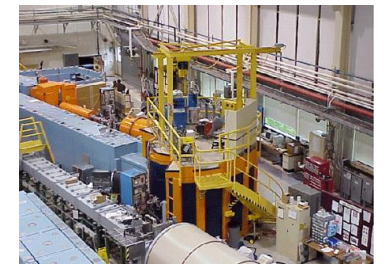
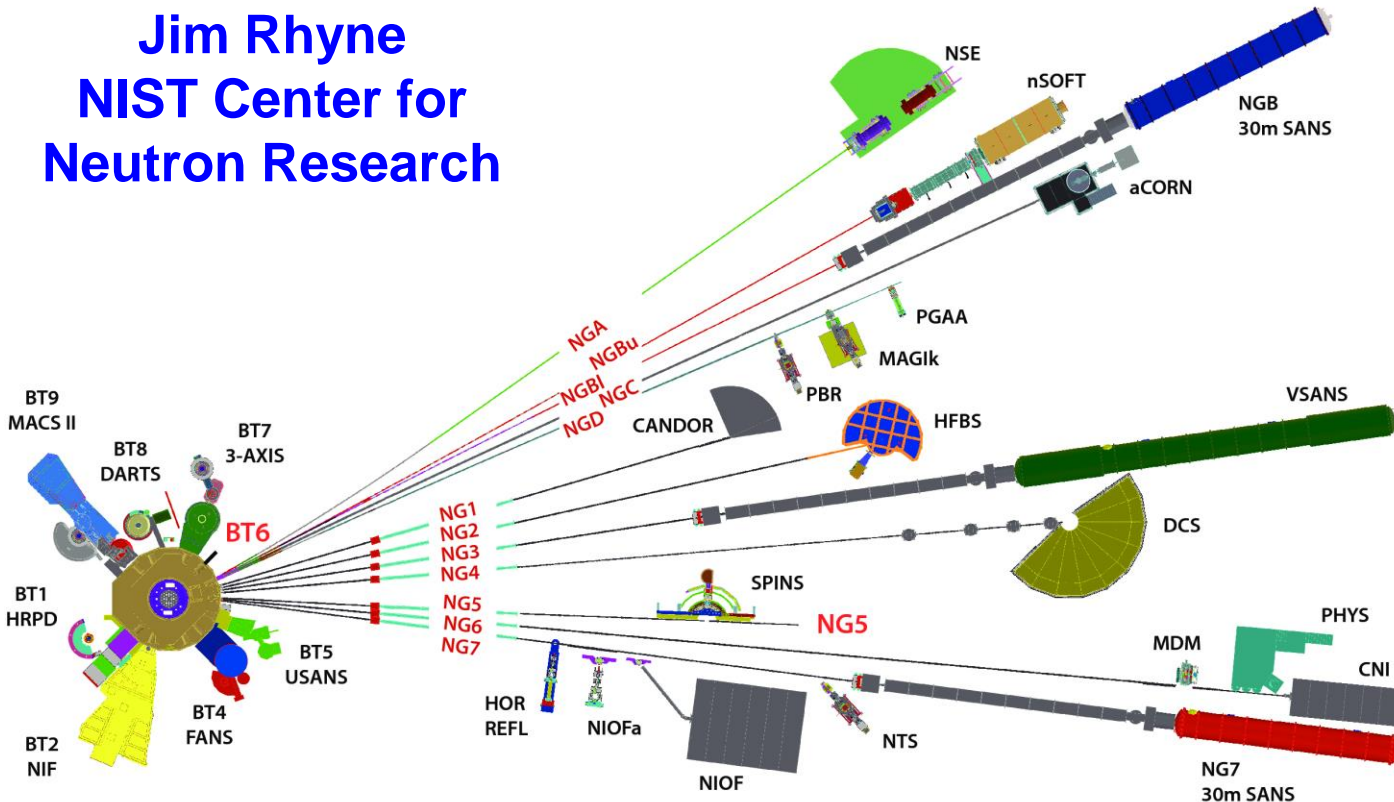


Choosing the Right Spectrometer

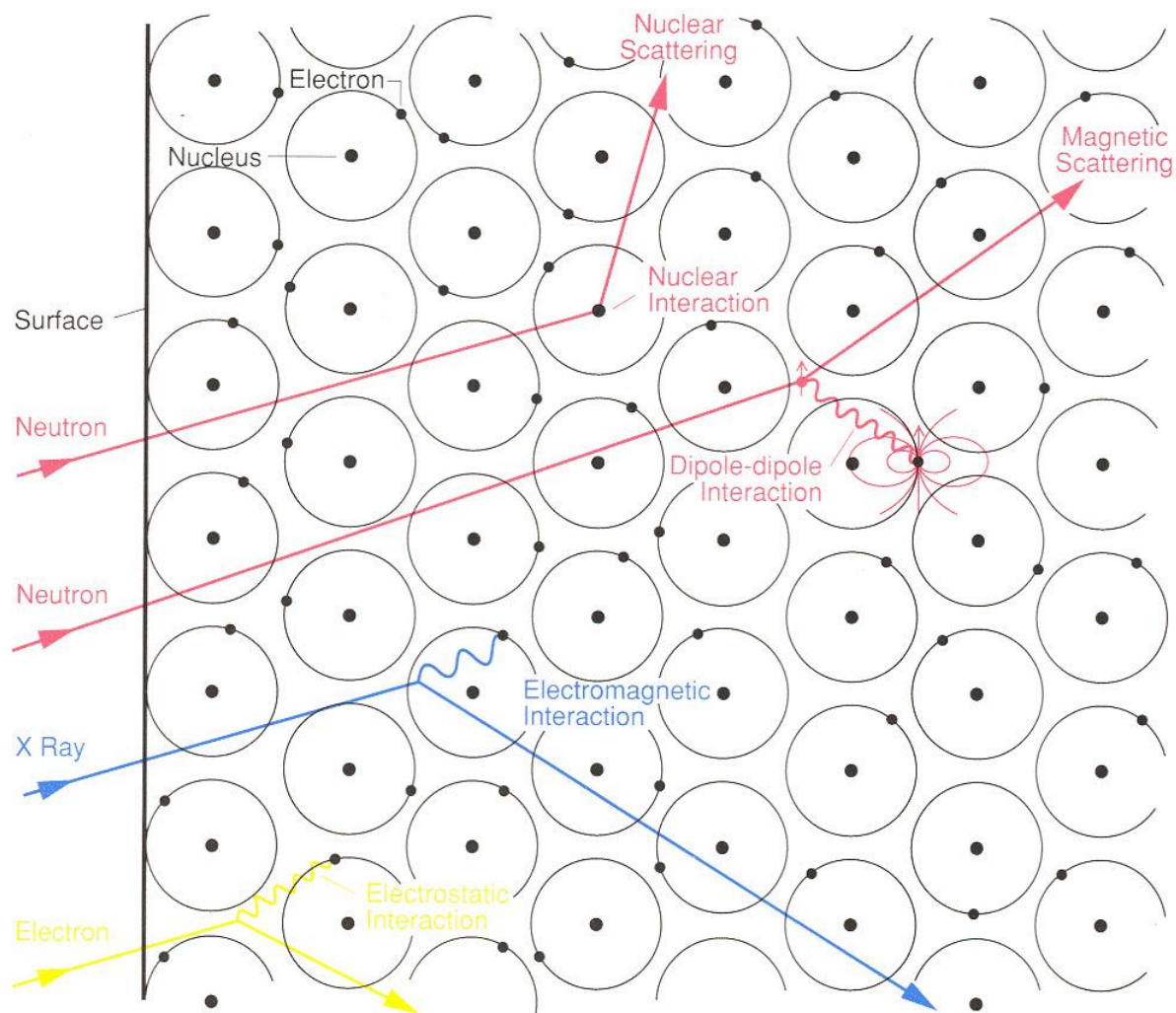


Jim Rhyne
NIST Center for
Neutron Research



Thanks to Peter Gehring, Jeff Lynn, and Dan Neumann for preparing many of the slides

Interaction of radiation with materials

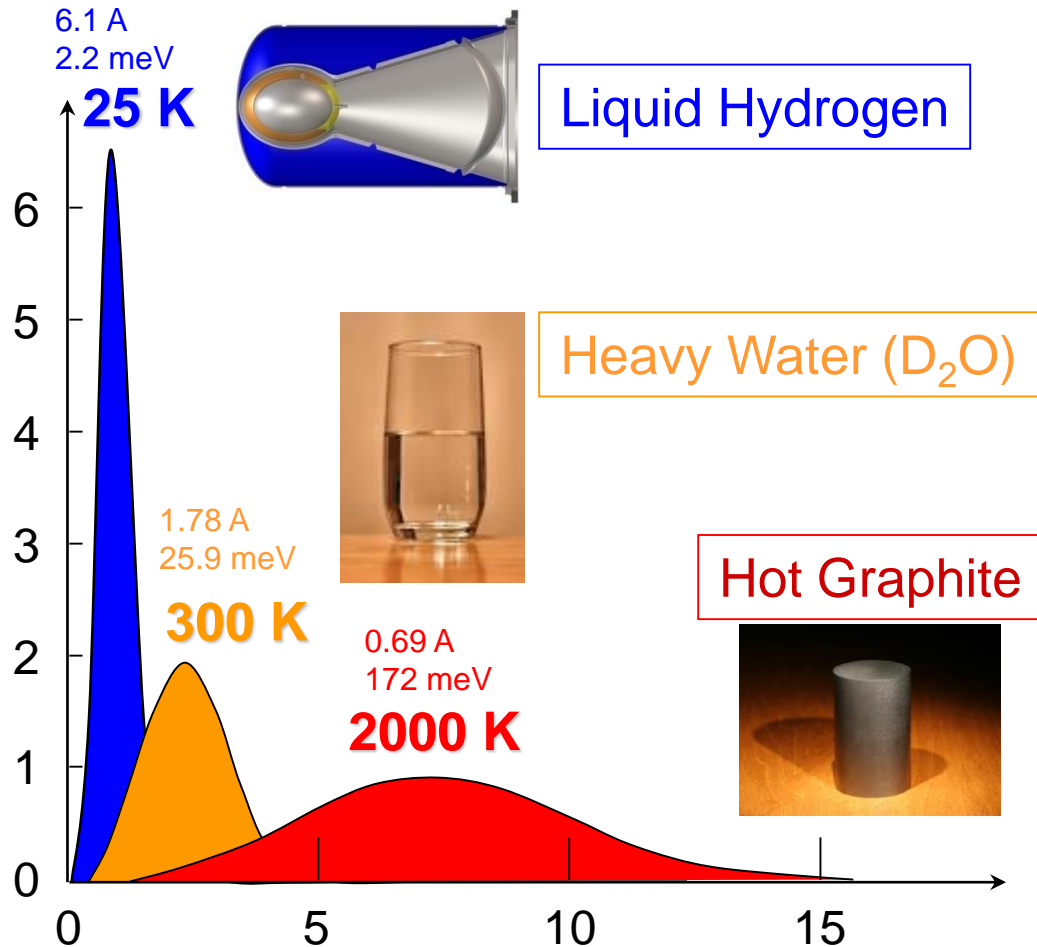
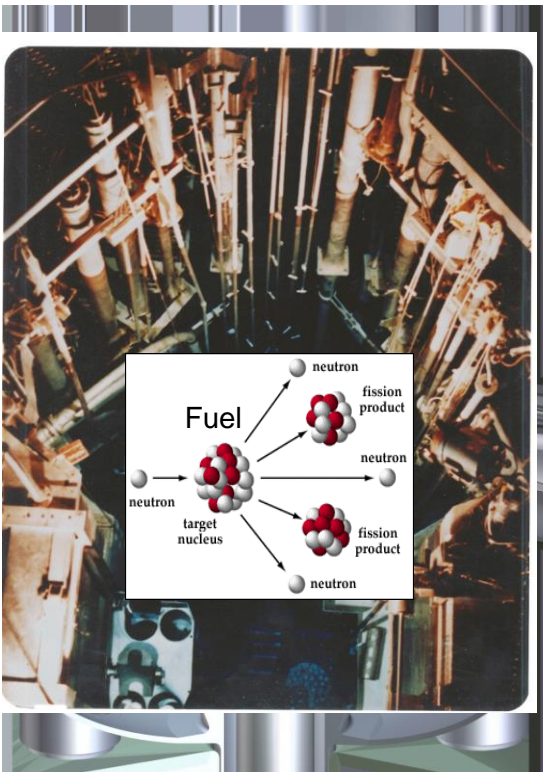


From Roger Pynn
– *Neutron Primer*

Neutron Source: Moderation

Maxwellian
Distribution

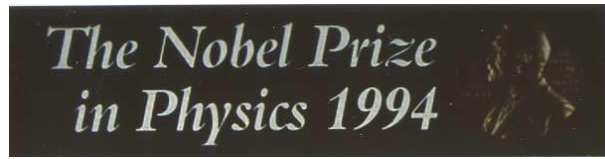
$$\Phi \sim v^3 e^{(-mv^2/2k_B T)}$$



“Fast” neutrons: $v = 20,000$ km/sec

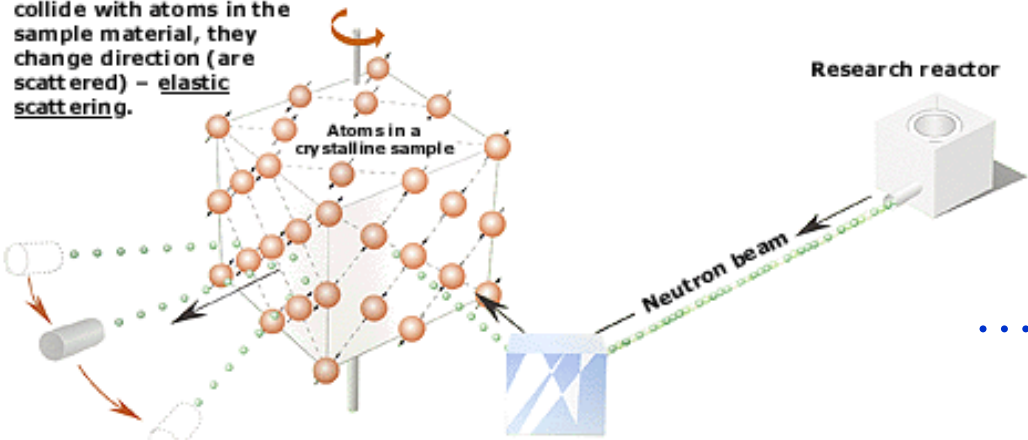
Neutron velocity v (km/sec)

Neutron Scattering

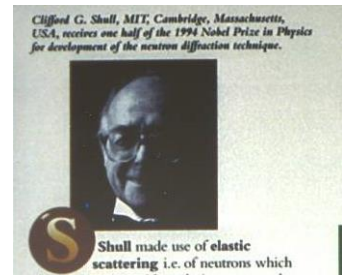


Neutrons show where the atoms are....

When the neutrons collide with atoms in the sample material, they change direction (are scattered) - elastic scattering.

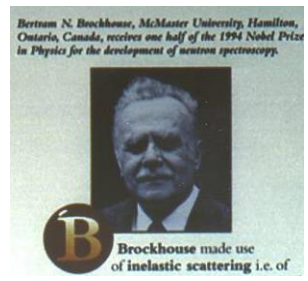
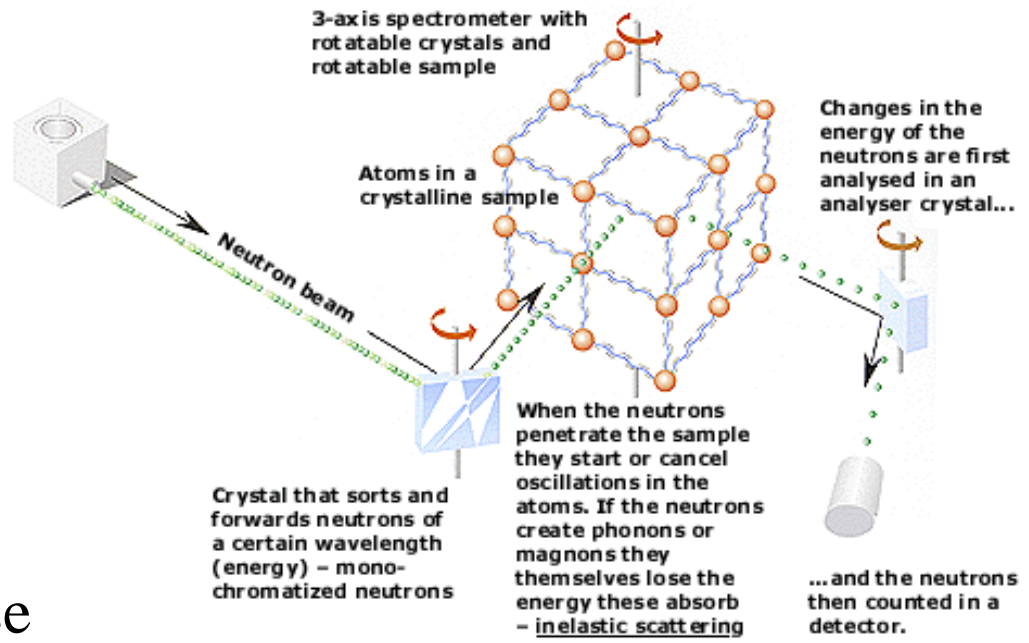


Detectors record the directions of the neutrons and a diffraction pattern is obtained. The pattern shows the positions of the atoms relative to one another.



Cliff Shull

...and what the atoms do.



Bertram Brockhouse

Scattering of neutrons by nuclei

- A single isolated nucleus will scatter neutrons with an intensity (isotropic)

$$- I = I_0 \sigma = I_0 [4\pi b^2]$$

where I_0 = incident neutron intensity,
 b = scattering amplitude for nucleus

- What happens when we put nucleus (atom) in lattice?

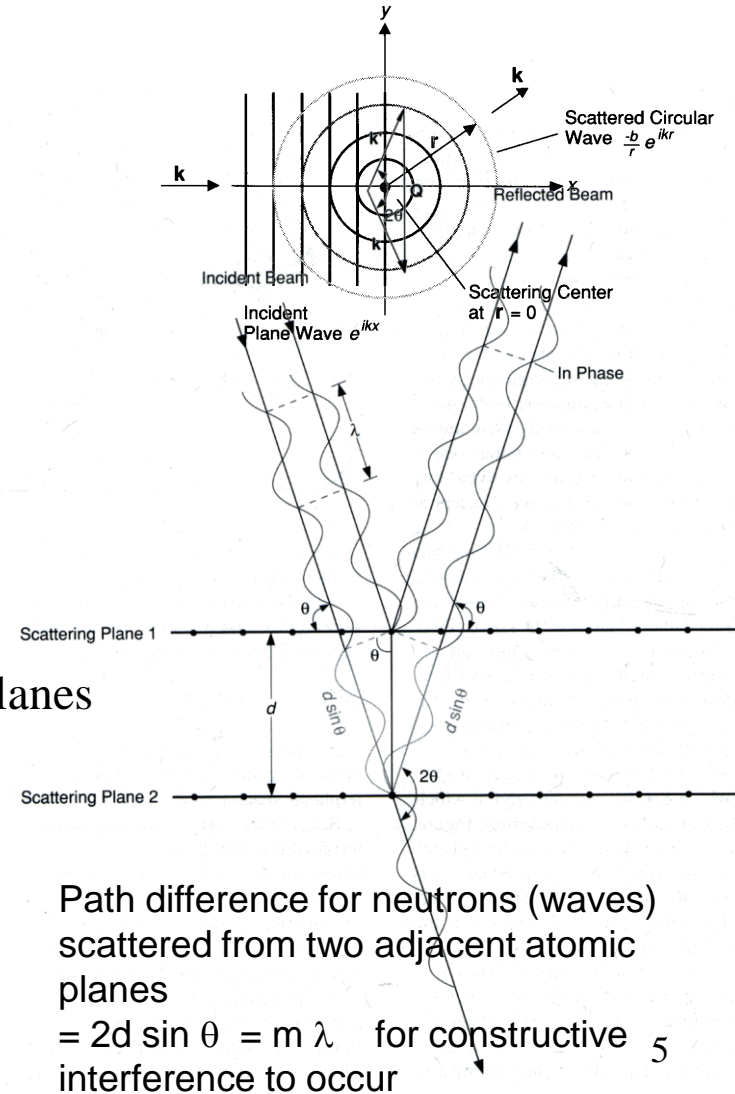
- Scattering from N nuclei can add up because they are on a lattice (constructive interference)
- Adding is controlled by phase relationship between waves scattered from different lattice planes
- Intensity is no longer isotropic

Bragg law gives directional dependence

$$\lambda = 2d \sin \theta$$

- Wave vector $|\mathbf{k}| = 2\pi/\lambda$

-- Intensity $I(Q, \text{ or } \theta)$ is given by a scattering cross-section



How do we find the wavelength to make the Bragg law work?

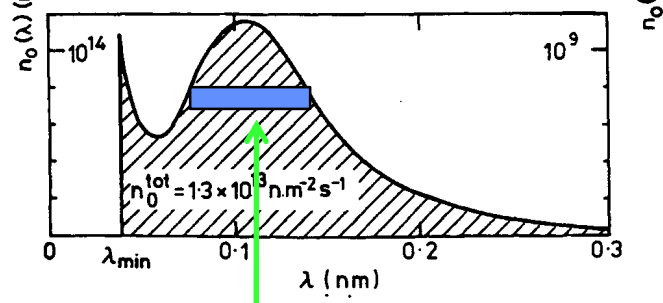
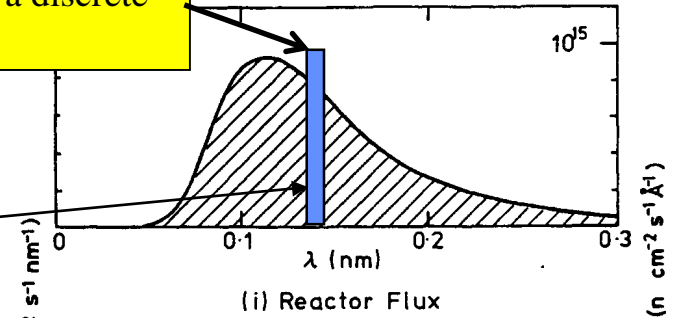
- **Reactor**

- Fission of U^{235} produces neutrons
- Fission spectrum moderated (slowed down) by either D_2O or H_2O (less effective moderator) and neutrons are extracted through beam tubes for spectrometers – fixed wavelength used

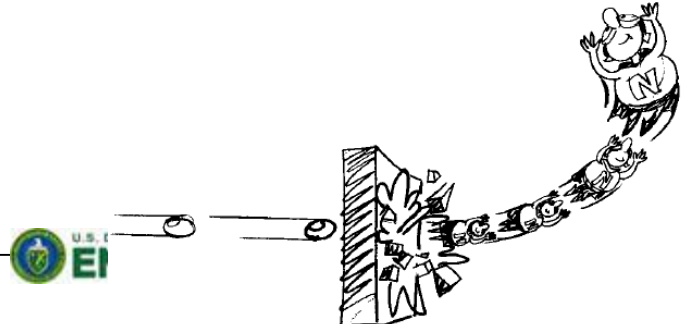
- **Spallation source**

- High E protons (e.g., 800 MeV) impinge on target (W, Hg or U)
- Nucleus of target is “exploded” by proton impact and emits 15 – 25 neutrons per proton with average E = 55 MeV (+ γ s, nucleons and neutrinos)
- Neutrons moderated by liquid H, H_2O or methane
- Spallation sources generally operate in pulse mode – 60 Hz at SNS

Monochromator crystal is used to saw-out a discrete wavelength



Time of flight is used to sort out wavelengths

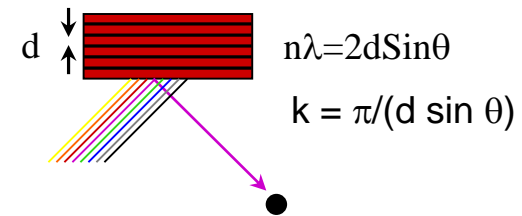


Methods of Specifying and Measuring \vec{k}_i and \vec{k}_f



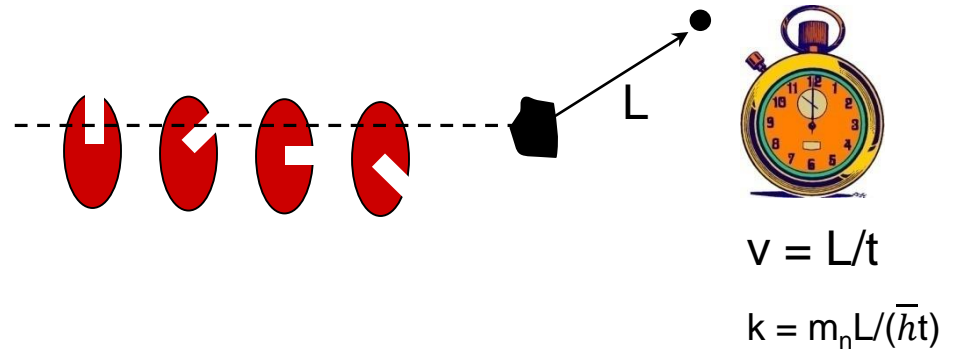
1. Bragg Diffraction

BT7, MACS, HFBS



2. Time-of-Flight (TOF)

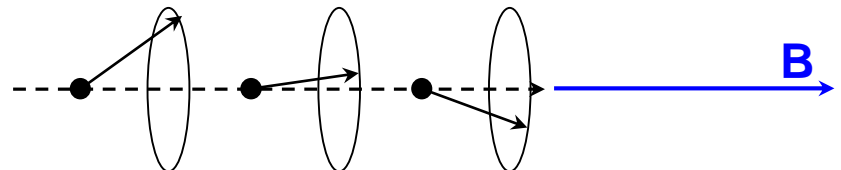
DCS, HFBS



3. Larmor Precession

NSE

Larmor precession angle of neutron mag moment acts as a clock – if $\Delta E \neq 0$ precession angles before and after sample are different.



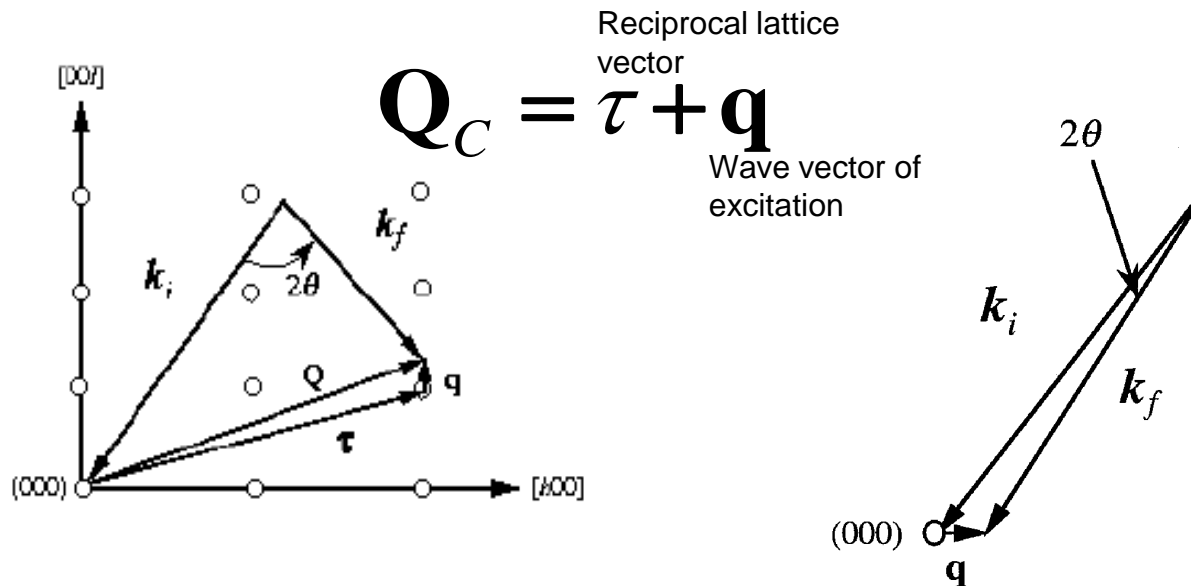
It's all about Conservation of Momentum

$$(\mathbf{p} = \hbar\mathbf{k} \text{ and Energy } (E = \hbar\omega = \left| \frac{p^2}{2m} \right|$$



$$\mathbf{Q} = \mathbf{k}_i - \mathbf{k}_f \text{ Wave vector transfer to excitation}$$

$$\Delta E = \frac{\hbar^2 k_i^2}{2m} - \frac{\hbar^2 k_f^2}{2m} \text{ Energy transfer to/from excitation}$$



Energy, wave vector, and wavelength relations for various probes



$$E_{neutron} (meV) = 2.0719k^2 = 81.7968 / \lambda^2$$

$$E_{photon} (keV) = 2.0k = 12.4 / \lambda$$

$$E_{electron} (eV) = 3.8k^2 = 150 / \lambda^2$$

$$1 meV = 11.6 K \quad (k_B T)$$

$$1 meV = 8.06 cm^{-1} \quad (E / hc)$$

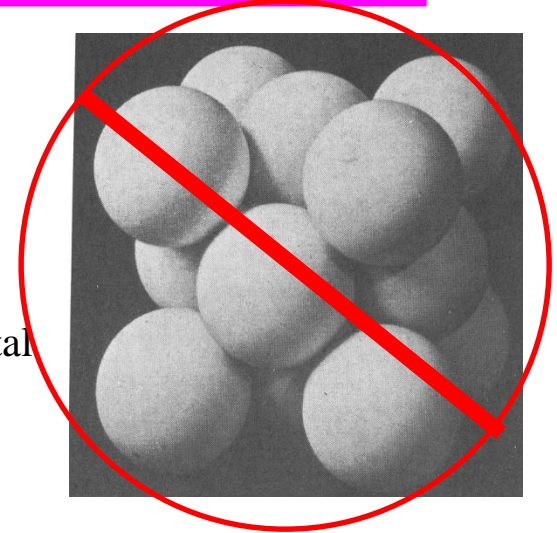
$$1 meV = 0.2418 THz \quad (E / h)$$

$$1 meV / \mu_B = 17.3 T \quad (E / \mu_B)$$

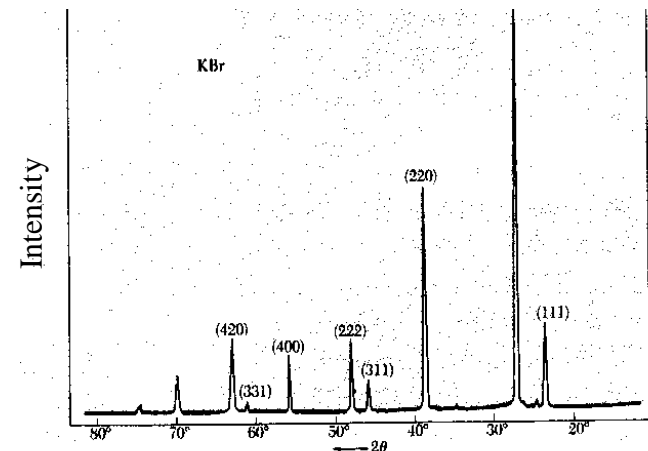
Golden Rule of Neutron Scattering

- We don't take pictures of atoms!

Atoms in fcc crystal



- Job security for neutron scatterers – we live in *reciprocal space*



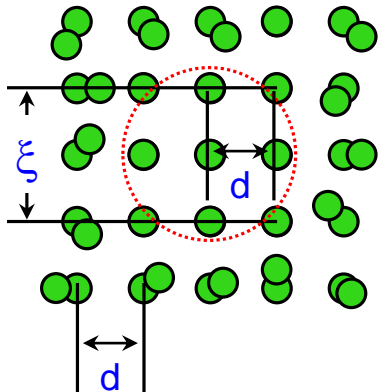
Review: Main Messages of the Week



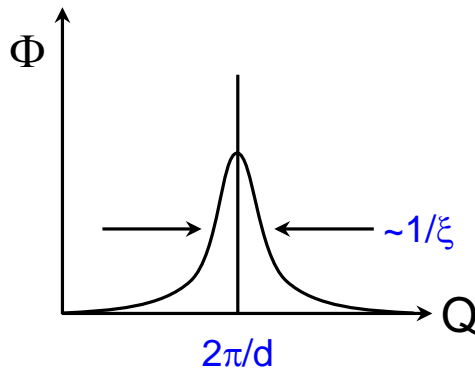
(3) The scattered neutron flux $\Phi(\vec{Q}, \hbar\omega)$ is proportional to the space (\vec{r}) and time (t) Fourier transform of the probability $G(\vec{r}, t)$ of finding one or two atoms separated by a particular distance at a particular time.

$$\Phi \propto \frac{\partial^2 \sigma}{\partial \Omega \partial \omega} \propto \iint e^{i(\vec{Q} \cdot \vec{r} - \omega t)} G(\vec{r}, t) d^3 \vec{r} dt$$

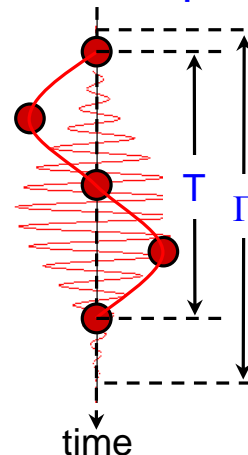
Real space



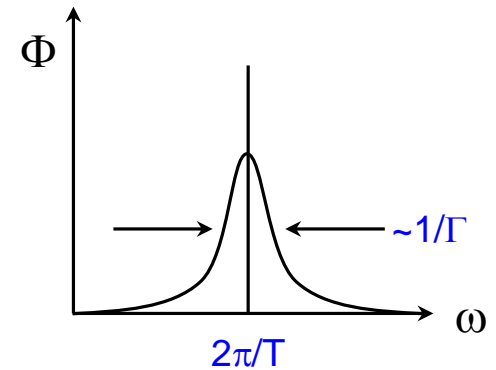
Q-space



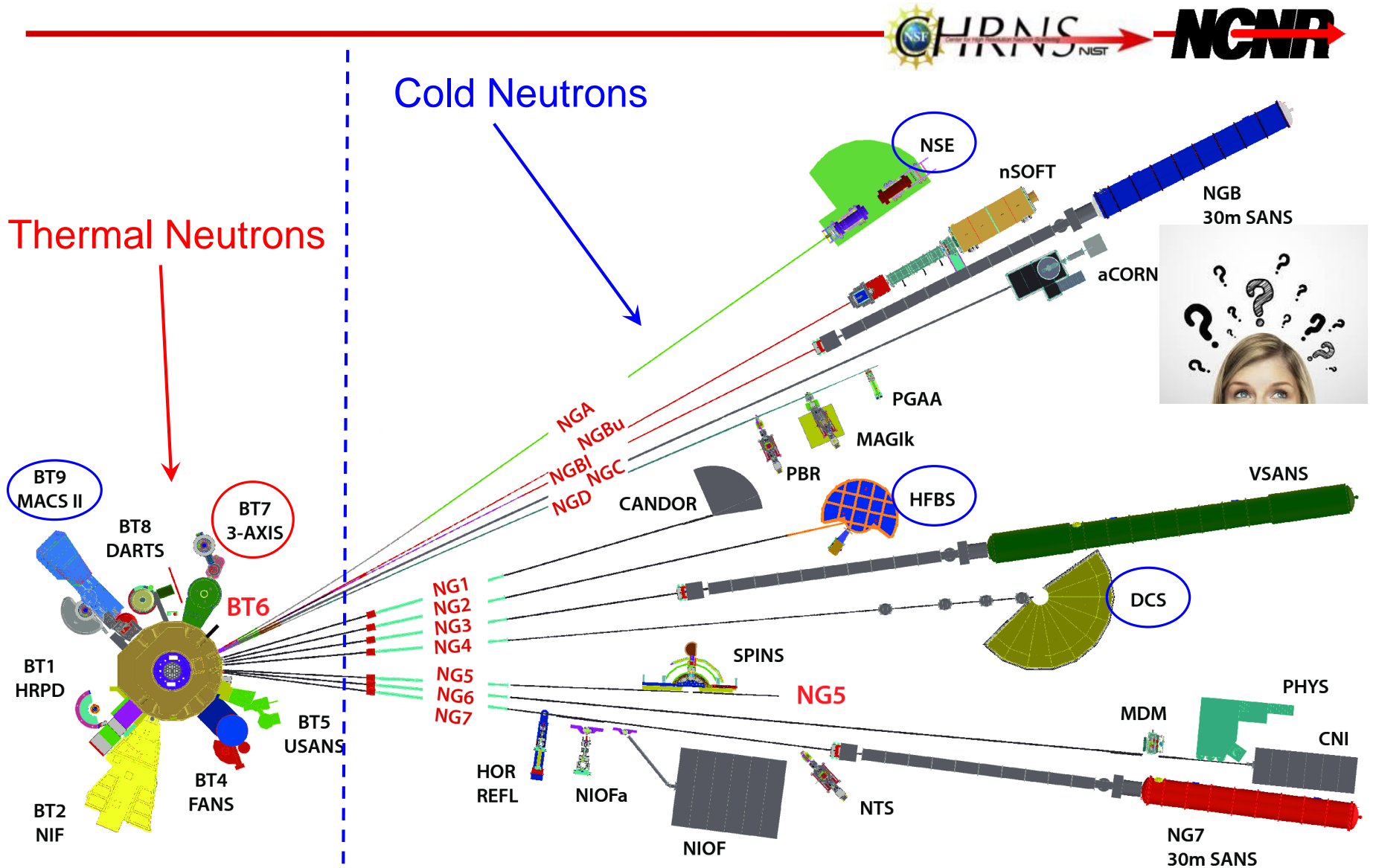
Time space



ω -space



The NCNR Menagerie of Instruments



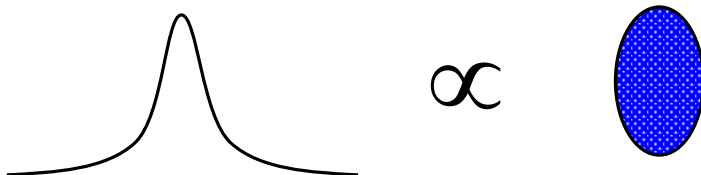
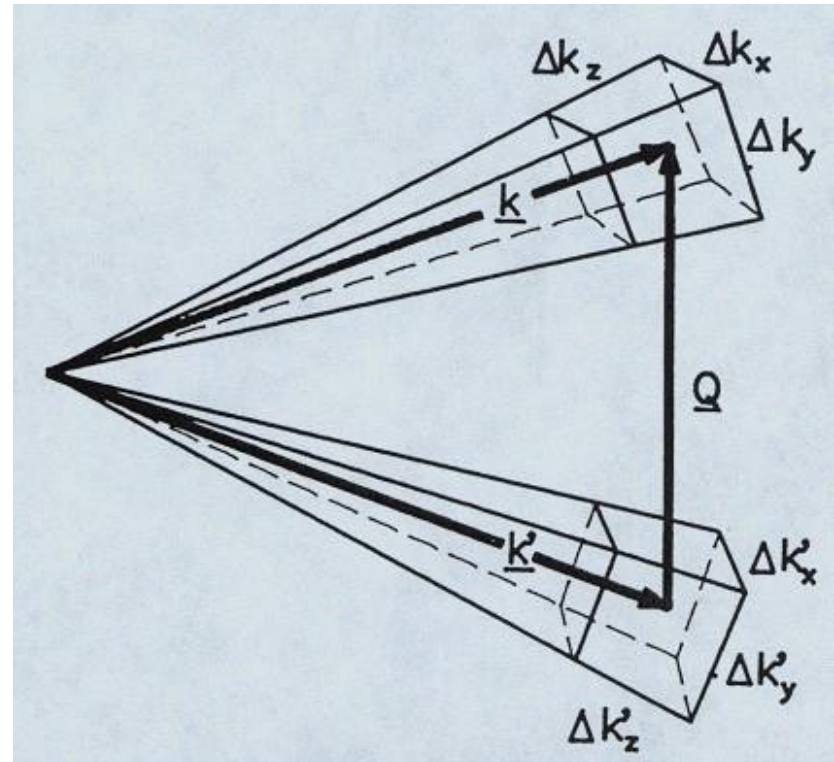
Why So Many Different Spectrometers?



Because neutron scattering is an intensity-limited technique. Thus detector coverage and resolution **MUST** be tailored to the science.

Uncertainties in the neutron wavelength and direction imply \mathbf{Q} and $\hbar\omega$ are only defined with a finite selectable precision.

The total signal in a scattering experiment is proportional to the resolution volume \rightarrow better resolution leads to lower count rates! *Choose carefully* ...



Courtesy of R. Pynn

How do I Choose the Right Spectrometer?



Two basic considerations:

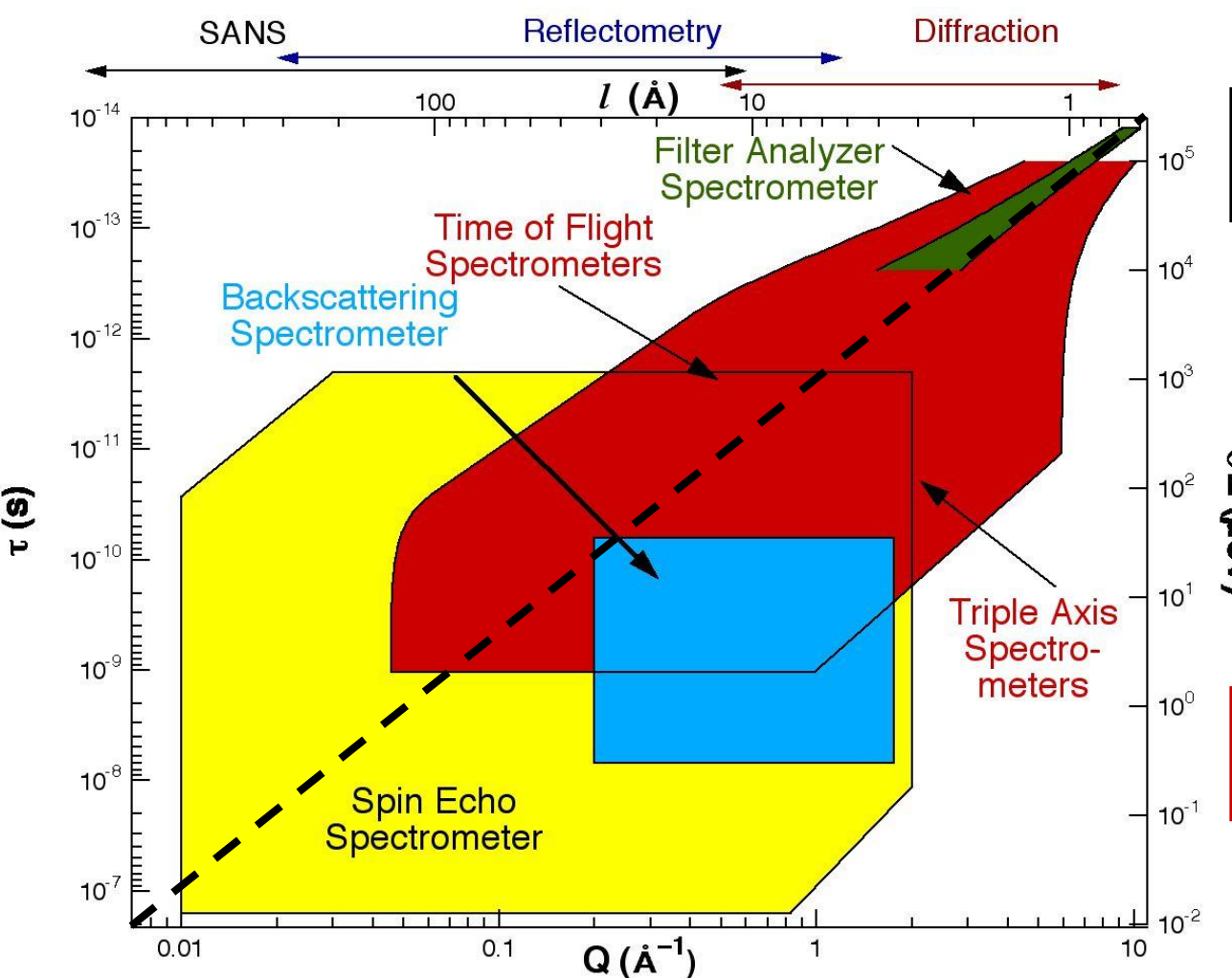
1. What are the **time** scales ($\hbar\omega$) of interest?
2. What are the **length** scales (Q) of interest?

(Some spectrometers overlap →
the choice may boil down to one of **resolution**)

Two additional considerations:

1. What **energy** resolution ($\Delta\hbar\omega$) is required?
2. What **momentum** resolution (ΔQ) is required?

Different Spectrometers Cover Different Regions of Phase Space



Do you see a pattern here?

Larger "objects" tend to exhibit slower motions.

Inelastic Spectrometers



Approx. Resol.

Thermal triple-axis instruments (BT-7) (BT-4) 1 meV

Cold neutron triple-axis instrument (MACS) (SPINS)

$S(\mathbf{Q}, E)$ Disk chopper time-of-flight spectrometer (DCS) (FANS) $\sim 250 \mu\text{eV}$

High flux backscattering spectrometer (HFBS) 1 μeV

$S(\mathbf{Q}, t)$ Spin-echo spectrometer (NSE) $\delta t \rightarrow \sim 10 \text{ neV}$

All these different spectrometers are designed differently to optimize intensity and resolution for different measurement requirements

Rules of Thumb



1. What are the energies ($\hbar\omega$), i.e. time scales ($\Delta t \sim 1/\omega$), of interest?

$\hbar\omega \approx 1-100 \text{ meV}$ - use a thermal triple-axis spectrometer like BT7.

$\hbar\omega \approx 20-30 \text{ } \mu\text{eV}$ - use HFBS or NSE.

In between - use MACS or DCS or a cold-neutron triple-axis spectrometer like SPINS.

2. Make sure that the length scales \mathbf{L} of the relevant motions lie within the range of the spectrometer. For example, consider the HFBS

$$Q_{\min} = 0.25 \text{ } \text{\AA}^{-1} \rightarrow L_{\max} \sim 25 \text{ } \text{\AA}$$

$$Q_{\max} \approx 1.75 \text{ } \text{\AA}^{-1} \rightarrow L_{\min} \sim 3.5 \text{ } \text{\AA}$$

$$Q = 2\pi/L$$

REMEMBER - Q_{\min} and Q_{\max} are inversely proportional to the incident neutron wavelength

More Rules of Thumb



Is your sample polycrystalline or amorphous?

Does **ONLY** the magnitude (not the direction) of Q matter?

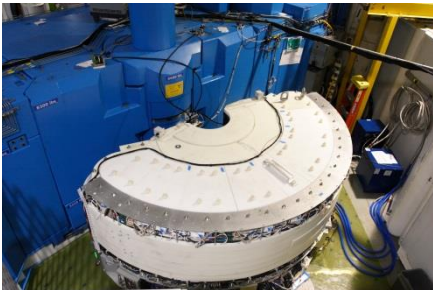
Is the expected Q -dependence of the scattering weak?

This often means that you want to look at a large region of $Q, \hbar\omega$ space, or that you can sum the data over a large region of $Q, \hbar\omega$ space.

YES? Consider instruments with large analyzer areas.

NO? Consider using BT7, SPINS, or NSE.

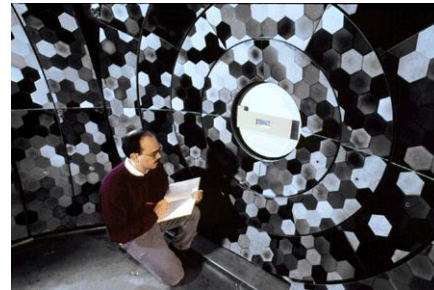
MACS



DCS



HFBS



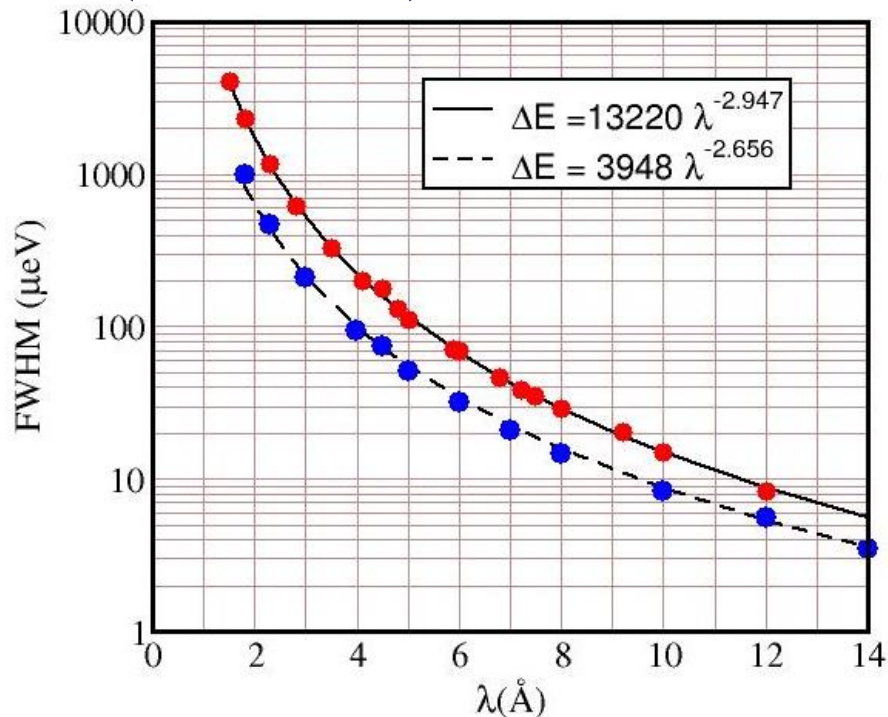
BT7



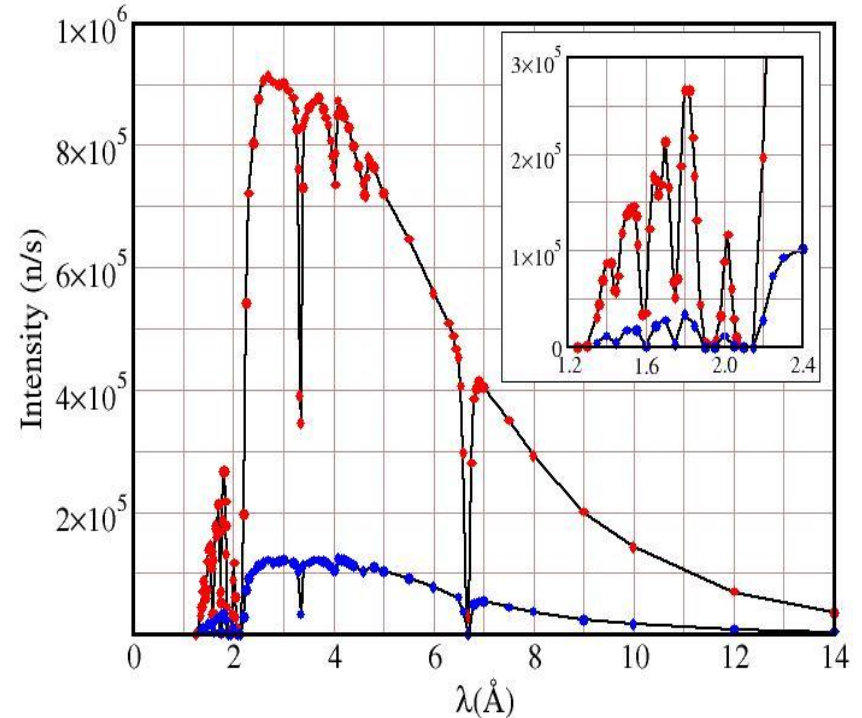
Things to Consider When Choosing DCS



ΔE (resolution)



Incident Intens.



Quantities varied

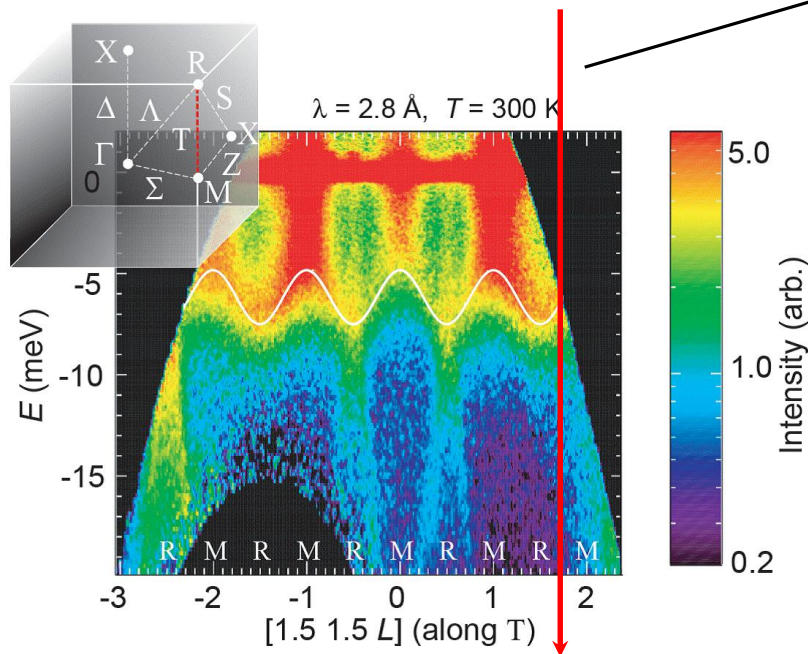
- wavelength λ
- chopper slot widths W

Remember – Intensity ↓
Resolution ↑

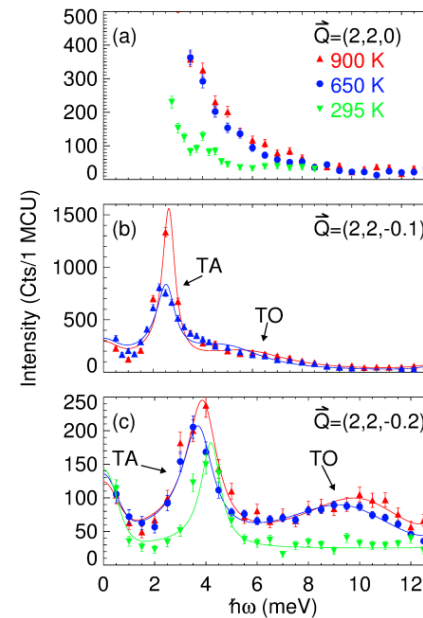
Example: DCS versus BT7



DCS Broad surveys in $Q-\omega$



BT7 Limited regions in $Q-\omega$



Rules of Thumb: (think carefully before violating)

DCS, MACS – systems requiring resolution $< 400 \mu\text{eV}$

BT7 – single crystals – resolution $> 100 \mu\text{eV}$

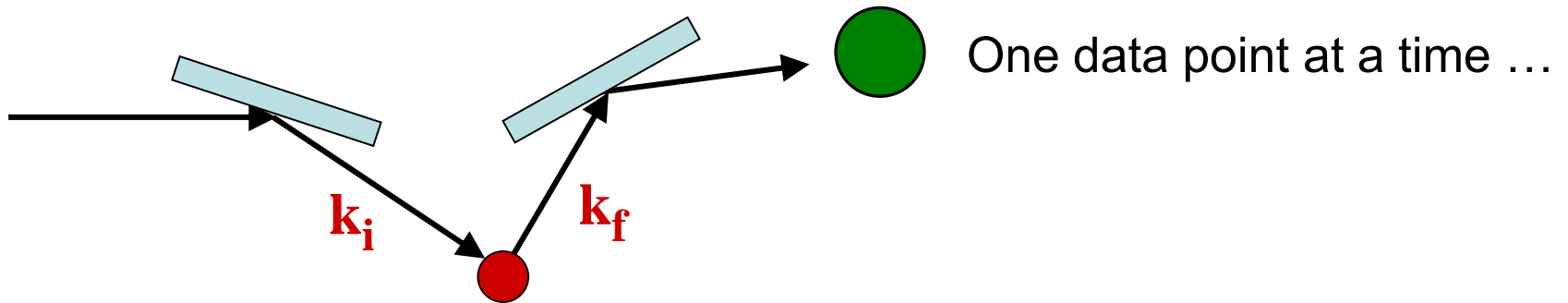
depends on collimation and monochromator/analyzer

Things to Consider When Choosing BT7



Triple axis spectrometers are typically used when either -

- (1) the *direction* of \mathbf{Q} is important or
- (2) the interesting region of \mathbf{Q} - ω space is of *limited extent*.



Remember – **Intensity** ↓
Resolution ↑

Things to Consider When Choosing HFBS

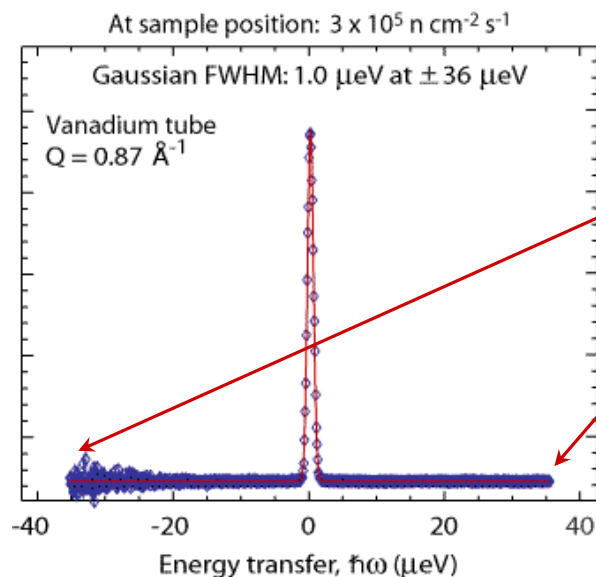


$$0.25 \text{ \AA}^{-1} < Q < 1.75 \text{ \AA}^{-1}$$

Do the length scales of interest lie within this Q-range?

$$\delta Q < 0.1 - 0.2 \text{ \AA}^{-1}$$

Can you live with such coarse Q-resolution?



Do the features of interest lie within this $\hbar\omega$ -range?

Do you really require such good energy resolution $\delta E \sim 1 \text{ \mu eV}$?

General Sample “Design”



Know as much about your sample as possible!!
(Beamtime costs ~ \$5000/day!!)

Other considerations:

What's the structure (in a general sense)?

Are there any phase transitions (or a glass transition)?

What isotopes are present?

Supplementary data from other measurements ...

Magnetization vs T

Muon spin relaxation

X-ray data

Specific heat vs T

Raman spectroscopy

General Sample “Design”



Try to avoid isotopes that are strongly absorbing.

${}^6\text{Li}$ ${}^{10}\text{B}$ ${}^{113}\text{Cd}$ ${}^{157}\text{Gd}$

For a complete listing go to

<http://www.ncnr.nist.gov/resources/n-lengths>

Sample “Design”



Single crystals yield the most information.

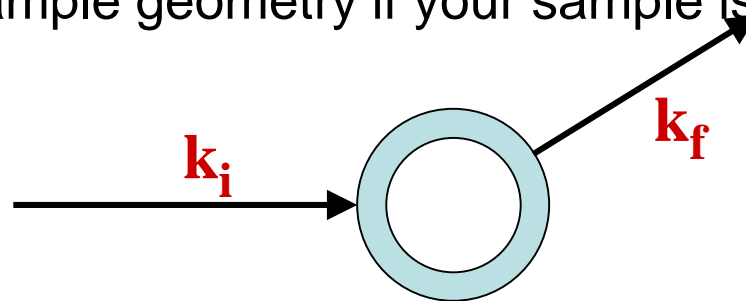
Increase the intensity by increasing the amount of sample.

If you have a powder, use a cylindrical container (rather than flat plate).

Annular may be the best sample geometry if your sample is absorbing.

Transmission of the beam
should be ~70-90%.

$$I/I_0 = \exp(-n\sigma_A T)$$



Almost all experiments of collective excitations involve coherent scattering
→ If sample contains H it should be deuterated (D).

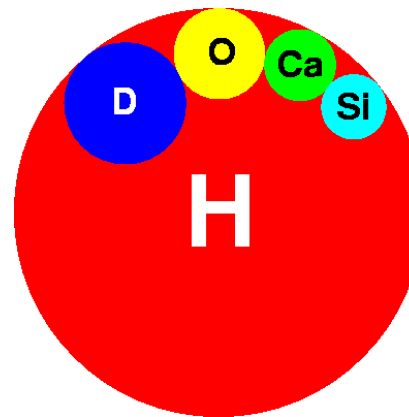
Sample Selective Deuteration



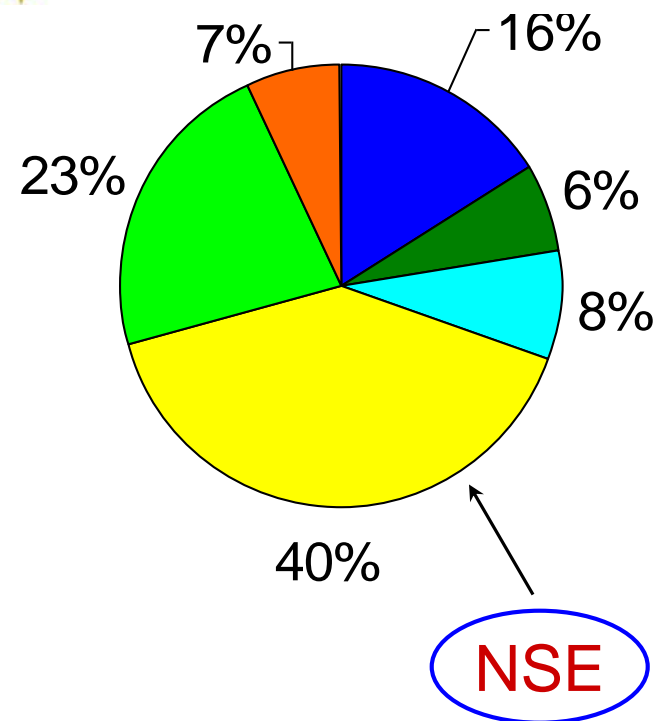
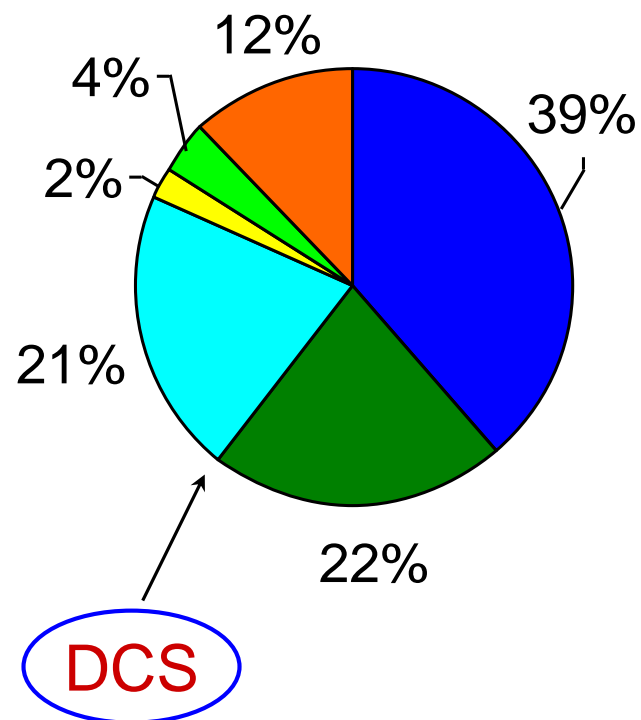
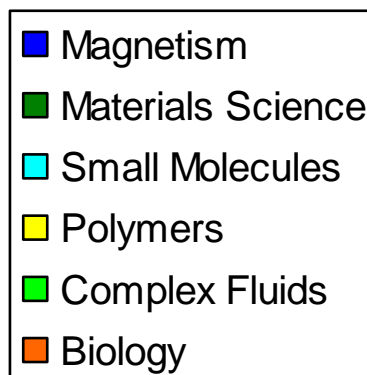
Does the sample contain H?

Remember: **Neutrons LOVE H!!**

Create a sample where -
the “interesting” portions are hydrogenated and
the “uninteresting” portions are deuterated.



Typical Distributions of Science by Instrument



Some Summer School Success Stories



2001



Jae-Ho Chung
University Prof.

2003



Vicky Garcia-Sakai
ISIS Staff Scientist

1999



William Ratcliff
NCNR Staff Physicist

1997



Rob Dimeo
NCNR Director

Acknowledgements



Organizers – Joe Dura and Yamali Hernandez

Administrative staff

Experiment teams



Enjoy the Science With Neutrons!