ \text{cm}^{-1}$ ($147.24 \pm 0.1\ \text{eV}$)

The doublet terms of $3s^2 3p$ and $3s 3p^2$ are from Smitt, Svensson, and Outred (1976). The uncertainty of their measurements is about $2\ \text{cm}^{-1}$.

The remaining terms are derived from the measurements and classifications of Ekberg and Svensson (1970). They obtained the position of the quartets by extrapolation.

The ionization energy was determined by Ekberg and Svensson from the $n/\ell F^{\circ}$ series.

References

- Ekberg, J. O. and Svensson, L. A. (1970), *Physica Scripta* **2**, 283.
 Smitt, R., Svensson, L. A., and Outred, M. (1976), *Physica Scripta* **13**, 293.

Ca VIII

Configuration	Term	<i>J</i>	Level (cm^{-1})	Configuration	Term	<i>J</i>	Level (cm^{-1})
$3s^2 3p$	$^2P^{\circ}$	$\frac{1}{2}$ $\frac{3}{2}$	0.0 4 308.3	$3s^2 4s$	2S	$\frac{1}{2}$	547 322
$3s 3p^2$	4P	$\frac{1}{2}$ $\frac{3}{2}$ $\frac{5}{2}$	129 100+x 130 678+x 133 042+x	$3s 3p 4s$	$^4P^{\circ}$	$\frac{1}{2}$ $\frac{3}{2}$ $\frac{5}{2}$	688 747+x 690 128+x 692 833+x
$3s 3p^2$	2D	$\frac{3}{2}$ $\frac{5}{2}$	171 572.2 171 830.7	$3s^2 4d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	698 232 698 420
$3s 3p^2$	2S	$\frac{1}{2}$	216 584.9	$3s^2 4f$	$^2F^{\circ}$	$\frac{5}{2}$ $\frac{7}{2}$	743 288 743 330
$3s 3p^2$	2P	$\frac{1}{2}$ $\frac{3}{2}$	231 016.3 233 592.8	$3s^2 5d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	885 693 885 750
$3s^2 3d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	282 356 282 577	$3s^2 5f$	$^2F^{\circ}$	$\frac{5}{2}$ $\frac{7}{2}$	905 052 905 087
$3p^3$	$^4S^{\circ}$	$\frac{3}{2}$	345 274+x	$3s^2 6d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	979 749 980 089
$3s 3p 3d$	$^4P^{\circ}$	$\frac{5}{2}$ $\frac{3}{2}$	408 227+x 409 291+x	$3s^2 6f$	$^2F^{\circ}$	$\frac{5}{2}$ $\frac{7}{2}$	991 023 991 028
$3s 3p 3d$	$^4D^{\circ}$	$\frac{1}{2}$ $\frac{3}{2}$ $\frac{5}{2}$ $\frac{7}{2}$	411 816+x 412 388+x 412 772+x 412 881+x	$3s^2 7f$	$^2F^{\circ}$	$\frac{7}{2}$ $\frac{5}{2}$	1 043 207 1 043 275
				Ca IX (1S_0)	Limit		1 187 600

Ca ix

 $Z=20$

Mg I isoelectronic sequence

Ground state: $1s^2 2s^2 2p^6 3s^2$ 1S_0 Ionization energy = $1\ 519\ 000 \pm 1000$ cm $^{-1}$ (188.3 ± 0.1 eV)

Most of the levels for this spectrum are taken from Ekberg (1971). The identification of the two combinations with $3s5f$ ${}^1F^o_3$ has been revised by Edlén and Bodén (1976). Ekberg obtained his intersystem connection with two faint lines, but regards this as tentative.

The $3p^2$ 1S term and the $3p3d$ configuration are from Fawcett (1970). The $3p4f$ configuration is from Fawcett (1976).

The ionization energy was calculated by Ekberg from the first three members of the $3snf$ ${}^3F^o$ series ($n=4, 5$, and 6).

References

- Edlén, B., and Bodén, E. (1976), *Physica Scripta* **14**, 31.
 Ekberg, J. O. (1971), *Physica Scripta* **4**, 101.
 Fawcett, B. C. (1970), *J. Phys.* **B3**, 1732.
 Fawcett, B. C. (1976), *J. Opt. Soc. Am.* **66**, 632.

Ca ix

Configuration	Term	J	Level (cm $^{-1}$)	Configuration	Term	J	Level (cm $^{-1}$)
$3s^2$	1S	0	0	$3s4p$	${}^1P^o$	1	832 314
$3s3p$	${}^3P^o$	0	141 612	$3s4d$	3D	1	915 636
		1	143 111			2	915 750
		2	146 348			3	915 964
$3s3p$	${}^1P^o$	1	214 482	$3s4d$	1D	2	921 921
$3p^2$	1D	2	336 245	$3p4s$	${}^3P^o$	0	939 907
$3p^2$	3P	0	338 399			1	941 094
		1	340 308			2	944 814
		2	343 908	$3s4f$	${}^3F^o$	2-4	953 030
$3p^2$	1S	0	398 900	$3s4f$	${}^1F^o$	3	963 050
$3s3d$	3D	1	410 514	$3p4p$	3D	2	1 003 670
		2	410 627			3	1 007 010
		3	410 841				
$3s3d$	1D	2	467 631	$3p4p$	3P	0	1 009 330
$3p3d$	${}^3F^o$	2	562 150			1	1 010 470
		3	564 160			2	1 012 820
		4	566 630	$3p4p$	3S	1	1 014 060
$3p3d$	${}^1D^o$	2	571 900	$3s5s$	3S	1	1 067 240
$3p3d$	${}^3D^o$	2	599 640	$3s5s$	1S	0	1 076 110
		3	601 140	$3s5p$	${}^1P^o$	1	1 097 570
$3p3d$	${}^1P^o$	1	618 520	$3p4f$	3G	5	1 125 440
$3s4s$	3S	1	758 974	$3p4f$	3F	3	1 125 620
$3s4s$	1S	0	774 480			4	1 128 930

Ca ix—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
3s5d	¹ D	2	1 139 810	3s6d	³ D	1	1 258 080
3s5d	³ D	1	1 143 670			2	1 258 260
		3	1 144 290			3	1 258 450
3s5f	³ F°	2-4	1 158 110	3s6d	¹ D	2	1 260 390
3s5f	¹ F°	3	1 162 610	3s6f	³ F°	2-4	1 269 050
3s6s	³ S	1	1 218 220	3s7p	¹ P°	1	1 315 300
3s6p	¹ P°	1	1 235 830	3s7d	³ D	1-3	1 329 760
				3s8d	³ D	1-3	1 374 830
				Ca x (² S _{1/2})	<i>Limit</i>		1 519 000

Ca x

Z=20

Na I isoelectronic sequence

Ground state: $1s^2 2s^2 2p^6 3s\ ^2S_{1/2}$ Ionization energy = $1\ 704\ 047 \pm 3\text{ cm}^{-1}$ ($211.277 \pm .001\text{ eV}$)

The recent publications by Edlén and Bodén (1976), by Fawcett (1976), and by Cohen and Behring (1976) provide considerable extensions of the early work of Kruger and Phillips (1939) on this spectrum. We have quoted level values from Edlén and Bodén and added the higher nf series members ($n=10-11$) from Cohen and Behring. The measurements of the high members of the nf series by Fawcett do not agree well with the Ritz formulae given by Edlén and Bodén.

The ionization energy is derived by Edlén (1978) from a polarization formula applied to the nf series.

References

- Cohen, L., and Behring, W. E. (1976), J. Opt. Soc. Am. **66**, 899.
 Edlén, B. (1978), Physica Scripta **17**, 565.
 Edlén, B., and Bodén, E. (1976), Physica Scripta **14**, 31.
 Fawcett, B. C. (1976), J. Opt. Soc. Am. **66**, 632.
 Kruger, P. G., and Phillips, L. W. (1939), Phys. Rev. **55**, 352.

Ca x

Configuration	Term	J	Level (cm $^{-1}$)	Configuration	Term	J	Level (cm $^{-1}$)
$2p^6(^1S)3s$	2S	$\frac{1}{2}$	0	$2p^6(^1S)6d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	1 389 840 1 389 870
$2p^6(^1S)3p$	$^2P^\circ$	$\frac{1}{2}$ $\frac{3}{2}$	174 213 179 287	$2p^6(^1S)6f$	$^2F^\circ$	$\frac{5}{2}$ $\frac{7}{2}$	1 398 330 1 398 440
$2p^6(^1S)3d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	417 112 417 522	$2p^6(^1S)7s$	2S	$\frac{1}{2}$	1 448 710
$2p^6(^1S)4s$	2S	$\frac{1}{2}$	832 790	$2p^6(^1S)7p$	$^2P^\circ$	$\frac{1}{2}, \frac{3}{2}$	1 459 920
$2p^6(^1S)4p$	$^2P^\circ$	$\frac{1}{2}$ $\frac{3}{2}$	899 290 901 200	$2p^6(^1S)7d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	1 474 040 1 474 090
$2p^6(^1S)4d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	987 300 987 490	$2p^6(^1S)7f$	$^2F^\circ$	$\frac{5}{2}$ $\frac{7}{2}$	1 479 470 1 479 540
$2p^6(^1S)4f$	$^2F^\circ$	$\frac{5}{2}$ $\frac{7}{2}$	1 016 100 1 016 150	$2p^6(^1S)8s$	2S	$\frac{1}{2}$	1 511 780
$2p^6(^1S)5s$	2S	$\frac{1}{2}$	1 174 710	$2p^6(^1S)8p$	$^2P^\circ$	$\frac{1}{2}, \frac{3}{2}$	1 519 200
$2p^6(^1S)5p$	$^2P^\circ$	$\frac{1}{2}$ $\frac{3}{2}$	1 206 850 1 207 760	$2p^6(^1S)8d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	1 528 490 1 528 510
$2p^6(^1S)5d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	1 248 920 1 249 030	$2p^6(^1S)8f$	$^2F^\circ$	$\frac{5}{2}$ $\frac{7}{2}$	1 532 290 1 532 390
$2p^6(^1S)5f$	$^2F^\circ$	$\frac{5}{2}$ $\frac{7}{2}$	1 263 690 1 263 720	$2p^6(^1S)9p$	$^2P^\circ$	$\frac{1}{2}, \frac{3}{2}$	1 559 260
$2p^6(^1S)6s$	2S	$\frac{1}{2}$	1 348 380	$2p^6(^1S)9d$	2D	$\frac{3}{2}, \frac{5}{2}$	1 565 730
$2p^6(^1S)6p$	$^2P^\circ$	$\frac{1}{2}$ $\frac{3}{2}$	1 366 360 1 366 890	$2p^6(^1S)9f$	$^2F^\circ$	$\frac{5}{2}$ $\frac{7}{2}$	1 568 390 1 568 420

Ca x—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
$2p^6(^1S)10d$	2D	$\frac{5}{2}$	1 592 260	$2p^6(^1S)11f$	$^2F^\circ$	$\frac{7}{2}$	1 613 480
$2p^6(^1S)10f$	$^2P^\circ$	$\frac{7}{2}$ $\frac{5}{2}$	1 594 230 1 594 330	Ca XI (1S_0)	<i>Limit</i>	$\frac{5}{2}$	1 613 600

Ca xi

 $Z=20$

Ne I isoelectronic sequence

Ground state: $1s^2 2s^2 2p^6 \ ^1S_0$ Ionization energy = 4 774 000 cm⁻¹ (591.9 eV)

Only resonance lines between 25 and 36 Å are classified by this rare-gas-type system of energy levels. Edlén and Tyrén (1936) identified 11 transitions and extrapolated to obtain an ionization potential which agrees well with the present value.

We use jj -coupling designations for the $2p^5 ns$ levels and jl -coupling designations for the $2p^5 nd$ levels.

Kastner, Behring, and Cohen (1975) identified transitions between $2p^5 3p$ and $2p^5 4d$, but there is no connection with the levels given here.

We derived the ionization energy from the $2s^2 2p^5 (^2P_{3/2}) nd \ ^2[3/2]^o$ series for $n=3$ and 4 with the change in quantum defect Δn^* taken from Ti XIII. Our value agrees exactly with Lotz's (1967) value.

References

- Edlén, B., and Tyrén, F. (1936), Z. Phys. **101**, 206.
 Kastner, S. O., Behring, W. E., and Cohen, L. (1975), Astrophys. J. **199**, 777.
 Lotz, W. J. (1967), J. Opt. Soc. Am. **57**, 973.

Ca xi

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
$2s^2 2p^6$	1S	0	0	$2s2p^6 3p$	$^3P^o$	1	3 692 900
$2s^2 2p^5 (^2P_{3/2}) 3s$	$(^3/2, ^1/2)^o$	1	2 810 880	$2s2p^6 3p$	$^1P^o$	1	3 708 900
$2s^2 2p^5 (^2P_{1/2}) 3s$	$(^1/2, ^1/2)^o$	1	2 839 940	$2s^2 2p^5 (^2P_{3/2}) 4s$	$(^3/2, ^1/2)^o$	1	3 753 900
$2s^2 2p^5 (^2P_{3/2}) 3d$	$^2[1/2]^o$	1	3 199 300	$2s^2 2p^5 (^2P_{1/2}) 4s$	$(^1/2, ^1/2)^o$	1	3 781 900
$2s^2 2p^5 (^2P_{3/2}) 3d$	$^2[3/2]^o$	1	3 239 700	$2s^2 2p^5 (^2P_{3/2}) 4d$	$^2[3/2]^o$	1	3 919 000
$2s^2 2p^5 (^2P_{1/2}) 3d$	$^2[3/2]^o$	1	3 284 300	$2s^2 2p^5 (^2P_{1/2}) 4d$	$^2[3/2]^o$	1	3 948 400
				Ca XII ($^2P_{3/2}$)	<i>Limit</i>		4 774 000

Ca XII

 $Z = 20$

F I isoelectronic sequence

Ground state: $1s^2 2s^2 2p^5 \ ^2P^o_{3/2}$ Ionization energy = 5 301 000 cm⁻¹ (657.2 eV)

The first work on this spectrum was by Edlén and Tyrén (1936), who classified 10 lines of the $2s^2 2p^5 - 2s^2 2p^4 3s$ and $3d$ transition arrays between 27 and 33 Å. This work was extended by Feldman, Doschek, Cowan, and Cohen (1973), from whose wavelengths the $3s$ and $3d$ levels are determined. Fawcett, Burgess, and Peacock (1967) identified the $2s^2 2p^5 - 2s^2 2p^6$ doublet at ~ 140 Å. The $2s^2 2p^5 \ ^2P^o$ term interval is obtained from the solar flare line at 3327.5 Å (in air) identified by Edlén (1942, 1976).

The $2s^2 2p^5 3s \ ^2P^o$ term is from Feldman et al. (1973).

The ionization energy was obtained by extrapolation by Lotz (1967).

References

- Edlén, B. (1942), Z. Astrophys. **22**, 30.
 Edlén, B. (1976), Mem. Soc. Roy. Sc. Liege **9**, 235.
 Edlén, B., and Tyrén, F. (1936), Z. Phys. **101**, 206.
 Fawcett, B. C., Burgess, D. D., and Peacock, N. J. (1967), Proc. Phys. Soc. **91**, 970.
 Feldman, U., Doschek, G. A., Cowan, R. D., and Cohen, L. (1973), J. Opt. Soc. Am. **63**, 1445.
 Lotz, W. J. (1967), J. Opt. Soc. Am. **57**, 873.

Ca XIII

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
$2s^2 2p^5$	$^2P^o$	$\frac{3}{2}$ $\frac{1}{2}$	0 30 044	$2s^2 2p^4 (^3P) 3d$	2P	$\frac{1}{2}$ $\frac{3}{2}$	3 486 700 3 508 200
$2s^2 2p^6$	2S	$\frac{1}{2}$	709 000	$2s^2 2p^4 (^3P) 3d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	3 494 600 3 511 500
$2s^2 2p^4 (^3P) 3s$	4P	$\frac{5}{2}$ $\frac{3}{2}$ $\frac{1}{2}$	3 062 300 3 077 100 3 089 300	$2s^2 2p^4 (^3P) 3d$	2F	$\frac{5}{2}$	3 494 900
$2s^2 2p^4 (^3P) 3s$	2P	$\frac{3}{2}$ $\frac{1}{2}$	3 097 800 3 114 800	$2s^2 2p^4 (^1D) 3d$	2S	$\frac{1}{2}$	3 559 300
$2s^2 2p^4 (^1D) 3s$	2D	$\frac{5}{2}$ $\frac{3}{2}$	3 158 500 3 159 300	$2s^2 2p^4 (^1D) 3d$	2P	$\frac{3}{2}$	3 574 900
$2s^2 2p^4 (^1S) 3s$	2S	$\frac{1}{2}$	3 249 600	$2s^2 2p^4 (^1D) 3d$	2D	$\frac{5}{2}$ $\frac{3}{2}$	3 574 900 3 584 900
$2s^2 2p^4 (^3P) 3d$	4P	$\frac{1}{2}$ $\frac{3}{2}$ $\frac{5}{2}$	3 475 800 3 479 600 3 489 400	$2s^2 2p^4 (^1S) 3d$	2D	$\frac{5}{2}$ $\frac{3}{2}$	3 647 900 3 652 300
$2s^2 2p^4 (^3P) 3d$	4F	$\frac{5}{2}$	3 480 000	$2s^2 p^5 (^3P^o) 3s$	$^2P^o$	$\frac{3}{2}$ $\frac{1}{2}$	3 738 000 3 755 900
				Ca XIII (3P_2)	Limit		5 301 000

Ca XIII

 $Z=20$

O I isoelectronic sequence

Ground state: $1s^2 2s^2 2p^4 \ ^3P_2$ Ionization energy = 5 861 000 cm⁻¹ (726.6 eV)

The observed spectrum of Ca XIII consists of the strong transition array $2s^2 2p^4 - 2s2p^5$, which lies between 130 and 170 Å, and the arrays $2p^4 - 2p^3 3s$ at 29 Å and $2p^4 - 2p^3 3d$ at 26 Å. The $2s2p^5 \ ^3P^o$ is taken from Fawcett, Burgess, and Peacock (1967). Two lines assigned to the ${}^1P^o$ are inconsistent so the term value cannot be determined. The transition $2p^6 \ ^1S_0$ to $2s2p^5 \ ^1P^o$ lies at 169.49 Å, according to Fawcett, Galanti and Peacock (1974).

The configuration $2p^3 3s$ is from the observations of Doschek, Feldman, and Cohen (1973). Lines of the $2p^4 - 2p^3 3d$ array were classified by Fawcett and Hayes (1975). Subsequently, revisions of several line classifications were suggested by Bromage and Fawcett (1977) on the basis of new calculations. We regard these changes as tentative and therefore omit the corresponding levels.

Levels of the $2s^2 2p^4$ configuration are obtained from Edlén (1972). He gives a calculated value for the position of the 1D_2 , on which we base the singlet system.

The ionization energy is from Lotz's (1967) extrapolation.

References

- Bromage, G. E., and Fawcett, B. C. (1977), Mon. Not. R. Astr. Soc. **178**, 591.
 Doschek, G. A., Feldman, U., and Cohen, L. (1973), J. Opt. Soc. Am. **63**, 1463.
 Edlén, B. (1972), Sol. Phys. **24**, 356.
 Fawcett, B. C., Burgess, D. D., and Peacock, N. J. (1967), Proc. Phys. Soc. **91**, 970.
 Fawcett, B. C., Galanti, M., and Peacock, N. J. (1974), J. Phys. B **7**, 1149.
 Fawcett, B. C., and Hayes, R. W. (1975), Mon. Not. R. Astr. Soc. **170**, 185.
 Lotz, W. J. (1967), J. Opt. Soc. Am. **57**, 873.

Ca XIII

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Configuration	Term	<i>J</i>	Level (cm ⁻¹)
$2s^2 2p^4$	3P	2	0	$2s^2 2p^3 ({}^2P^o) 3s$	${}^1P^o$	1	$3\ 544\ 600 + x$
		1	24 460		${}^3D^o$	2	$3\ 739\ 000$
		0	28 830			3	$3\ 743\ 000$
$2s^2 2p^4$	1D	2	88 400 + <i>x</i>	$2s^2 2p^3 ({}^2D^o) 3d$	${}^3D^o$	1	$3\ 828\ 000$
$2s^2 2p^4$	1S	0	178 560 + <i>x</i>			2	$3\ 839\ 000$
$2s2p^5$	${}^3P^o$	2	618 260	$2s^2 2p^3 ({}^2D^o) 3d$	${}^3P^o$	2	$3\ 846\ 000$
		1	638 260				
		0	650 160	$2s^2 2p^3 ({}^2D^o) 3d$	${}^3S^o$	1	$3\ 864\ 000$
$2s^2 2p^3 ({}^4S^*) 3s$	${}^3S^o$	1	$3\ 374\ 600$	$2s^2 2p^3 ({}^2P^o) 3d$	${}^3D^o$	1	$3\ 909\ 000$
$s^2 2p^3 ({}^2D^o) 3s$	${}^3D^o$	1	$3\ 452\ 700$			3	$3\ 917\ 000$
		2	$3\ 453\ 200$			2	$3\ 924\ 000$
		3	$3\ 458\ 300$	Ca XIV (${}^4S_{3/2}^o$)	<i>Limit</i>		5 861 000
$2s^2 2p^3 ({}^2D^o) 3s$	${}^1D^o$	2	$3\ 474\ 800 + x$				

Ca XIV

Z = 20

N I isoelectronic sequence

Ground state: $1s^2 2s^2 2p^3 \ ^4S^o_{3/2}$ Ionization energy = 6 595 000 cm⁻¹ (817.6 eV)

The levels of the configurations $2s^2 2p^3$, $2s2p^4$, and $2p^5$ are from the measurements and classifications of Kononov, Koshelev, Podobedova, and Churilov (1976). The position of the doublets relative to the ground state is based on the estimated position of $2s^2 2p^3 \ ^2D^o_{5/2}$ by Eldén (1972).

The $2p^3 3d$ terms are from Fawcett and Hayes (1975); their percentage compositions were calculated by Bromage and Fawcett (1977).

The ionization energy is from Lotz's (1967) extrapolation.

References

- Bromage, G. E., and Fawcett, B. C. (1977), Mon. Not. R. Astr. Soc. **179**, 683.
 Edlén, B. (1972), Solar Physics **24**, 356.
 Fawcett, B. C., and Hayes, R. W. (1975), Mon. Not. R. Astr. Soc. **170**, 185.
 Kononov, E. Y., Koshelev, K. N., Podobedova, L. I., and Churilov, S. S. (1976), Opt. Spectrosc. **40**, 121.
 Lotz, W. (1967), J. Opt. Soc. Am. **57**, 873.

Ca XIV

Configuration	Term	J	Level (cm ⁻¹)	Leading percentages		
$2s^2 2p^3$	$^4S^o$	$\frac{3}{2}$	0			
$2s^2 2p^3$	$^2D^o$	$\frac{3}{2}$	105 120+x			
		$\frac{5}{2}$	112 740+x			
$2s^2 2p^3$	$^2P^o$	$\frac{1}{2}$	171 640+x			
		$\frac{3}{2}$	182 570+x			
$2s2p^4$	4P	$\frac{5}{2}$	515 840			
		$\frac{3}{2}$	535 910			
		$\frac{1}{2}$	545 110			
$2s2p^4$	2D	$\frac{3}{2}$	709 970+x			
		$\frac{5}{2}$	711 760+x			
$2s2p^4$	2S	$\frac{1}{2}$	824 360+x			
$2s2p^4$	2P	$\frac{3}{2}$	857 510+x			
		$\frac{1}{2}$	884 910+x			
$2p^5$	$^2P^o$	$\frac{3}{2}$	1 347 180+x			
		$\frac{1}{2}$	1 379 450+x			
$2s^2 2p^2(^3P)3d$	2P	$\frac{3}{2}$	4 113 200+x	57	28	$(^3P) \ ^4D$
$2s^2 2p^2(^3P)3d$	4P	$\frac{5}{2}$	4 143 700	76	19	$(^3P) \ ^4D$
		$\frac{3}{2}$	4 151 800	91		
$2s^2 2p^2(^3P)3d$	2F	$\frac{7}{2}$	4 153 300+x	57	20	$(^1D) \ ^2F$
$2s^2 2p^2(^3P)3d$	2D	$\frac{5}{2}$	4 198 100+x	78	14	$(^1D) \ ^2D$
$2s^2 2p^2(^1D)3d$	2F	$\frac{7}{2}$	4 229 800+x	60	32	$(^3P) \ ^2F$
		$\frac{5}{2}$	4 242 800+x	48	21	

Ca xiv—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Leading percentages	
$2s^2 2p^2 (^1D) 3d$	² D	$\frac{5}{2}$	4 241 000+ <i>x</i>	67	18 (^1D) ² F
$2s^2 2p^2 (^1D) 3d$	² P	$\frac{3}{2}$	4 250 400+ <i>x</i>	90	
Ca xv (³ P ₀)	<i>Limit</i>		6 595 000		

Ca xv

 $Z=20$ C₁ isoelectronic sequenceGround state: 1s²2s²2p² 3P₀Ionization energy = 7 215 000 cm⁻¹ (894.5 eV)

The levels of the 2s²2p², 2s2p³, and 2p⁴ configurations were determined by Kononov, Koshelev, Podobedova, and Churilov (1976), from observations of a laser plasma between 130 and 270 Å. The fine-structure of the 2s²2p² 3P term is determined from the solar coronal lines 5444 Å and 5693.6 Å as given by Edlén (1972).

The levels of the higher configurations are from the measurements of Fawcett and Hayes (1975) and Bromage and Fawcett (1977).

No intersystem combinations have been observed in the laboratory. Edlén (1972) has extrapolated the position of 2p² 1D₂ to 108 561 cm⁻¹. Sandlin, Brueckner, and Tousey (1977) have identified the 3P₂-1D₂ forbidden transition in 2s²2p² at 1375.95 Å in the solar corona. This put 1D₂ at 108 600 cm⁻¹ above the ground state. We have used that value as the reference value for the singlet system.

The percentage compositions for the 2p3d configuration were calculated by Bromage and Fawcett.

The ionization energy is Lotz's (1967) extrapolation.

References

- Bromage, G. E., and Fawcett, B. C. (1977), Mon. Not. R. Astr. Soc. **178**, 605.
 Edlén, B. (1972), Solar Physics **24**, 356.
 Fawcett, B. C., and Hayes, R. W. (1975), Mon. Not. R. Astr. Soc. **170**, 185.
 Kononov, E. Y., Koshelev, K. N., Podobedova, L. I., and Churilov, S. S. (1976), Opt. Spectrosc. **40**, 121.
 Lotz, W. (1967), J. Opt. Soc. Am. **57**, 873.
 Sandlin, G. D., Brueckner, G. E., and Tousey, R. (1977), Astrophys. J. **241**, 898.

Ca xv

Configuration	Term	J	Level (cm ⁻¹)	Leading percentages
2s ² 2p ²	3P	0	0	
		1	17 559	
		2	35 923	
2s ² 2p ²	1D	2	108 600	
2s ² 2p ²	1S	0	197 620	
2s2p ³	3D°	2	496 680	
		1	497 590	
		3	500 240	
2s2p ³	3P°	0	581 730	
		1	582 840	
		2	585 660	
2s2p ³	3S°	1	728 910	
2s2p ³	1D°	2	729 720	
2s2p ³	1P°	1	814 870	
2p ⁴	3P	2	1 107 570	
		1	1 133 870	
		0	1 140 140	
2p ⁴	1D	2	1 195 200	
2p ⁴	1S	0	1 254 120	

Ca xv—Continued

Configuration	Term	<i>J</i>	Level (cm ⁻¹)	Leading percentages		
$2s^2 2p3d$	$^3F^\circ$	2	$4\ 363\ 000$	74	25	$^1D^\circ$
		3	$4\ 379\ 000$	88		
$2s^2 2p3d$	$^3D^\circ$	1	$4\ 399\ 000$	78	15	$^3P^\circ$
		2	$4\ 412\ 000$	45	24	$^1D^\circ$
		3	$4\ 426\ 000$	88	10	$^3F^\circ$
$2s^2 2p3d$	$^3P^\circ$	1	$4\ 435\ 000$	82	17	$^3D^\circ$
		2	$4\ 435\ 000$	61	36	
$2s^2 2p3d$	$^1P^\circ$	1	$4\ 473\ 000$	92		
$2s^2 2p3d$	$^1F^\circ$	3	$4\ 475\ 000$	95		
$2s2p^2(^4P)3d$	3F	4	$4\ 727\ 000$			
Ca XVI (${}^3P_{1/2}$)	<i>Limit</i>		$7\ 215\ 000$			

Ca xvi

 $Z = 20$

B 1 isoelectronic sequence

Ground state: $1s^2 2s^2 2p\ ^2P^o_{1/2}$ Ionization energy = 7 860 000 cm^{-1} (974 eV)

The low lying configurations $2s^2 2p$, $2s 2p^2$, and $2p^3$ are from the observations of Kononov, Koshelev, Podobedova and Churilov (1975). We have assigned the lowest quartet level the value calculated by them, since there are no observed connections of the quartet terms with the doublet terms.

The high levels are from Fawcett and Hayes (1975) and the ionization energy is from the extrapolation of Lotz (1967).

References

- Kononov, E. Y., Koshelev, K. N., Podobedova, L. I., and Churilov, S. S. (1975), Opt. Spectrosc. **39**, 458.
 Fawcett, B. C., and Hayes, R. W. (1975), Mon. Not. R. Astr. Soc. **170**, 185.
 Lotz, W. (1967), J. Opt. Soc. Am. **57**, 873.

Ca xvi

Configuration	Term	J	Level (cm^{-1})	Configuration	Term	J	Level (cm^{-1})
$2s^2 2p$	$^2P^o$	$\frac{1}{2}$ $\frac{3}{2}$	0 36 600	$2s^2 3d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	4 662 000 4 664 000
$2s 2p^2$	4P	$\frac{1}{2}$ $\frac{3}{2}$ $\frac{5}{2}$	269 000 + x 283 560 + x 301 790 + x	$2s 2p (^3P^o) 3p$	2P	$\frac{1}{2}$ $\frac{3}{2}$	4 773 000 4 794 000
$2s 2p^2$	2D	$\frac{3}{2}$ $\frac{5}{2}$	479 460 481 960	$2s 2p (^3P^o) 3d$	$^4D^o$	$\frac{3}{2}$ $\frac{5}{2}$ $\frac{7}{2}$	4 931 000 + x 4 935 000 + x 4 953 000 + x
$2s 2p^2$	2S	$\frac{1}{2}$	592 240	$2s 2p (^3P^o) 3d$	$^4P^o$	$\frac{5}{2}$	4 964 000 + x
$2s 2p^2$	2P	$\frac{1}{2}$ $\frac{3}{2}$	633 890 645 770	$2s 2p (^3P^o) 3d$	$^2F^o$	$\frac{5}{2}$ $\frac{7}{2}$	5 000 000 5 022 000
$2p^3$	$^4S^o$	$\frac{5}{2}$	835 920 + x	$2s 2p (^1P^o) 3d$	$^2F^o$	$\frac{7}{2}$	5 170 000
$2p^3$	$^2D^o$	$\frac{3}{2}$ $\frac{5}{2}$	940 060 944 830	$2s 2p (^1P^o) 3d$	$^2D^o$	$\frac{5}{2}$	5 198 000
$2p^3$	$^2P^o$	$\frac{1}{2}$ $\frac{3}{2}$	1 053 000 1 062 090	Ca XVII (1S_0)	<i>Limit</i>		7 860 000

Ca xvii

 $Z=20$

Be I isoelectronic sequence

Ground state: $1s^2 2s^2 ^1S_0$ Ionization energy = 8 770 000 cm⁻¹ (1087 eV)

The strong singlet resonance line $2s^2 ^1S_0 - 2s2p ^1P^{\circ}_1$ at 192.8 Å has been observed in the laboratory by Kononov, Koshelev, Podobedova, and Churilov (1975), by Fawcett and Hayes (1975) and in solar flares. We used the wavelength of Kononov et al. to establish the value of the $^1P^{\circ}$ term.

The intersystem transition $2s^2 ^1S_0 - 2s2p ^3P^{\circ}$, has not been observed in the laboratory but has been identified in a solar flare spectrum by Sandlin, Brueckner, Scherrer, and Tousey (1976). We have adopted their value to locate the triplet system relative to the singlets. The 3P terms of $2s2p$ and $2p^2$ and the 1S term of $2p^2$ are from Kononov et al. The $2p^2 ^1D$ is from Goldsmith, Oren, Crooker, and Cohen (1973).

The higher configurations are from Fawcett and Hayes. The ionization energy is from Lotz's (1967) extrapolation.

References

- Fawcett, B. C., and Hayes, R. W. (1975), Mon. Not. R. Astr. Soc. **170**, 185.
 Goldsmith, S., Oren, L., Crooker, A. M., and Cohen, L. (1973), Astrophys. J. **184**, 1021.
 Kononov, E. Y., Koshelev, K. N., Podobedova, L. I., and Churilov, S. S. (1975), Opt. Spectrosc. **39**, 458.
 Lotz, W. (1967), J. Opt. Soc. Am. **57**, 873.
 Sandlin, G. D., Brueckner, G. E., Scherrer, V. E., and Tousey, R. (1976), Astrophys. J. **205**, L47.

Ca xvii

Configuration	Term	J	Level (cm ⁻¹)	Configuration	Term	J	Level (cm ⁻¹)
$2s^2$	1S	0	0	$2s3p$	$^1P^{\circ}$	1	5 113 000
$2s2p$	$^3P^{\circ}$	0	258 290	$2s3d$	3D	2	5 186 000
		1	269 460			3	5 190 000
		2	296 950		1D	2	5 236 000
$2s2p$	$^1P^{\circ}$	1	518 620	$2p3p$	3D	3	5 448 000
$2p^2$	3P	1	706 680	$2p3d$	$^3D^{\circ}$	2	5 533 000
		2	726 450			3	5 546 000
$2p^2$	1D	2	801 710	$2p3d$	$^1F^{\circ}$	3	5 602 000
$2p^2$	1S	0	967 330	Ca XVIII ($^2S_{1/2}$)	<i>Limit</i>		8 770 000

Ca xviii

 $Z=20$

Li I isoelectronic sequence

Ground state: $1s^2 2s \ ^2S_{1/2}$ Ionization energy = 9 332 000 cm⁻¹ (1157.0 eV)

The $2s-3p$, $2s-4p$, $2s-5p$, $2p-3s$, $2p-3d$, $2p-4d$, $2p-5d$, and $2p-6d$ transitions were reported by Goldsmith, Feldman, Oren, and Cohen (1972). The value of the $2p \ ^2P^\circ$ term is from the $2s-2p$ transition observed at 300 Å in a solar flare by Widing and Purcell (1976).

Boiko, Faenov, and Pikuz (1978) confirmed the lines identified by Goldsmith et al. and added the $6p$, $7p$, $6d$, and $7d$ terms. The doubly excited levels were obtained from lines observed by Aglitskii, Boiko, Zakharov, Pikuz, and Faenov (1974) at 3.2 Å in a laser-produced plasma.

We derived the ionization energy from the nd Rydberg series ($n=4$ to 6).

References

- Aglitskii, E. V., Boiko, V. A., Zakharov, S. M., Pikuz, S. A., and Faenov, A. Y. (1974), Kvantovaya Elektron. (Moscow) **1**, 908.
 Boiko, V. A., Faenov, A. Y., and Pikuz, S. A. (1978), J. Quant. Spectrosc. Radiat. Transfer **19**, 11.
 Goldsmith, S., Feldman, U., Oren, L., and Cohen, L. (1972), Astrophys. J. **174**, 209.
 Widing, K. G. and Purcell, J. D. (1976), Astrophys. J. **204**, L151.

Ca xviii

Configuration	Term	J	Level (cm ⁻¹)	Configuration	Term	J	Level (cm ⁻¹)
$1s^2 2s$	2S	$\frac{1}{2}$	0	$1s^2 5p$	$^2P^\circ$	$\frac{1}{2}, \frac{3}{2}$	7 914 000
$1s^2 2p$	$^2P^\circ$	$\frac{1}{2}, \frac{3}{2}$	290 060 330 920	$1s^2 6p$	$^2P^\circ$	$\frac{1}{2}, \frac{3}{2}$	8 341 000
$1s^2 3s$	2S	$\frac{1}{2}$	5 277 000	$1s^2 6d$	2D	$\frac{5}{2}$	8 345 000
$1s^2 3p$	$^2P^\circ$	$\frac{1}{2}, \frac{3}{2}$	5 338 000 5 350 000	$1s^2 7p$	$^2P^\circ$	$\frac{1}{2}, \frac{3}{2}$	8 605 000
$1s^2 3d$	2D	$\frac{3}{2}, \frac{5}{2}$	5 381 000 5 384 000	$1s^2 7d$	2D	$\frac{3}{2}, \frac{5}{2}$	8 610 000
$1s^2 4p$	$^2P^\circ$	$\frac{1}{2}, \frac{3}{2}$	7 101 000	Ca XIX (1S_0)	Limit		9 332 000
$1s^2 4d$	2D	$\frac{3}{2}, \frac{5}{2}$	7 112 000 7 116 000	$1s(^2S)2s2p(^1P^\circ)$	$^2P^\circ$	$\frac{3}{2}$	31 352 000
$1s^2 5d$	2D	$\frac{5}{2}, \frac{3}{2}$	7 912 000 7 913 000	$1s2p^2$	2D	$\frac{3}{2}, \frac{5}{2}$	31 484 000 31 486 000
				$1s2p^2$	2P	$\frac{3}{2}$	31 551 000
				$1s2p3p$			36 784 000

Ca xix

 $Z=20$

He I isoelectronic sequence

Ground state: $1s^2 \ ^1S_0$ Ionization energy = $41\ 369\ 120 \pm 40\ \text{cm}^{-1}$ ($5129.16 \pm 0.01\ \text{eV}$)

The theoretical values calculated by Ermolaev and Jones (1974) for the singlet and triplet S and P terms of this two-electron ion are more accurate than the observed values, and we have quoted them up to $n=4$. The uncertainty in the calculation is estimated by the authors to be one part in 10^{-6} or 10^{-7} . For comparison, the $1s^2$ - $1s2p$ transition of this ion has been observed by Aglitskii et al.

(1974) in a laser-produced plasma. They place $1s2p$ $^3P^o$ at $31\ 322\ 000\ \text{cm}^{-1}$ and $1s2p$ $^1P^o_1$ at $31\ 480\ 000\ \text{cm}^{-1}$.

References

Ermolaev, A. M. and Jones, M. (1974), J. Phys. B 7, 199.
 Aglitskii, E. V., Bolko, V. A., Zakharov, S. M., Pikuz, S. A., and Faenov, A. Y. (1974), Kvantovaya Elektron. (Moscow) 1, 908.

Ca xix

Configuration	Term	J	Level (cm^{-1})	Configuration	Term	J	Level (cm^{-1})
$1s^2$	1S	0	0	$1s3s$	1S	0	$36\ 922\ 600$
$1s2s$	3S	1	$31\ 145\ 160$	$1s3p$	$^1P^o$	1	$36\ 965\ 520$
$1s2p$	$^3P^o$	0	$31\ 316\ 020$	$1s4s$	3S	1	$38\ 853\ 340$
		1	$31\ 323\ 600$	$1s4p$	$^3P^o$	0	$38\ 872\ 900$
		2	$31\ 358\ 980$			1	$38\ 873\ 820$
$1s2s$	1S	0	$31\ 330\ 780$			2	$38\ 878\ 300$
$1s2p$	$^1P^o$	1	$31\ 477\ 040$	$1s4s$	1S	0	$38\ 873\ 200$
$1s3s$	3S	1	$36\ 873\ 600$	$1s4p$	$^1P^o$	1	$38\ 891\ 350$
$1s3p$	$^3P^o$	0	$36\ 920\ 730$	Ca XX (${}^2S_{1/2}$)		<i>Limit</i>	
		1	$36\ 922\ 920$				
		2	$36\ 933\ 520$				

Ca xx $Z=20$ **H I isoelectronic sequence**Ground state: $1s^2S_{1/2}$ Ionization energy = $44\ 117\ 200 \pm 200\ \text{cm}^{-1}$ ($5469.88 \pm 0.02\ \text{eV}$)

The theoretical values calculated by Erikson (1977) for terms of this hydrogen-like ion are given below. No laboratory data were found. The binding energy of the $1s$ electron is given with an uncertainty of $\pm 200\ \text{cm}^{-1}$; the levels measured from the ground state taken as zero will also have this uncertainty.

Doschek (1972) reports wavelengths of several lines of this spectrum identified in solar flares.

References

- Doschek, G. A. (1972), Space Sci. Rev. **13**, 765.
 Erikson, G. W. (1977), J. Phys. Chem. Ref. Data **6**, 831.

Ca xx

Configuration	Term	<i>J</i>	Level (cm^{-1})	Configuration	Term	<i>J</i>	Level (cm^{-1})
$1s$	2S	$\frac{1}{2}$	0	$4p$	$^2P^\circ$	$\frac{1}{2}$ $\frac{3}{2}$	$41\ 361\ 800$ $41\ 369\ 200$
$2p$	$^2P^\circ$	$\frac{1}{2}$ $\frac{3}{2}$	$33\ 069\ 700$ $33\ 129\ 100$	$4s$	2S	$\frac{1}{2}$	$41\ 362\ 100$
$2s$	2S	$\frac{1}{2}$	$33\ 071\ 600$	$4d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	$41\ 369\ 200$ $41\ 371\ 700$
$3p$	$^2P^\circ$	$\frac{1}{2}$ $\frac{3}{2}$	$39\ 213\ 800$ $39\ 231\ 400$	$4f$	$^2F^\circ$	$\frac{5}{2}$ $\frac{7}{2}$	$41\ 371\ 700$ $41\ 372\ 900$
$3s$	2S	$\frac{1}{2}$	$39\ 214\ 400$			<i>Limit</i>	$44\ 117\ 200$
$3d$	2D	$\frac{3}{2}$ $\frac{5}{2}$	$39\ 231\ 400$ $39\ 237\ 200$				