

Current Status of CDSAXS: Is it Fab-Ready?

Daniel F. Sunday, Donald Windover, Wen-li Wu, R. Joseph Kline



National Institute of Standards and Technology

Technology Administration U.S. Department of Commerce

Contributors



NIST

CD-SAXS

- Adam Hannon*
- R. Joseph Kline
- Christopher Liman*
- Christopher Soles
- Daniel Sunday
- Donald Windover
- Wen-li Wu (retired)

CD-SEM

- Andras Vladar
- John Villarrubia

Intel Corporation

- Scott List
- Jasmeet Chawla

SEMATECH

Ben Bunday

APS - DND-CAT

- Steven J. Weigand
- Denis T. Keane





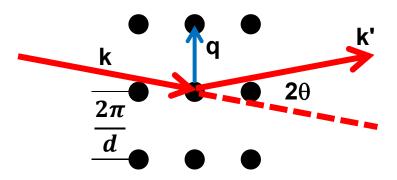


*NIST National Research Council Postdoctoral Fellowship

2 year, \$66k/year stipend to work at NIST, requires US citizenship Other potential opportunities for foreign postdocs Contact me at Joe.Kline@nist.gov for more info



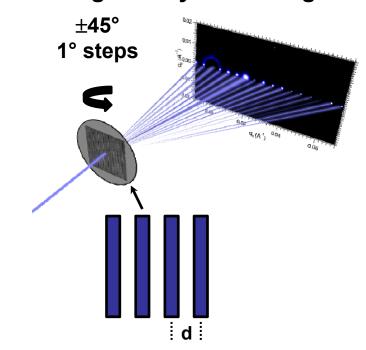
Reciprocal Lattice



 $n\lambda = 2d \sin \theta$ (Bragg's Law)

$$k - k' = q$$
, where $|k| = \frac{2\pi}{\lambda}$

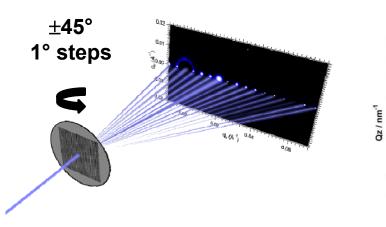
Small Angle X-ray Scattering

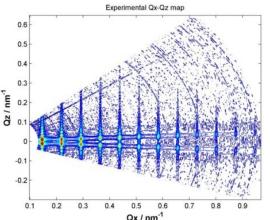


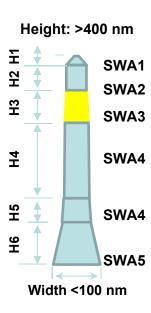
- X-ray diffraction allows the measurement of atomic scale spacings, periodicity, and orientation
- At small angles, the diffraction comes from a periodic nanopattern instead of atoms in a crystal

Critical-Dimension Small Angle X-ray Scattering







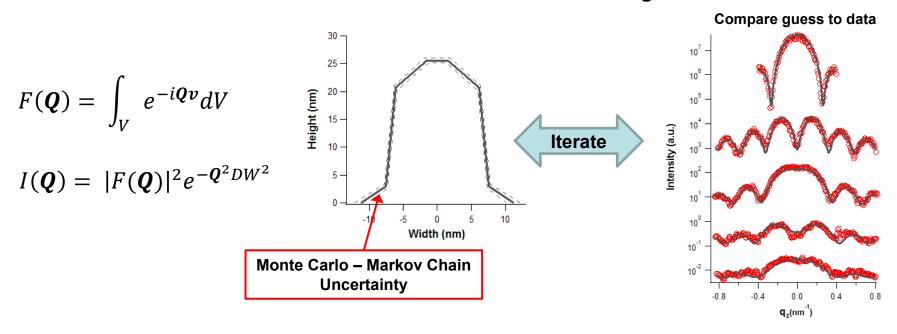


Variable-angle transmission SAXS:

- "Single" crystal diffraction with nanopattern as lattice, pattern shape as atom
- Hard X-rays (>15 keV) for transmission through wafer
- Small spot size (<100 μm)
- Measure 1D, 2D, or 3D periodic nanopatterns
- Result is the average shape of the repeated nanostructure



Inverse Fitting

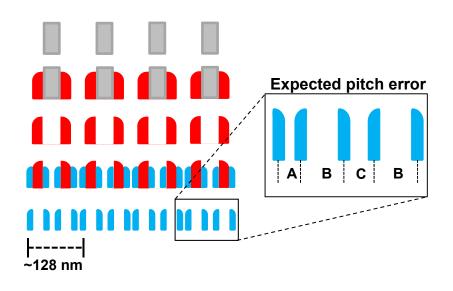


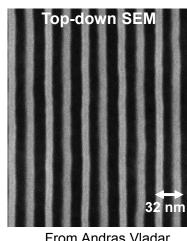
- Diffraction pattern is Fourier transform of electron density distribution
- Partially coherent scattering cannot be directly converted to image
- Use inverse, iterative method to calculate simulated pattern from trial solution
- Use Monte Carlo Markov Chain algorithm to sample parameter space for parameter sensitivity to data



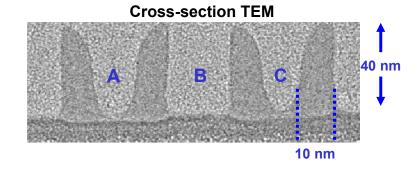


Spacer-Assisted Quadruple Patterning (SAQP)





From Andras Vladar

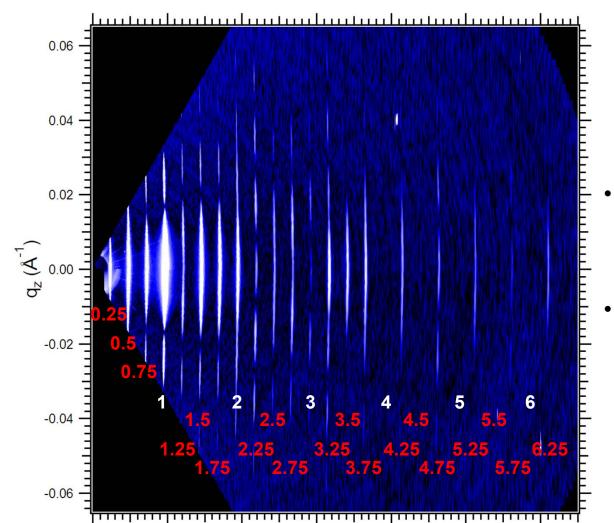


- Measurement goals
 - Average 2D shape, pitch error, and edge roughness
- Sample set with sub-nm controlled variation in pitch error and nominal 32 nm pitch

van Veenhuizen et al. Interconnect Technology Conference, 2012

CDSAXS on 32 nm Pitch Nanopattern

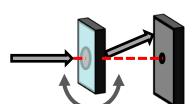




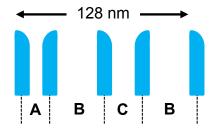
0.08

0.10

0.12



- Composite from 121 images
 - ±60° sample angle
 - 10 s/image = ~30 min/scan
 - Data highly oversampled
- 32 nm, 64 nm, and 128 nm pitches clearly visible
 - Non-integer peaks from pitch quartering error



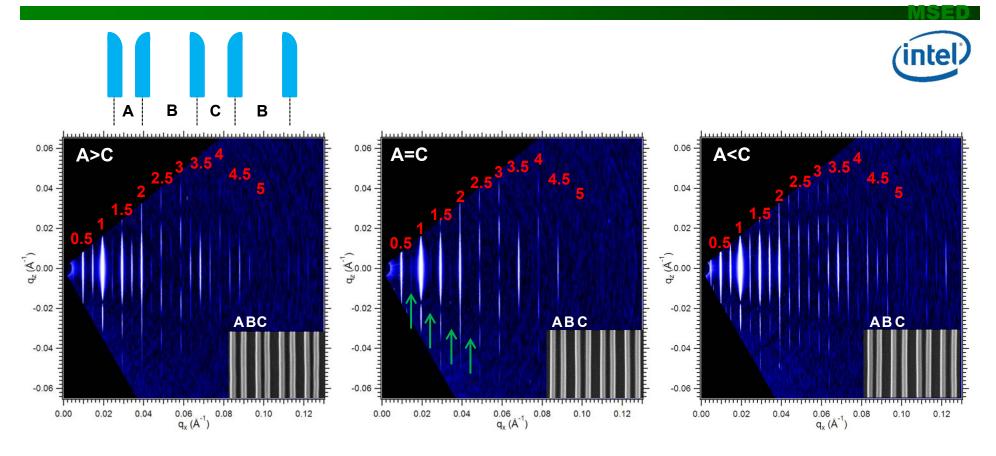
0.02

0.04

0.00

CD-SAXS Measurements on Offset Bias Series



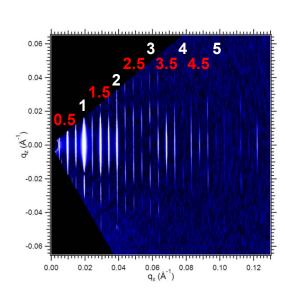


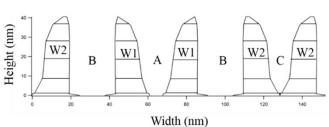
- Sample series has controlled sub-nm variations in A and C
- CD-SAXS measurement is highly sensitive to changes as patterns are visually different
- Sample A=C has weak x.25 and x.75 peaks, indicating 64 nm pitch

$$-A=C\neq B$$

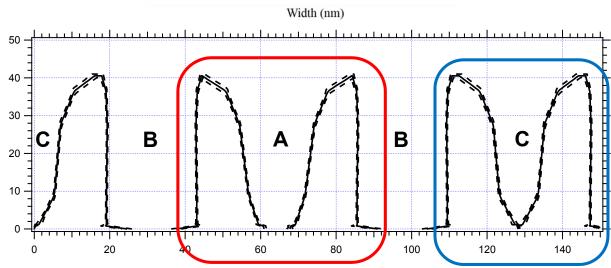
Example CD-SAXS Fit on 32 nm Pitch Nanopattern



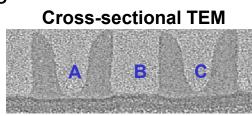






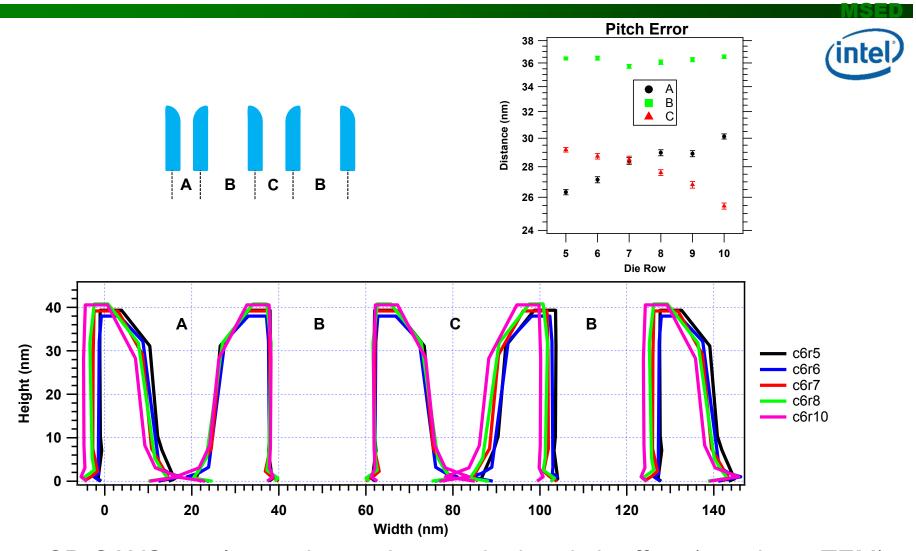


- Sample periodicity is 128 nm
- Uses 6-trapezoid stack with two mirrored line pairs
- Shape matches up to expected value
- Fit uncertainty very small



CD-SAXS Measurement of Pitch Error

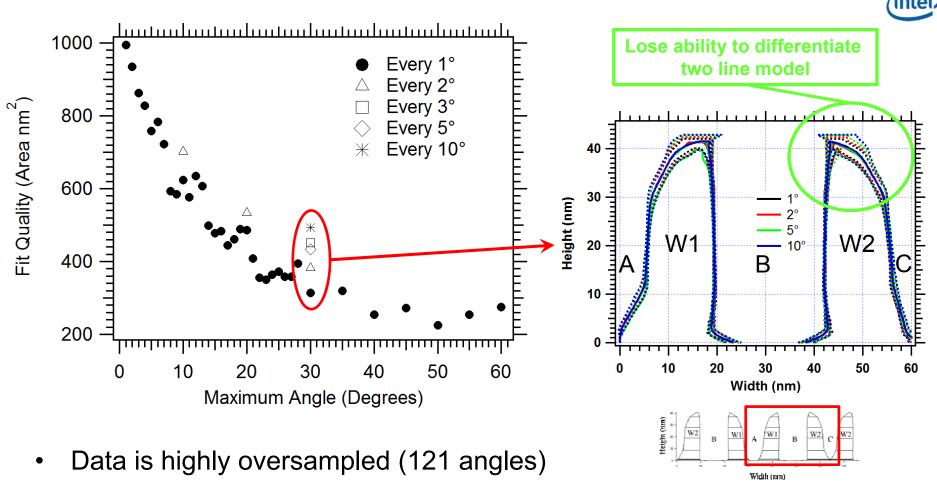




- CD-SAXS resolves sub-nm changes in the pitch offset (matches xTEM)
- Samples had similar line shape with primary difference being shift

Effect of Number of Angles on Fit



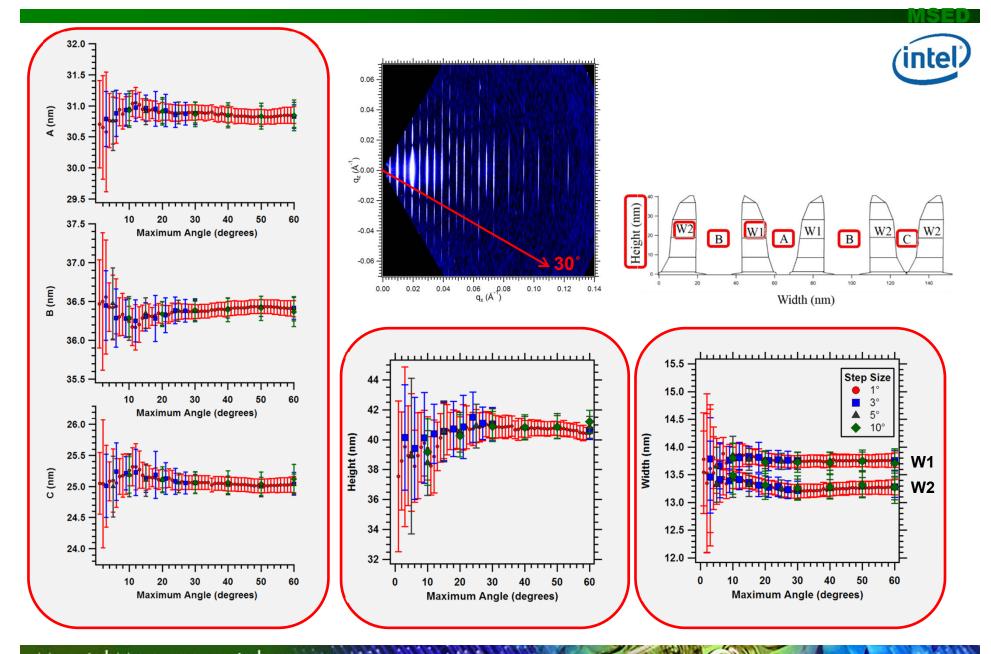


- Compare uncertainty from MCMC algorithm vs number of angles used
- Four angles with 10° step gives reasonable fit

*Fit quality = area between error bars

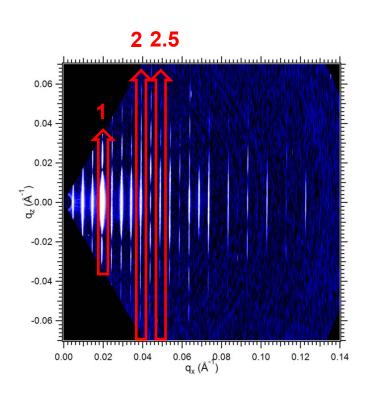
Effect of Maximum Angle on Fit Parameters

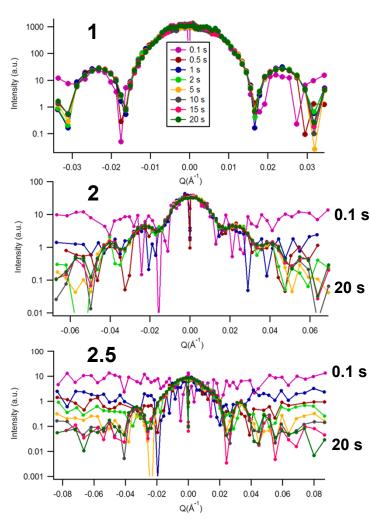




Effect of Signal to Noise on Fit







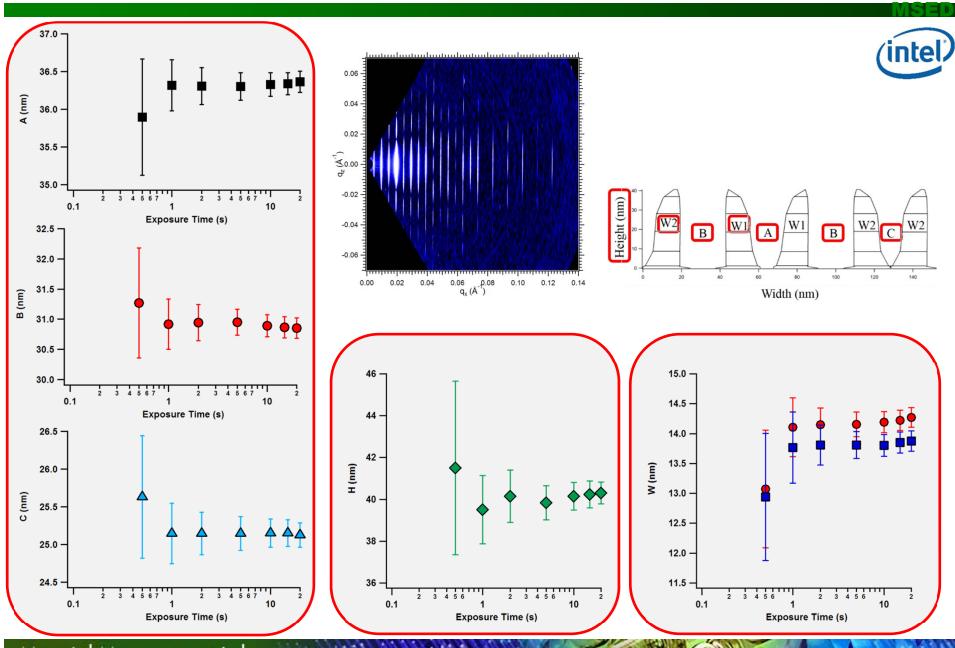
- Exposure times from 0.1 s to 20 s
- Compare uncertainty from MCMC algorithm vs. time



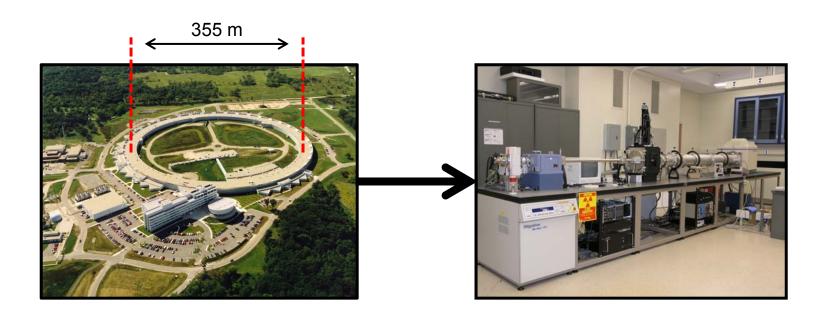
intel

Effect of Signal to Noise on Fit Parameters







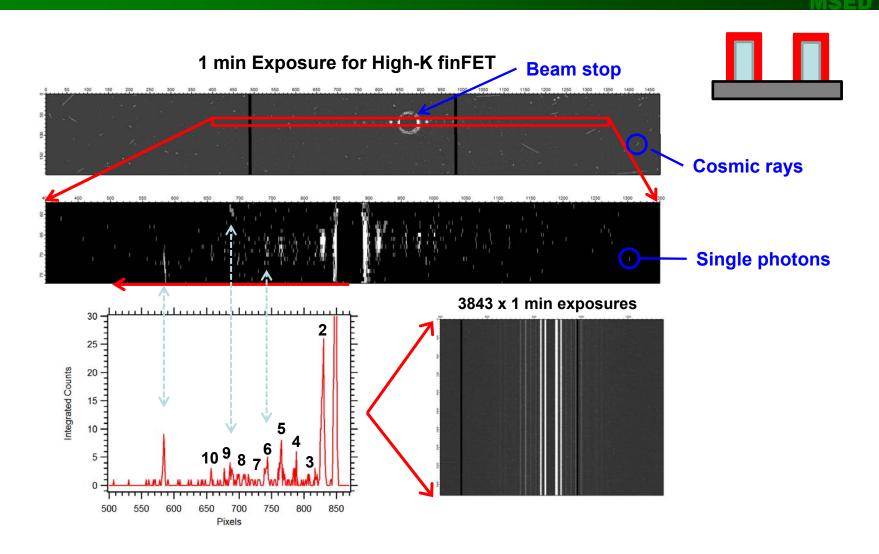


Current system

- Mo Kα micro-focus rotating anode with a multilayer mirror
- Beam size on sample = 300 μm
- Small sample chamber (no 300 mm wafers)
- Noiseless, fast detector
 - Single photon counting with 10⁶ dynamic range

Initial Results from New Detector with Mo Ka



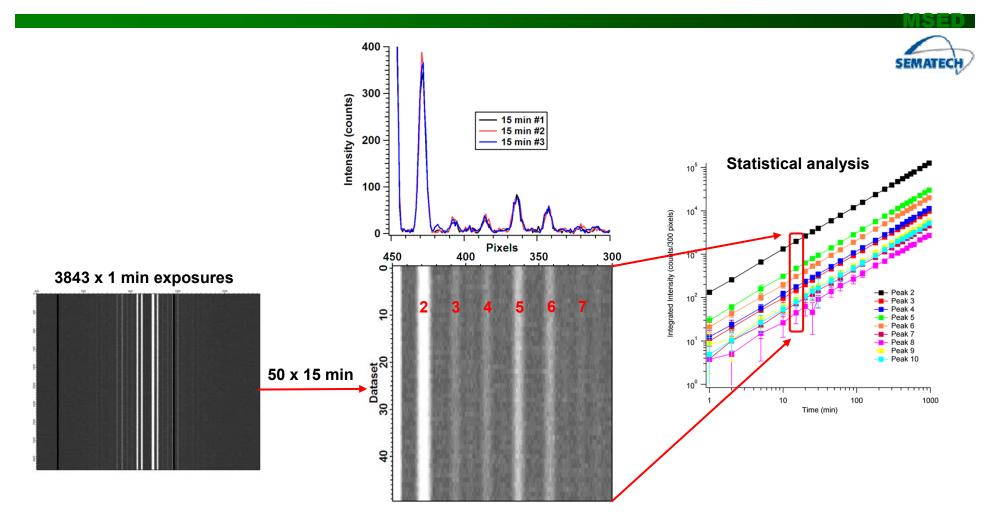


- Detector measures single photons and has no readout noise with fast readout
 - Many short exposures can be combined and allows separation of cosmic from sample scattering



Synthetic Exposure Series and Statistical Analysis

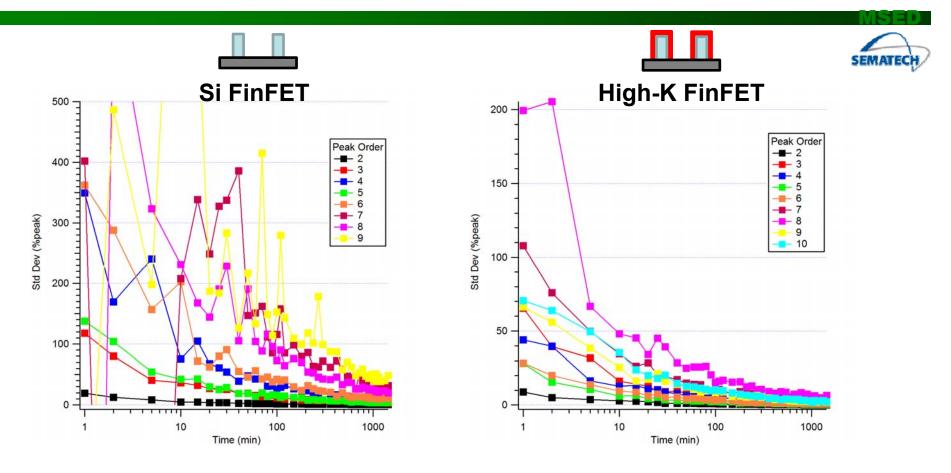




- Resample datasets by randomly combining 1 min exposures into longer time exposure
- Run statistics on synthetic datasets
 - Compare variance in peak intensities across 50 synthetic exposures at each exposure time

Peak Noise vs. Exposure Time

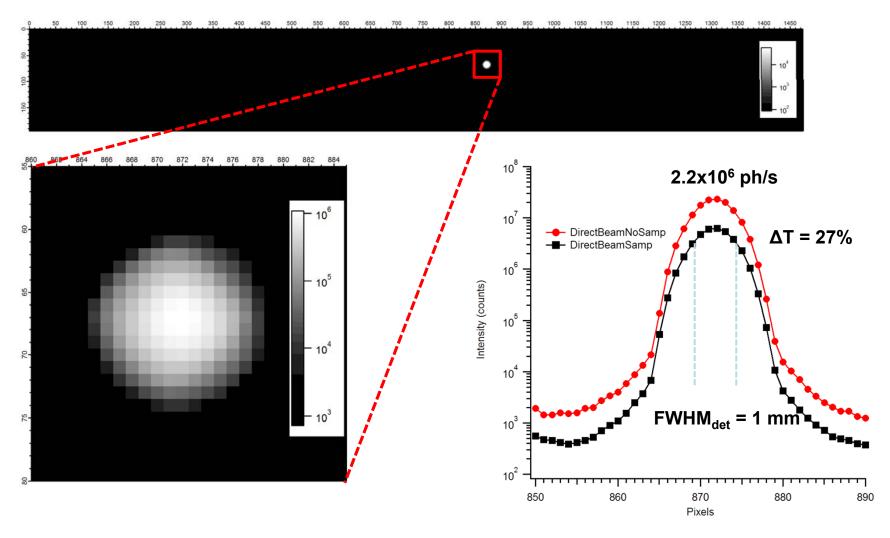




- Values are based on std. dev. of 50 resampled datasets
 - Plot of std. dev. relative to peak intensity
- Noise level much higher in Si FinFET sample
 - Noise is combination of Poisson statistics and cosmic background

Direct Beam Imaging

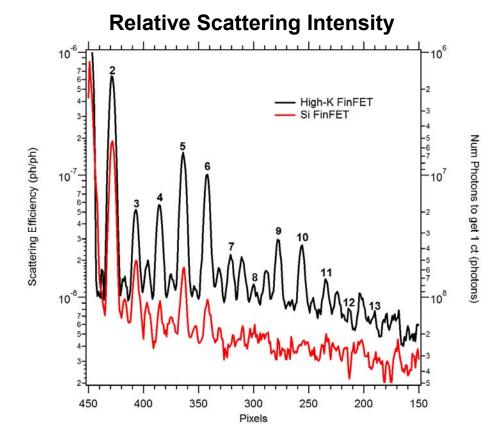




- Direct measurement of beam flux and size
- Allows easy comparison between scattering and beam intensity

Absolute Scattering Intensity and Source Requirements NUST





Ideal Source Requirements

Property	Value	
Energy	>20 keV	
Energy res	<2 % (5 %)	
Divergence	<0.5 x 1.5 mrad	
Spot size	<100 μm	

Flux for 10 sec Measurement

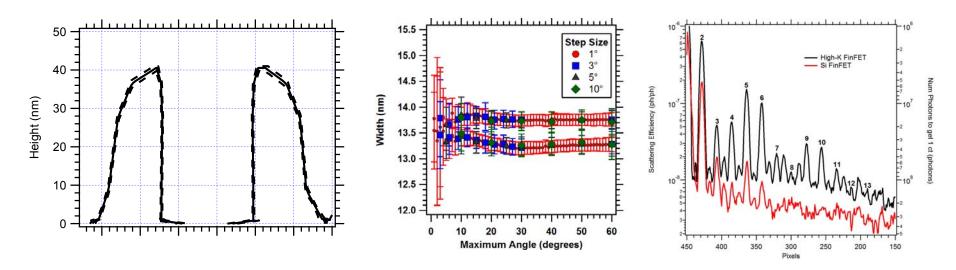
Material	High-K	Si	Resist
Photons	108 ph/s	10 ⁹ ph/s	10 ¹⁰ ph/s

- Data is normalized to beam intensity and silicon absorption
- Assume >10 cts/peak for significance at high q (10th order)



CDSAXS: Is it Fab Ready?





- CDSAXS works great in limit of excess photons
- Evaluated effects of data quantity and quality
- Number of potential new X-ray sources on the horizon
 - Identified critical source requirements for CDSAXS
 - New sources key to success of CDSAXS

*NIST National Research Council Postdoctoral Fellowship

2 year, \$66k/year stipend to work at NIST, requires US citizenship Other potential opportunities for foreign postdocs Contact me at Joe.Kline@nist.gov for more info